

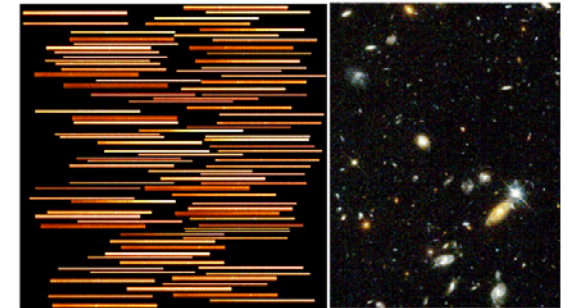
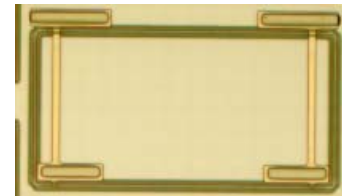
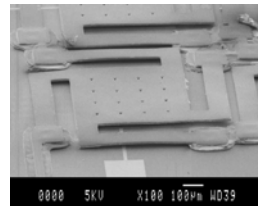
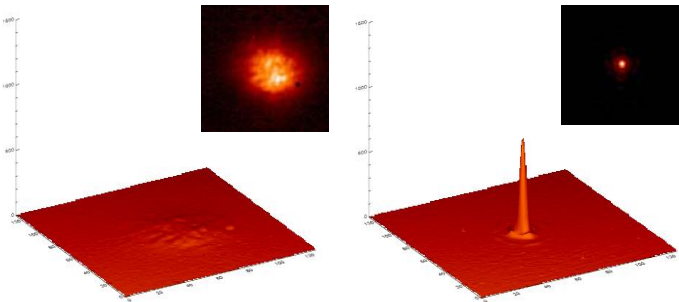
MOEMS devices for future astronomical instrumentation

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Véronique Conédéra², Norbert Fabre², Sylvaine Muratet², Henri Camon²,
Severin Waldis³, Wilfried Noell³

¹*Laboratoire d'Astrophysique de Marseille, France*

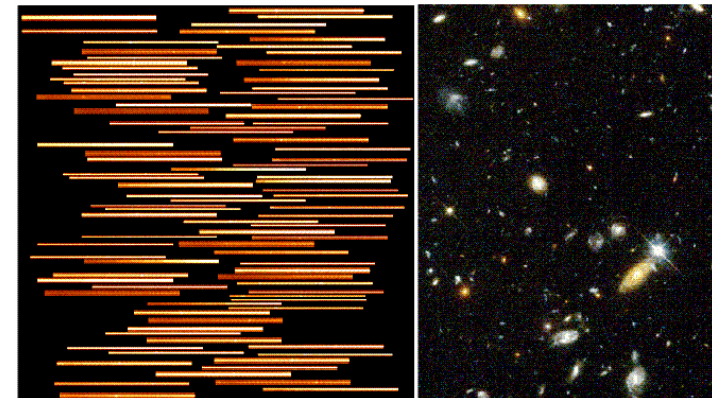
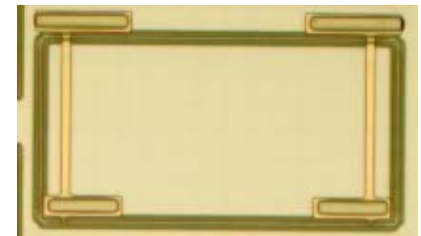
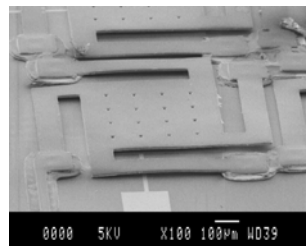
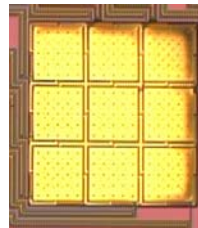
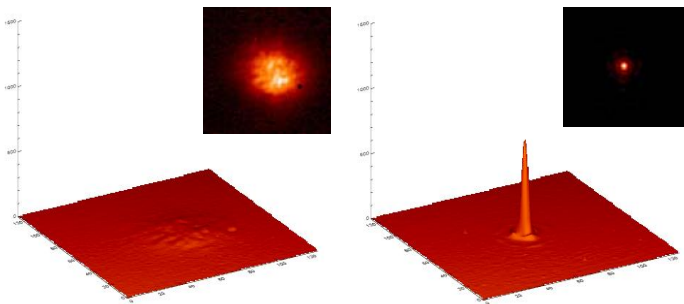
²*Laboratoire d'Analyse et d'Architecture des Systèmes, France*

³*Institut de Microtechnologies, U. de Neuchâtel, Switzerland*

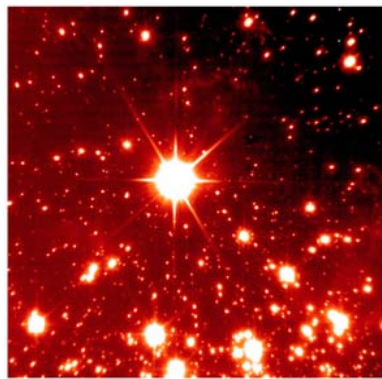
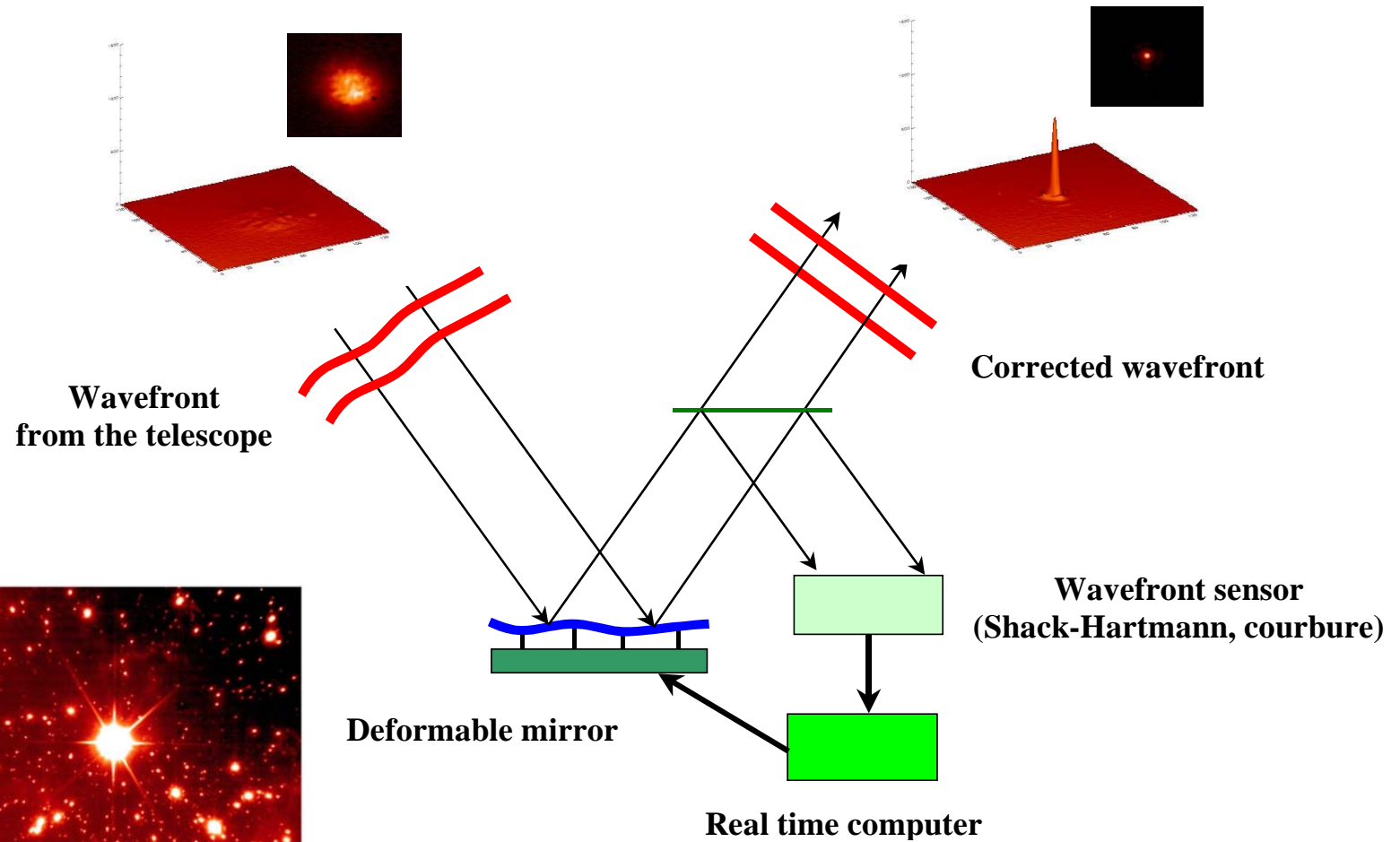


Astronomical cases

- ◆ **Origin's quest**
 - Galaxies formation and evolution
 - Stars and planetary systems formation
 - Life's Origin
- ◆ **Instrumental needs**
 - High resolution
 - Multiplex measurements



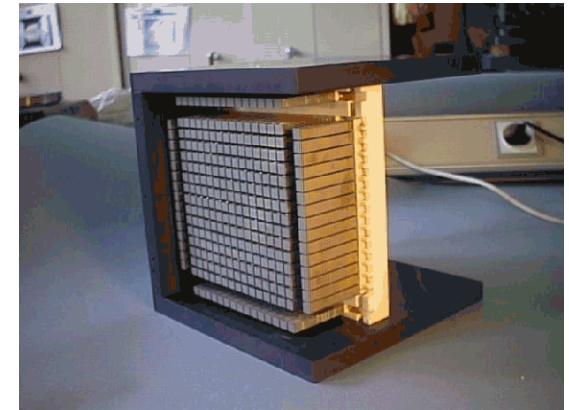
Wavefront correction



Deformable mirrors

◆ Mirrors with classical technology

- Piezo and magnetic actuators
- Few 100 actuators
- Inter-actuator spacing 5mm



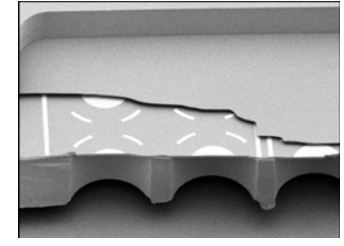
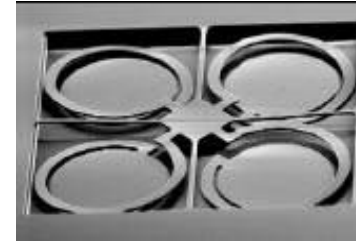
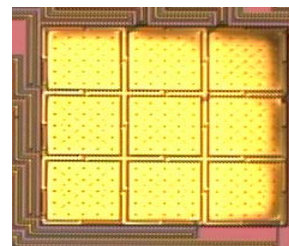
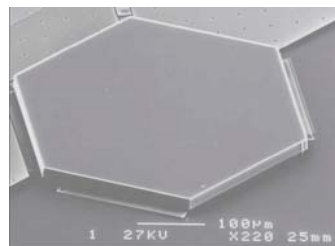
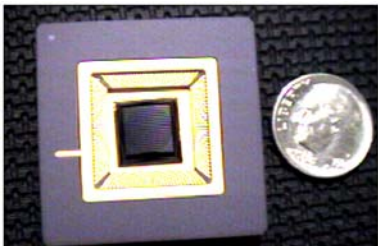
◆ Micro-deformable mirrors (MDM)

- Electrostatic force
- Up to 500 000 actuators (Extremely Large Telescope)
- Inter-actuator spacing 500 μ m-1mm

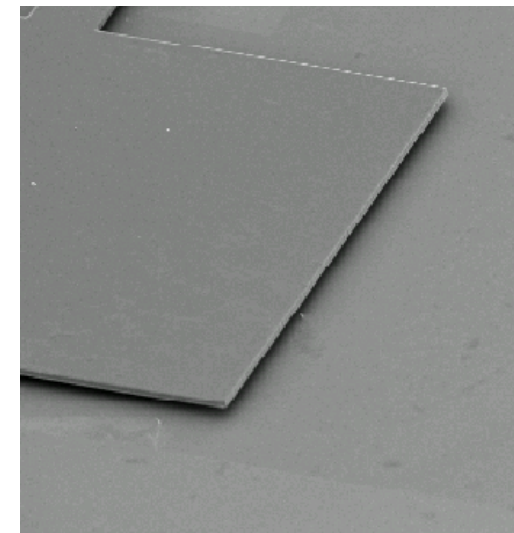
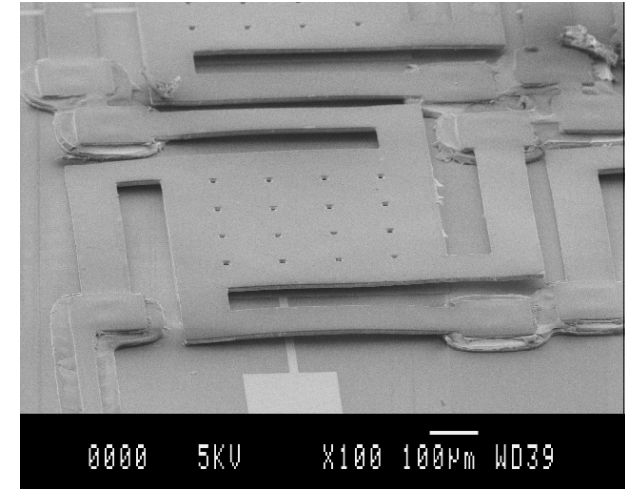
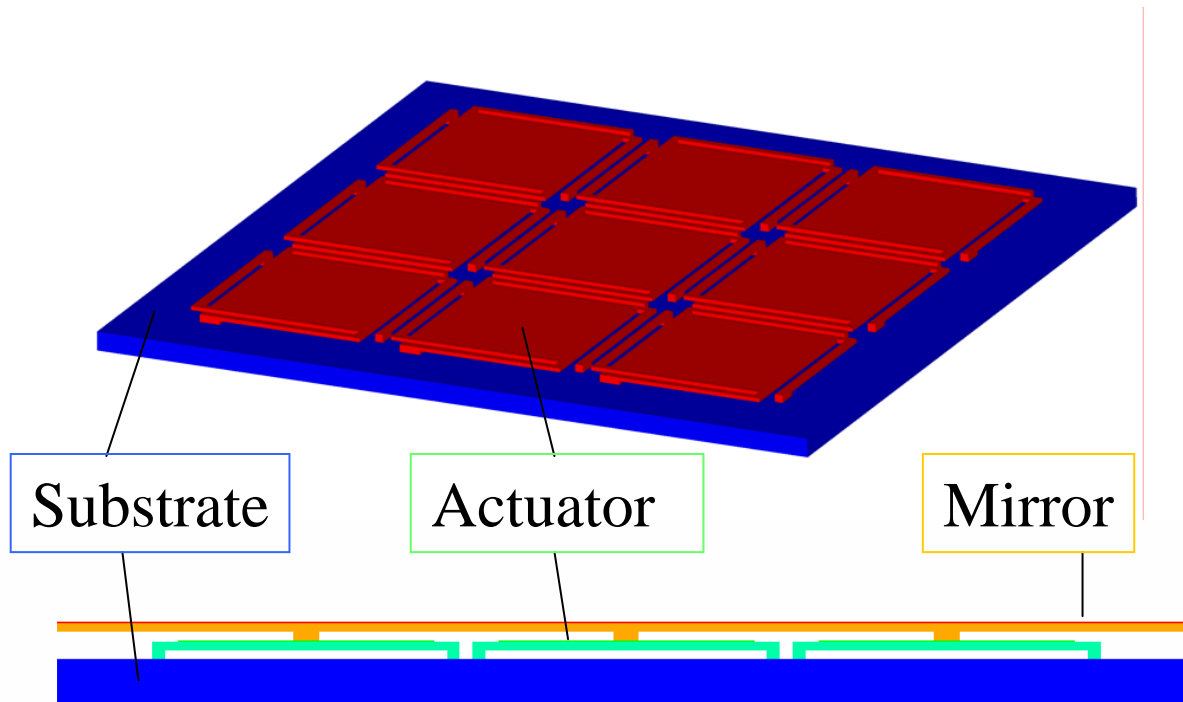


◆ MDM with membrane

◆ MDM with attached actuators



- ◆ Collaboration LAM-LAAS since 2001
- ◆ New materials : polymers
- ◆ Original process



MOEMS characterization bench

◆ Principle

- Twyman-Green interferometer
- Low coherence light source
- Different magnifications available (4X...0.25X)



CCD Camera

White light

Pin-hole

Interference filter

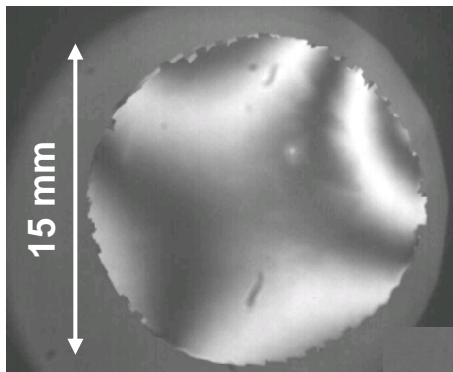
Sample

Reference flat

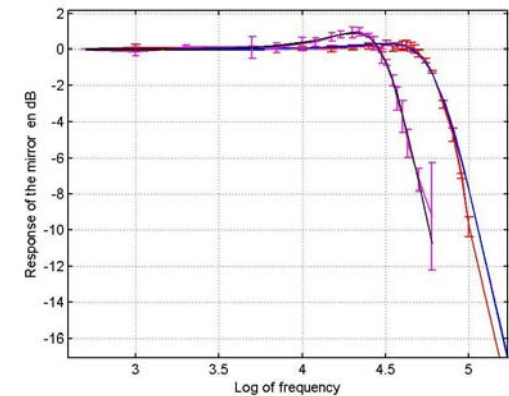
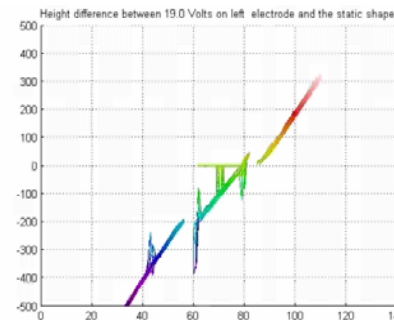
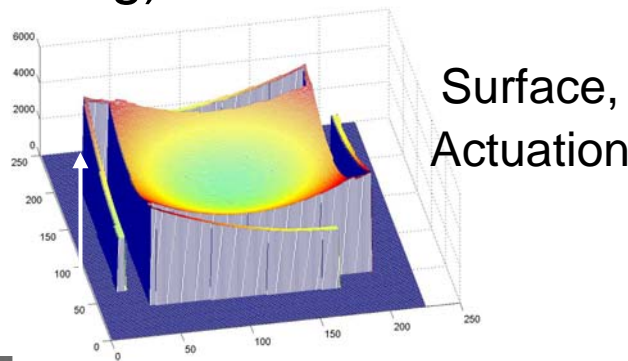
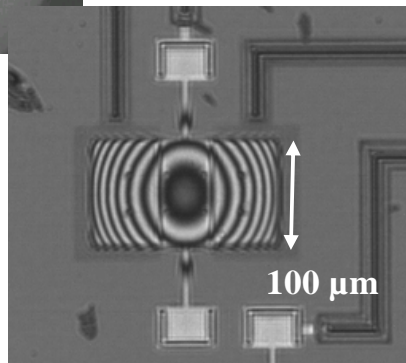
Surface measurement and actuation

◆ Measurement capabilities

- Surface measurement (phase-shift, two wavelengths)
- Actuation measurement (**resolution < 1nm**)
- Dynamical measurement (time-averaged interferometry)
- Probe tips (chip level testing)



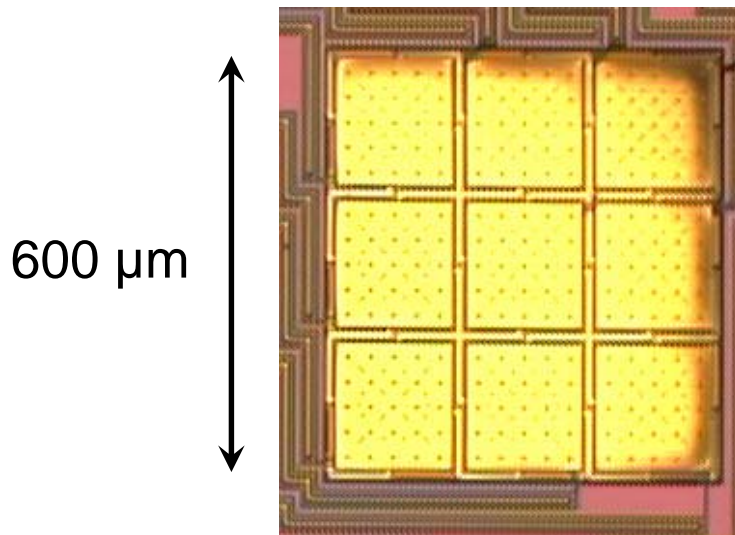
Imagery



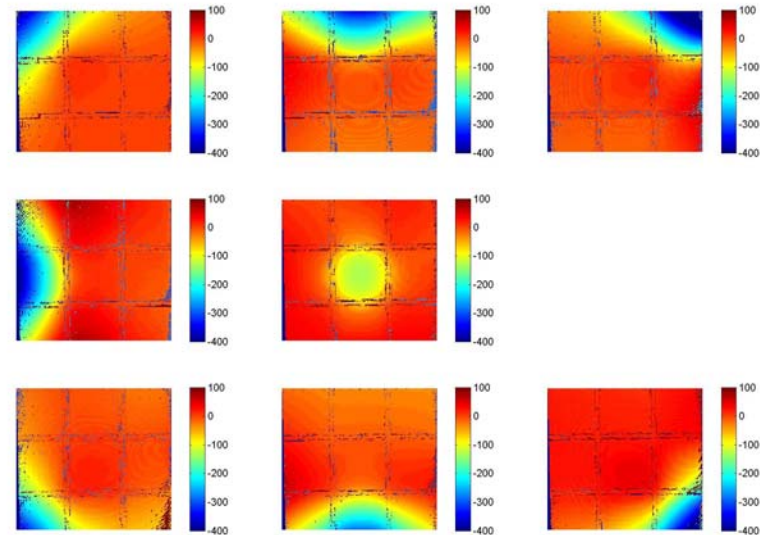
Dynamical response

- ◆ First prototype of an MDM designed by LAM-LAAS, realized in the Memscap foundry (US)

Material : Poly-Silicon



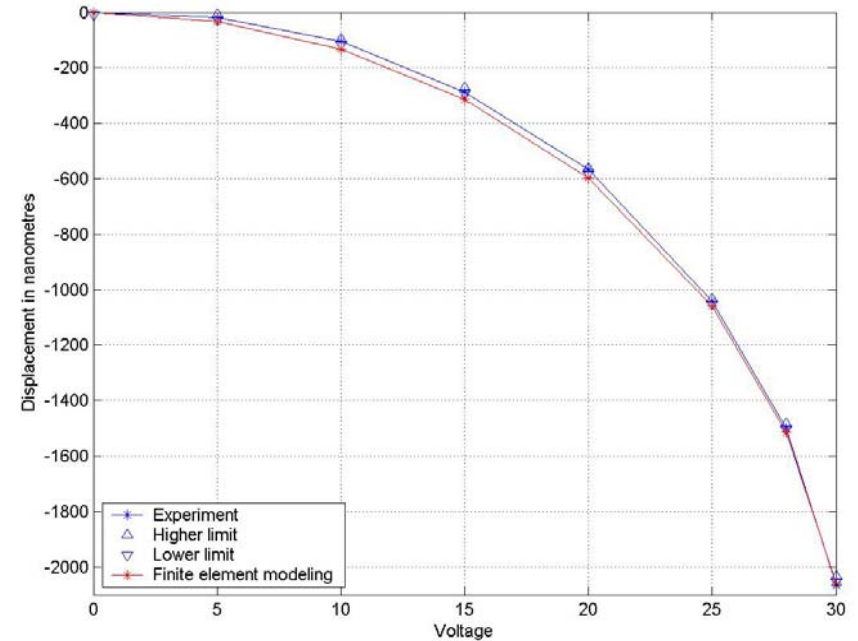
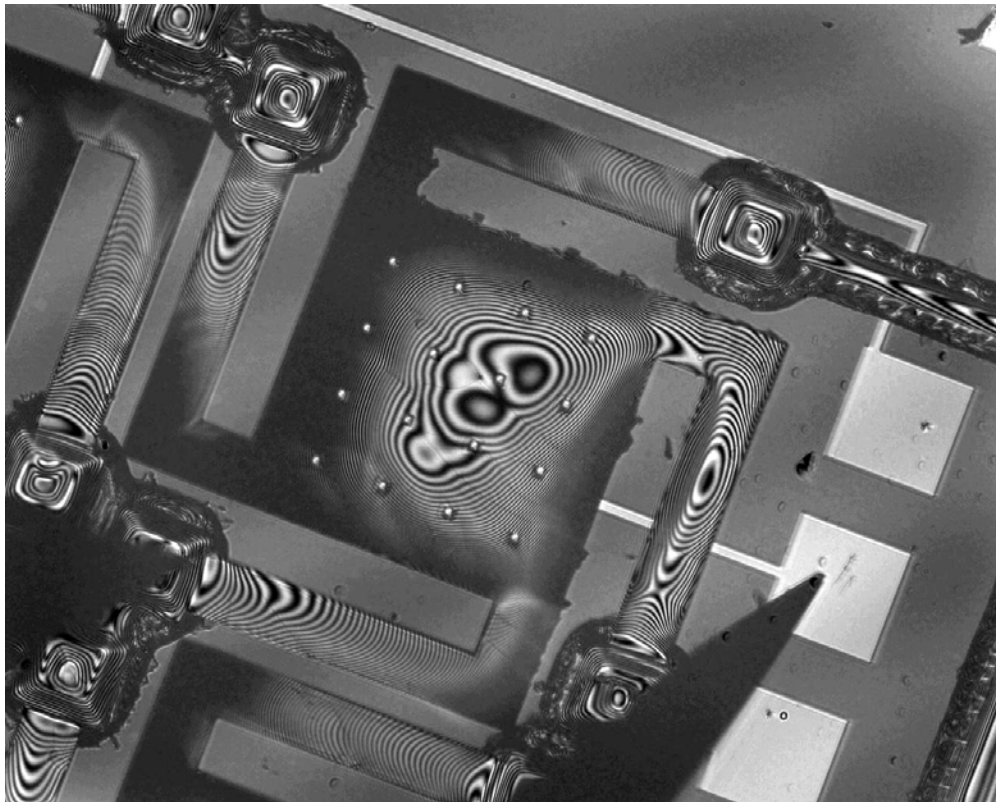
Mirror with 9 actuators



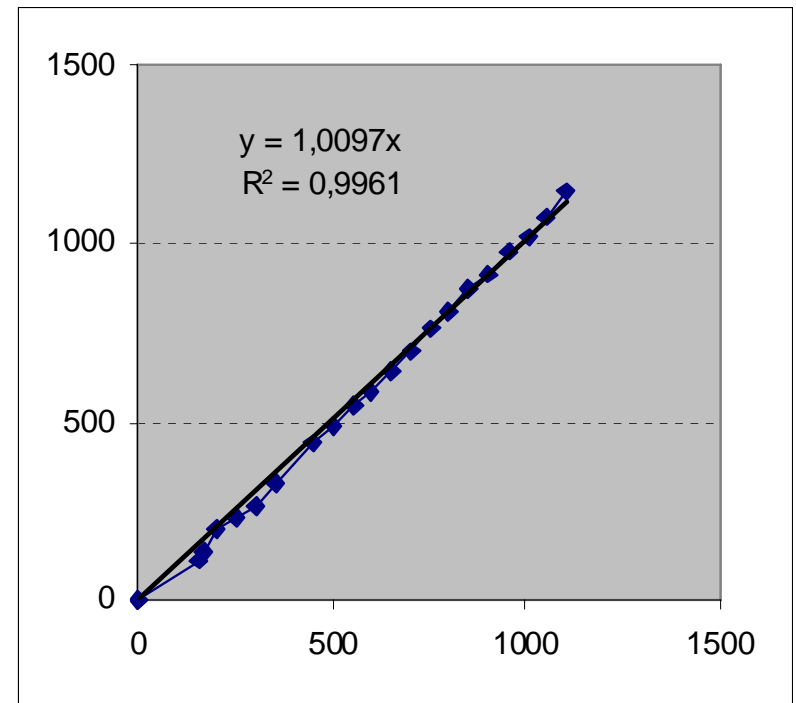
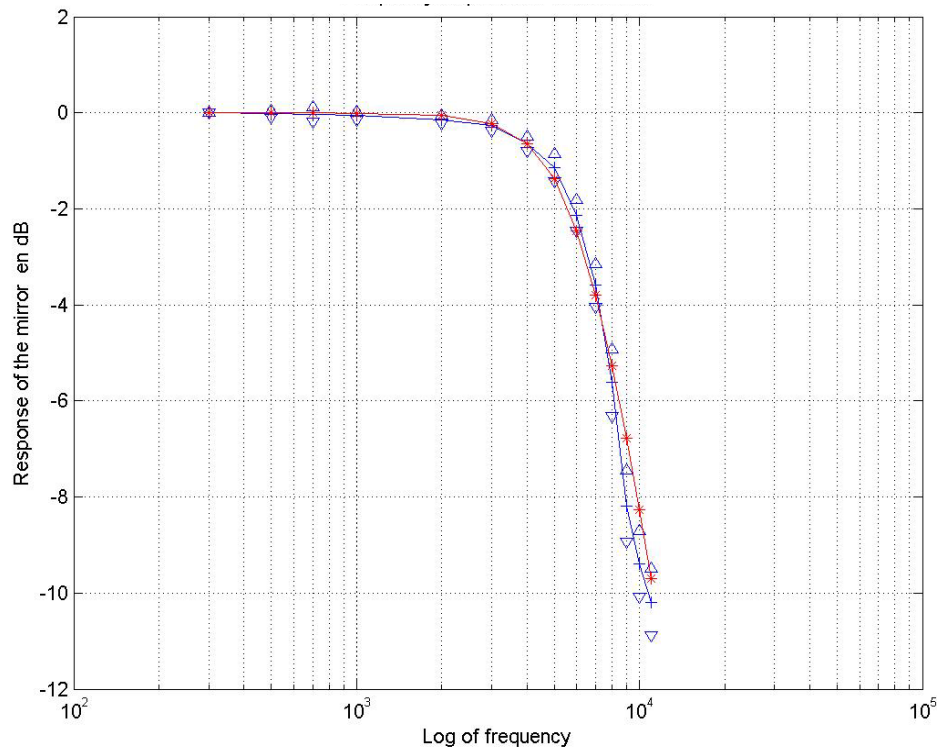
Influence function at 25V

**For a larger stroke with a low driving voltage:
change material**

- ◆ First polymer actuators realized
- ◆ Piston stroke: **2 μm at 30V**



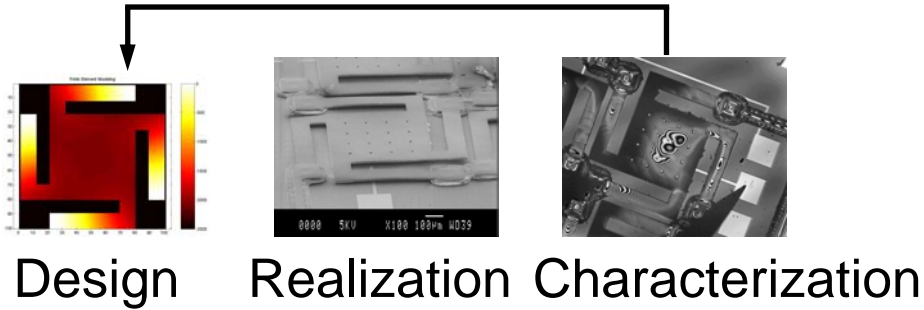
- ◆ Dynamical response: resonance frequency at 6.5 kHz
- ◆ Linearized actuation with a 14 bits dedicated electronics



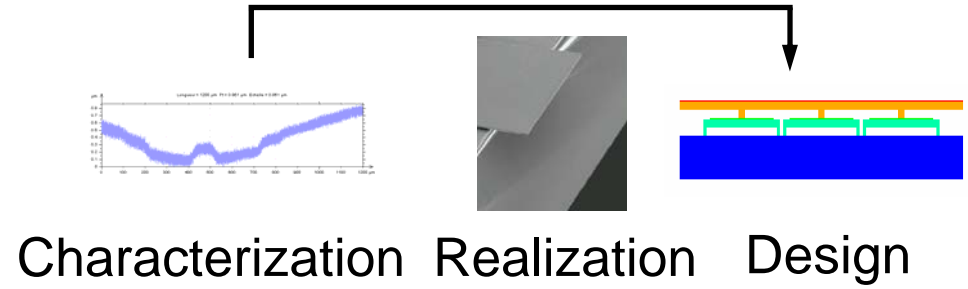
See also Liotard et al., LAM-LAAS paper at SPIE MOEMS conf. 2006, San Jose

MOEMS deformable mirror

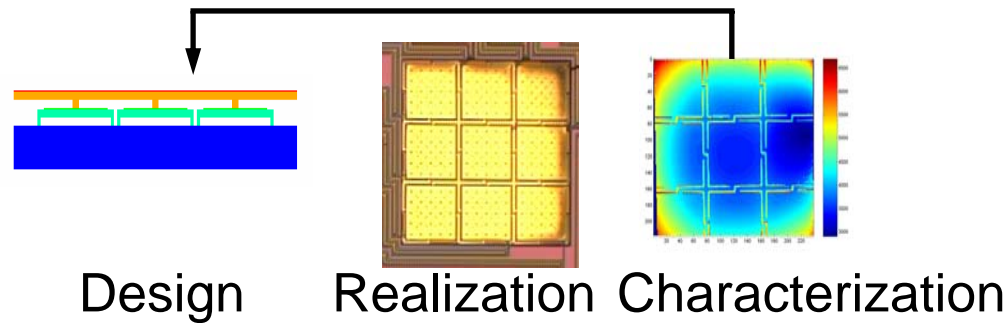
Actuator



Mirror

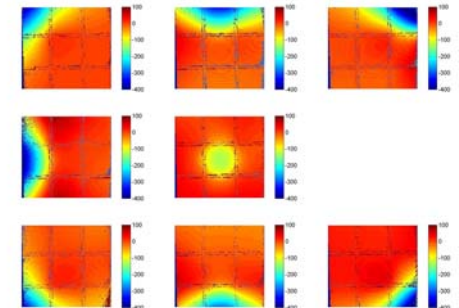


Actuators + Mirror

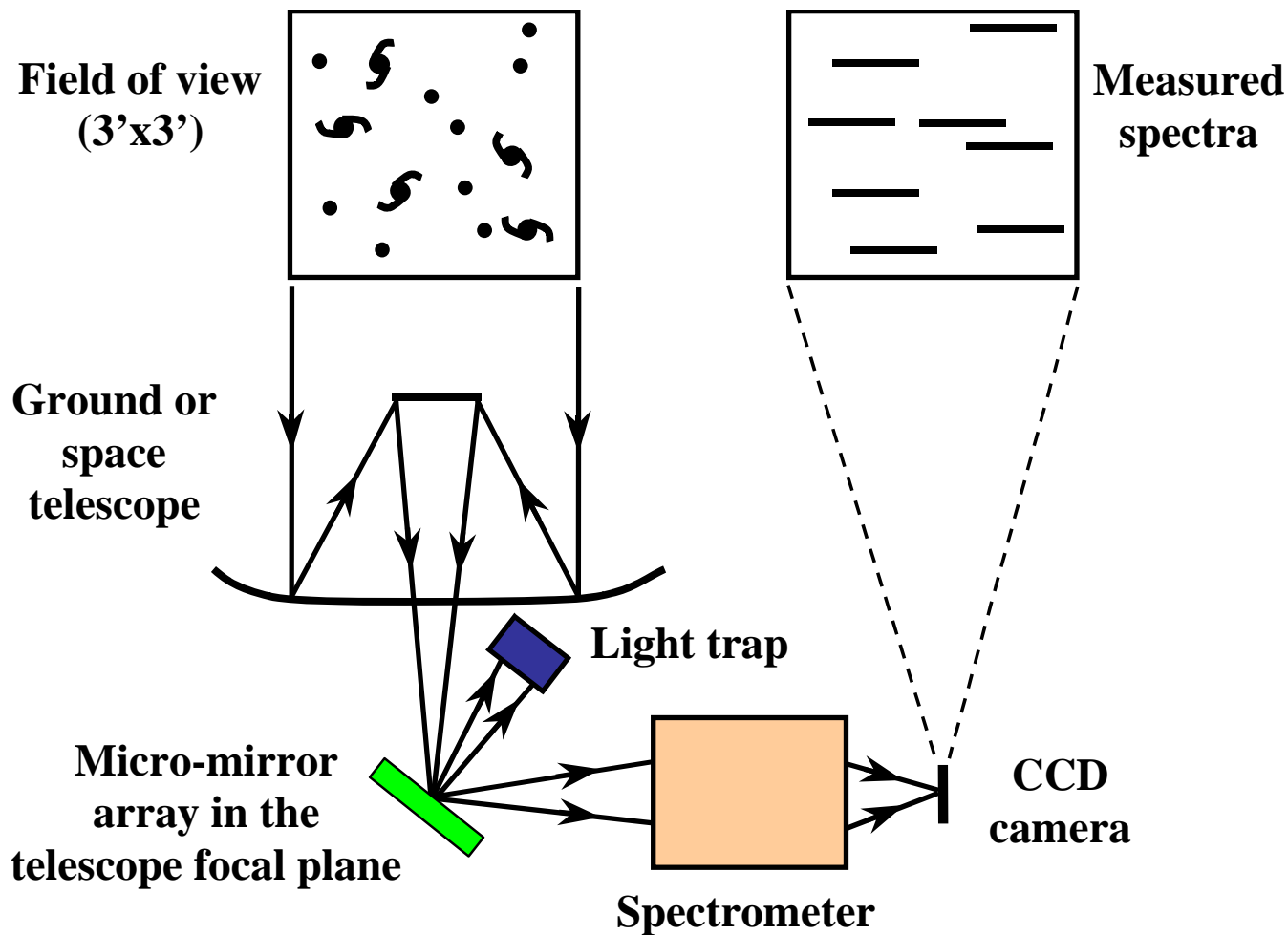


**MDM
prototype**

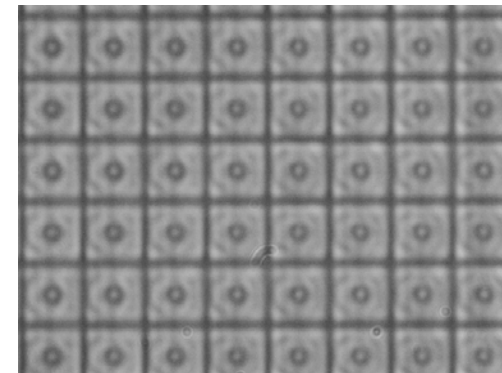
+



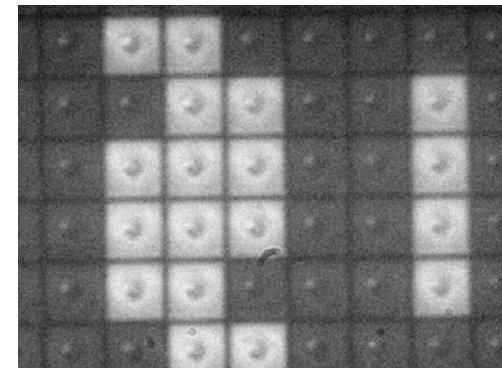
Multi-Object Spectrograph



Park position

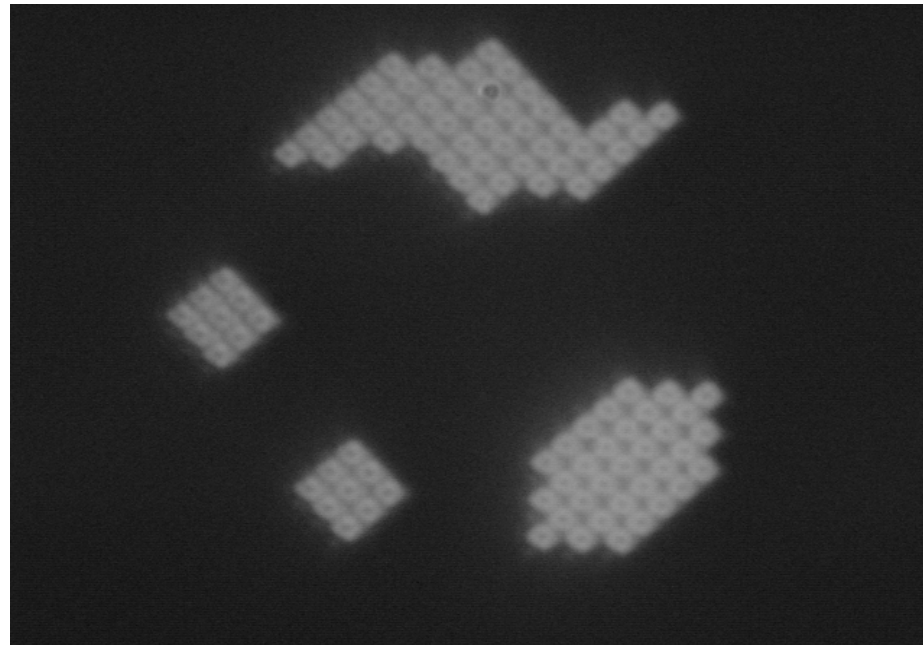


— 20 μm

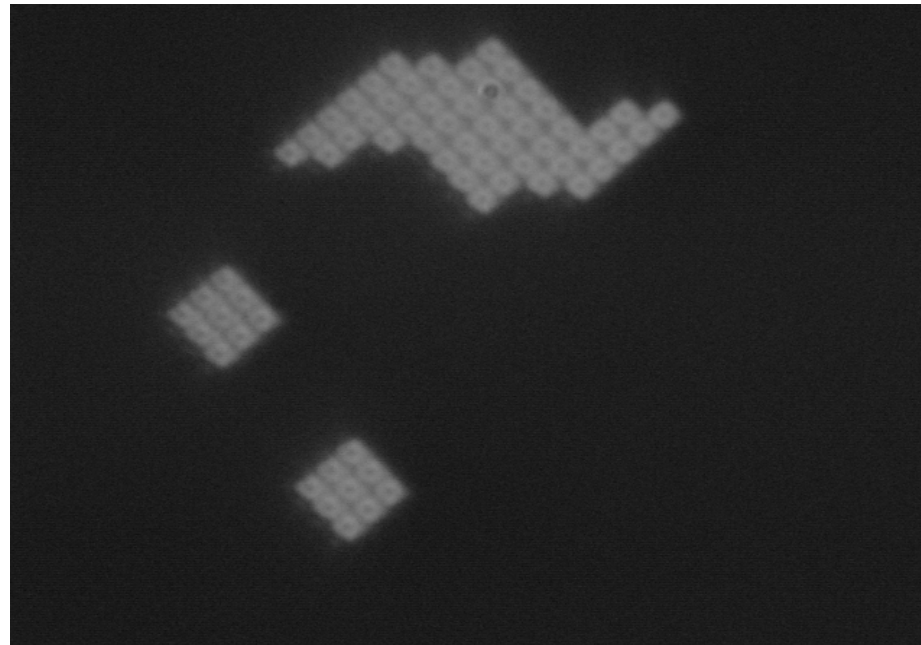


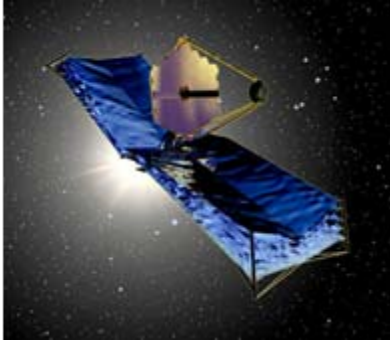
In action

Object selection

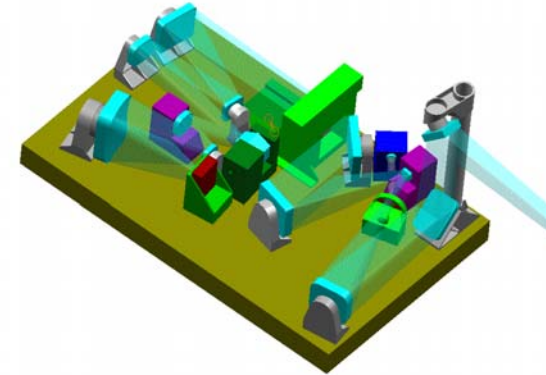


Object selection





- ◆ Visible + NIR Camera
- ◆ **Multi-Object NIR Spectrograph**
 - ❑ 0.6 - 5 μm
 - ❑ FOV : 3' x 3' (sampling 0,1")
 - ❑ R = 100 et R = 1000
- ◆ MIR Camera / Spectrograph



ESA pre-phase A, phase A and phase B1 studies

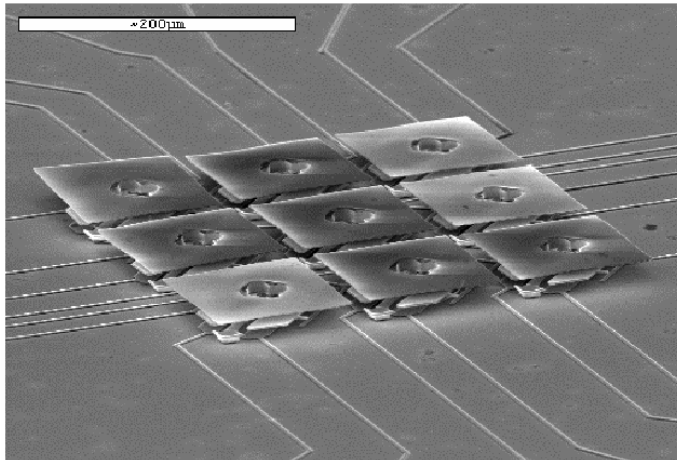
Instrument simulation

- ◆ Tool box
 - ❑ Photometry
 - ❑ Instrument design impact
 - ❑ Encircled energy

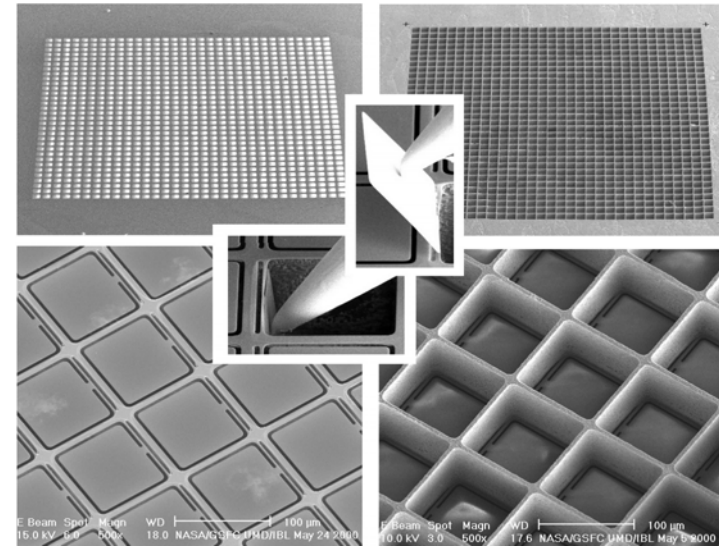
MOEMS slit masks

- ◆ Characterization
 - ❑ Surface quality
 - ❑ Contrast
 - ❑ Operability

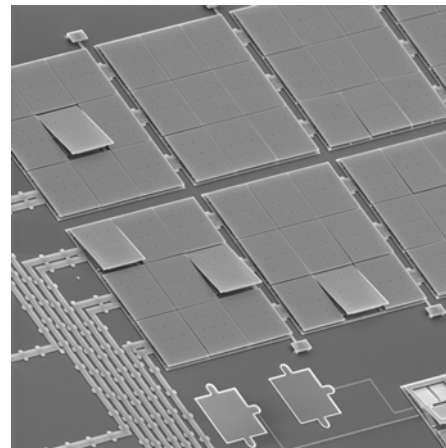
◆ NASA-GSFC Micro-mirrors



◆ NASA-GSFC Micro-shutters



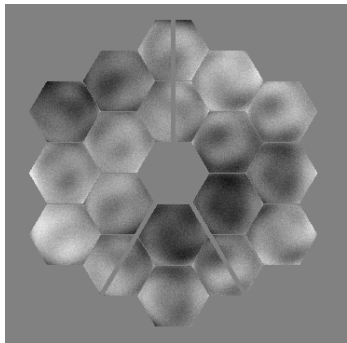
◆ Sandia Labs Micro-mirrors



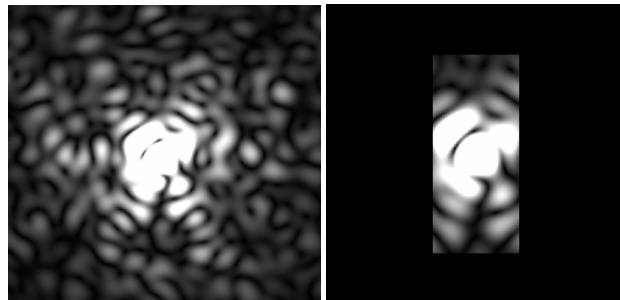
Programmable slit modeling

Fourier model

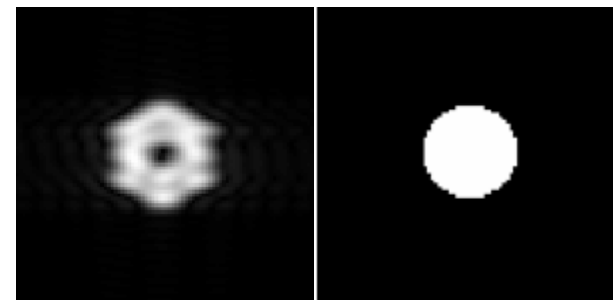
Telescope (JWST)



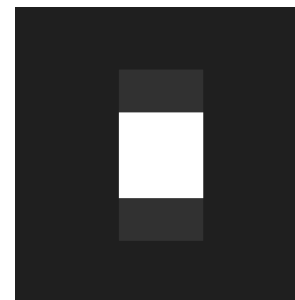
Field



Pupil



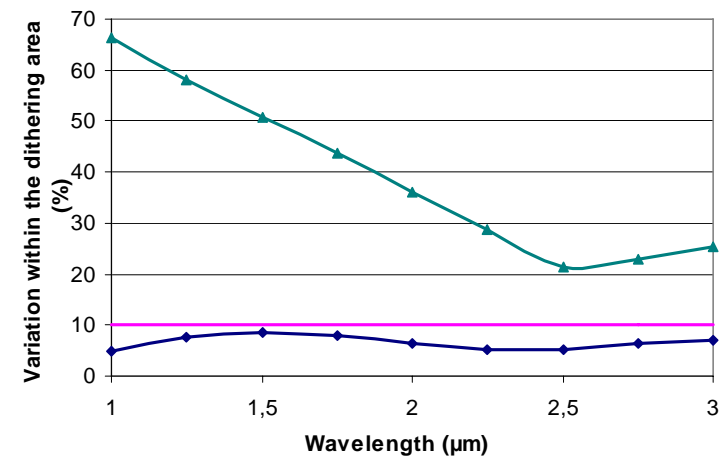
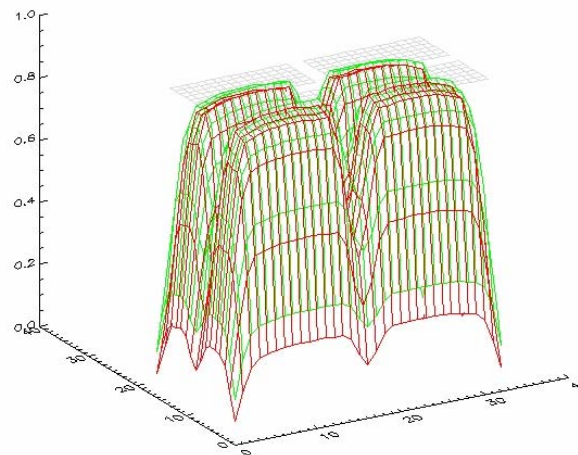
Detector



Spectral photometric variation modeling

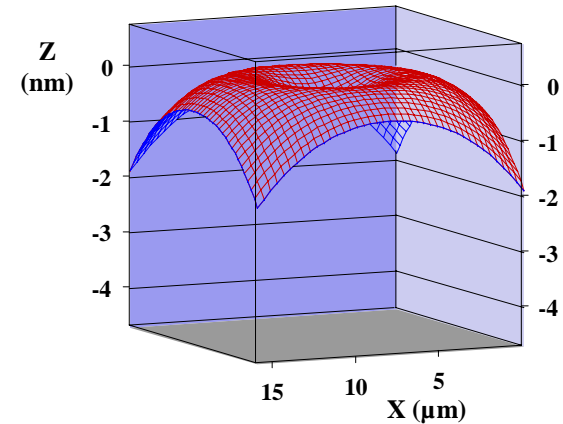
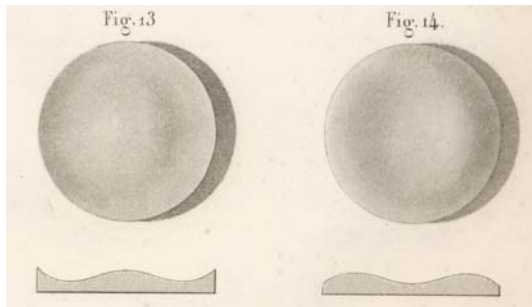
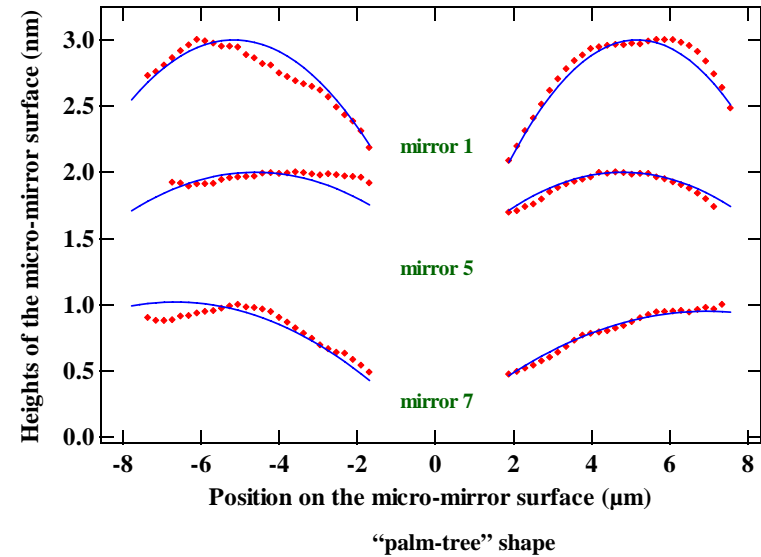
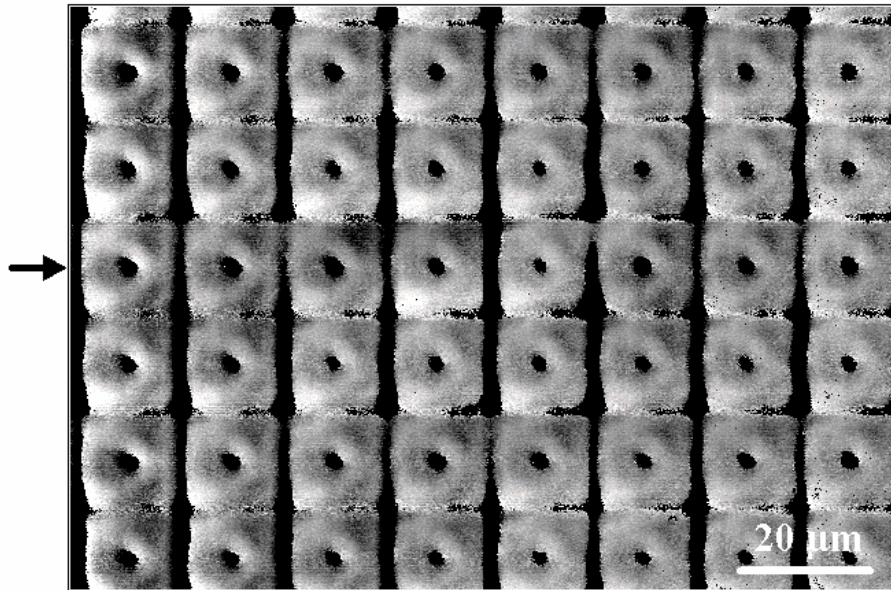
Geometrical effect

Diffraction effect



Surface measurement

Micro knife-edge test



Characterization bench

Characterization bench



Vacuum and cryo test facility
(down to 30K) under development

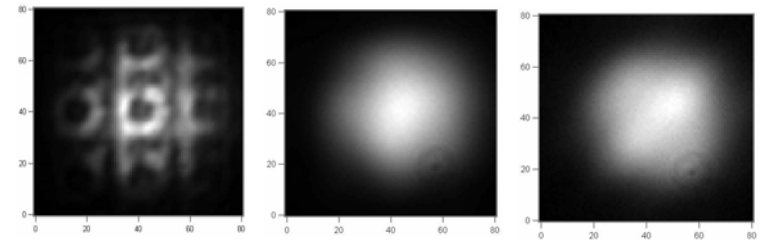
Images

$$F_{IN} = F/20$$

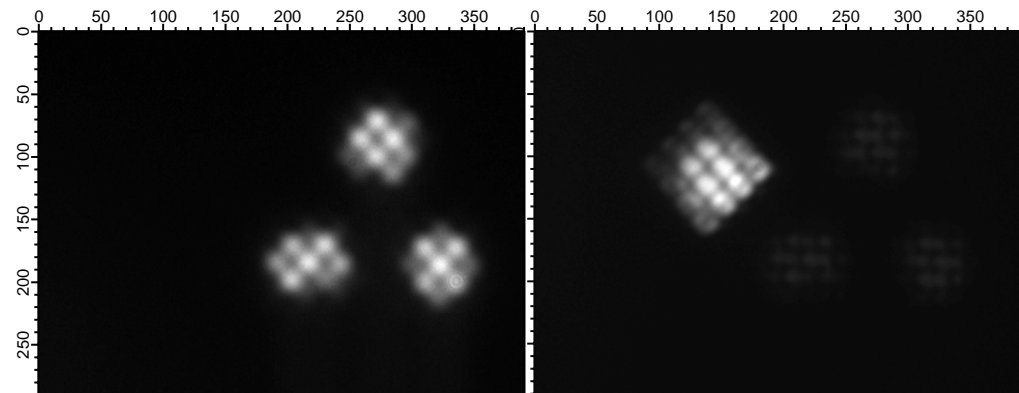
F/2

F/17 ON

F/17 OFF

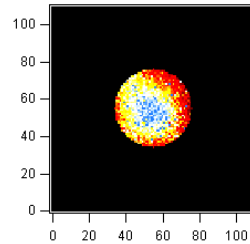
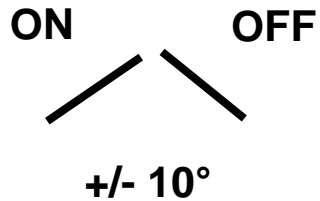


Field simulator

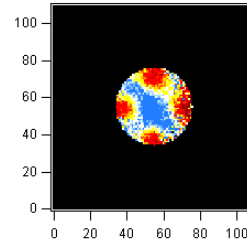


Contrast measurements

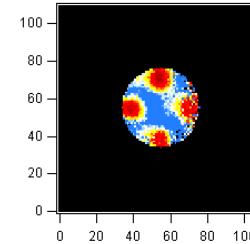
Micro-mirrors



F/34



F/34 +50%



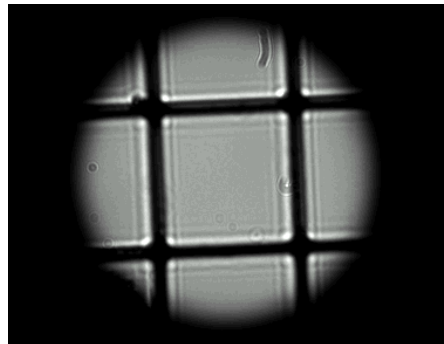
F/34 +100%



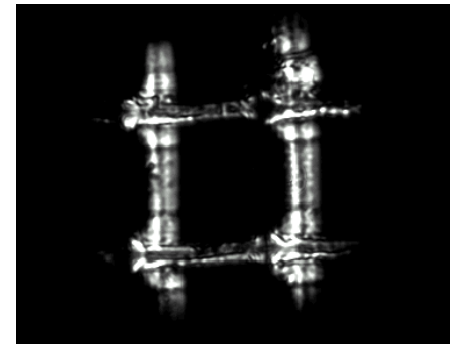
Effective Contrast
> 3000

Micro-shutters

OPEN



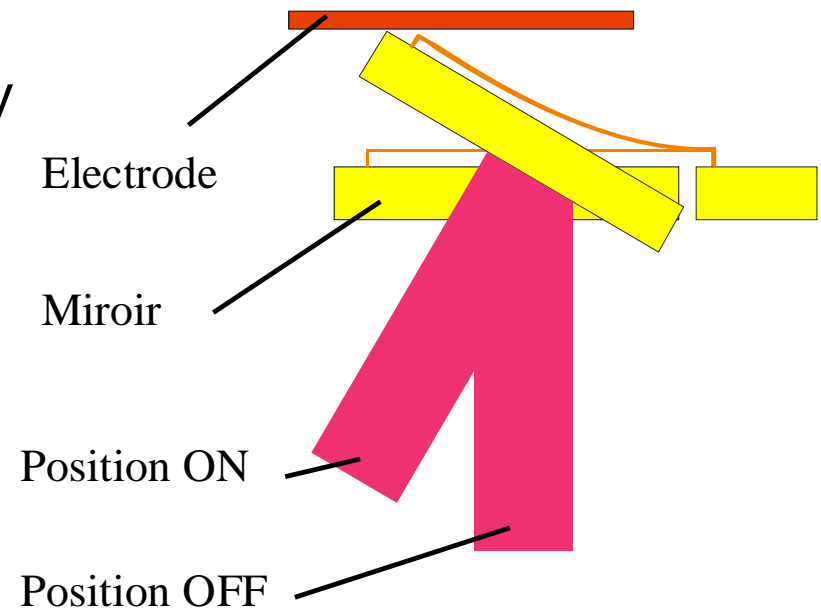
CLOSED



Collaboration engaged with University of Neuchâtel (Switzerland) for the development in Europe of programmable slits with MOEMS technology (FP6/Opticon/JRA Smart Focal Planes)

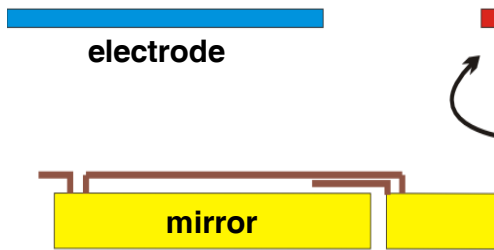
□ Micro-mirrors key parameters

- Design based on underneath cantilevers or torsion bars
- Deflection angle: 20° (ON position)
- Micro-element size: $100\ \mu\text{m} \times 200\ \mu\text{m}$, or bigger
- Fill factor: $> 90\%$
- Mirror surface of good quality

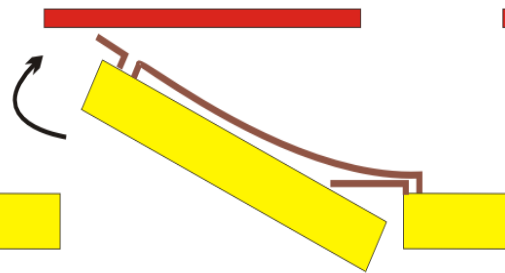


□ Principle

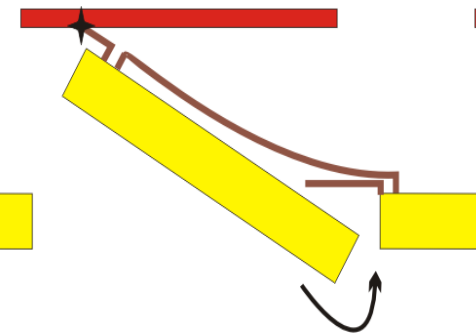
a) Rest position



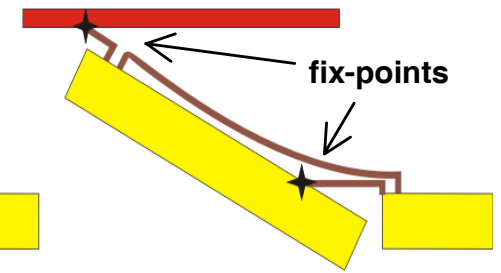
b) 1st tilting



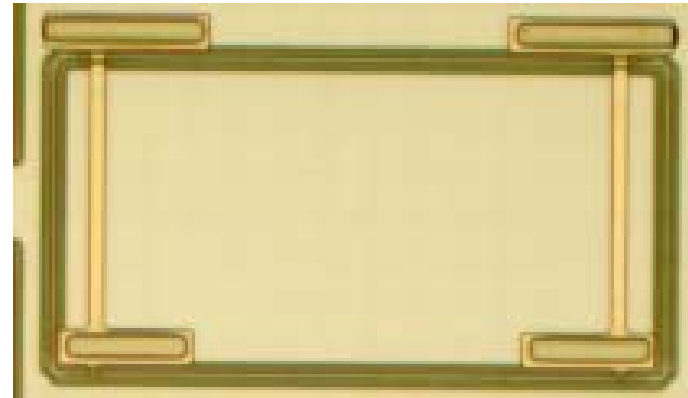
c) Stopper & 2nd tilting



d) Electrostatic latching



□ Realization

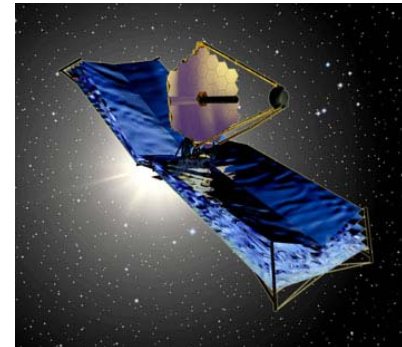


See also Waldis et al., IMT-LAM paper at SPIE MOEMS conf. 2006, San Jose

Projects

◆ Programmable slits

- ❑ JRA Smart Focal Planes 2004-2007
- ❑ JWST (NGST) 2011



◆ Micro-deformable mirrors

- ❑ JRA Adaptive Optics 2004-2007
- ❑ VLT 2nd generation (FALCON) 2004-
- ❑ OWL 2015-2020

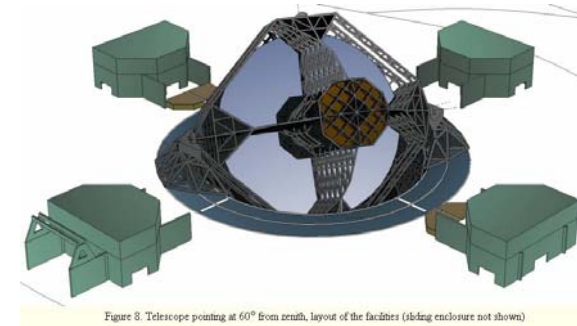


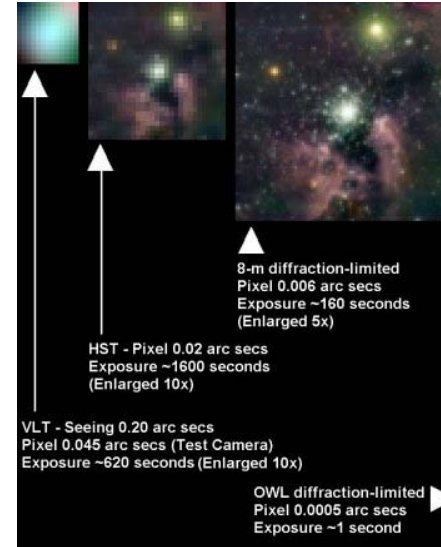
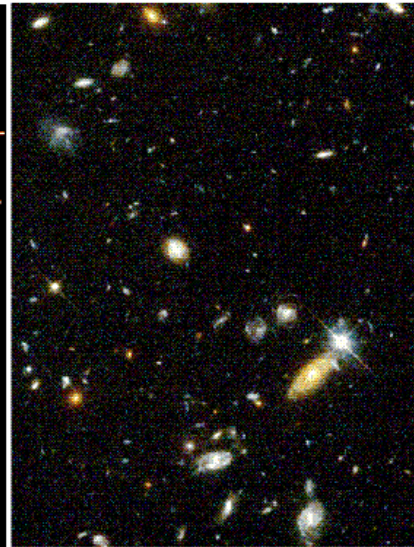
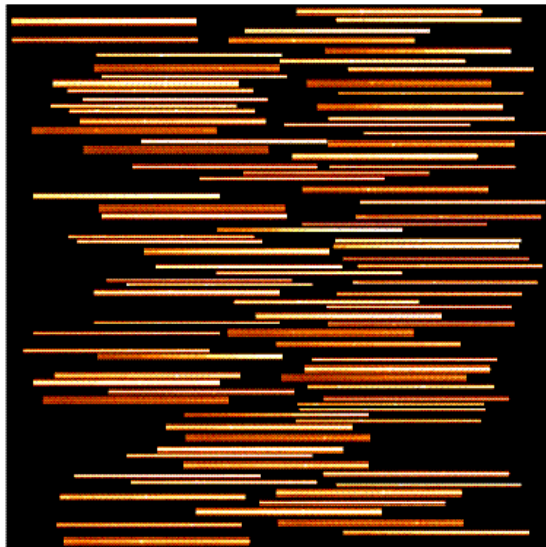
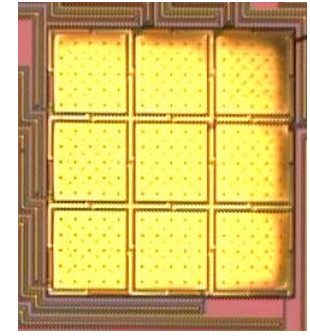
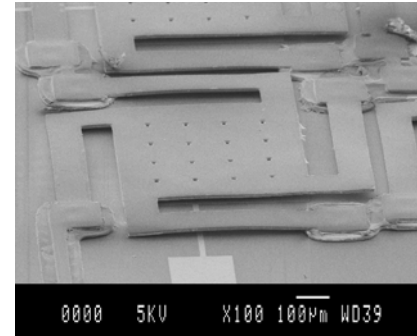
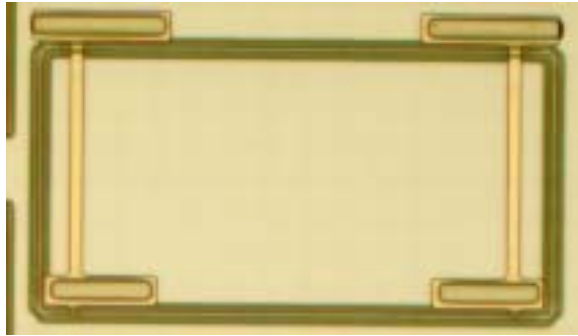
Figure 8. Telescope pointing at 60° from zenith, layout of the facilities (siding enclosure not shown)

◆ ESA study on MOEMS Delta space qualification methodology



- ❑ Study driven by Alcatel Alenia Space March 2005 – end 2006
- ❑ LAM expertise for Earth and Space Observation

MOEMS for astronomy



Funding: INSU, CNRS, Min. Recherche, CNES, ESA, UE, Région PACA, CG13

frederic.zamkotsian@oamp.fr

Bibliography



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