

SEE Modelling / Prediction

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“LET-Requirements and Testing for Space Applications”

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Outline

- Basics of SEE rate Modelling
- Playing with RPP method using CREME96
- Examples; Device data with IRPP method using CREME96
- Conclusions

This presentation addresses only Heavy ion SEE modelling

All predictions have been made in CREME96. We expect similar results in any other tool using RPP/IRPP method

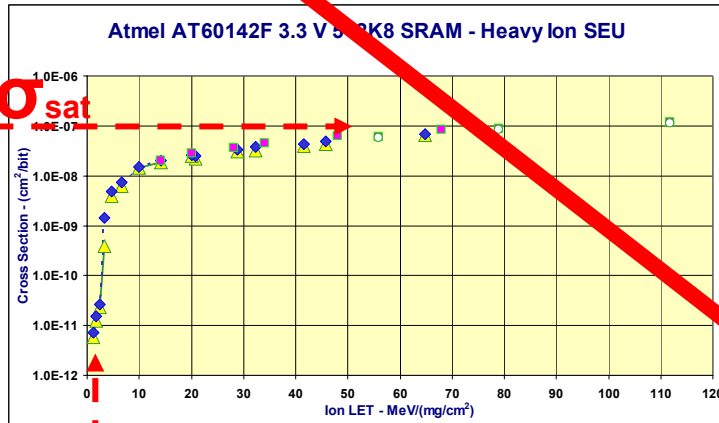
Correlation to flight data is unknown. It only shows the outcome of the modelling.

Basics of Heavy ion SEE rates Predictions

Accelerator Test Data

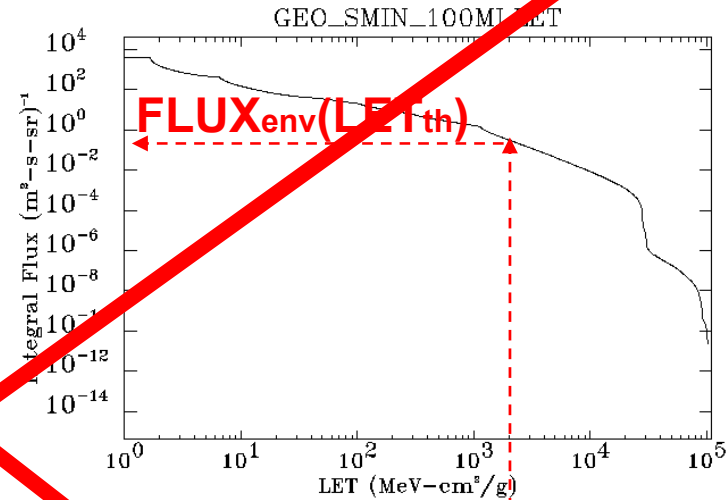
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Space environment Model



σ_{sat}

LET_{th}



$FLUX_{env}(LET_{th})$

Wed Jan 17 07:56:30 2007

LET_{th}

\Downarrow

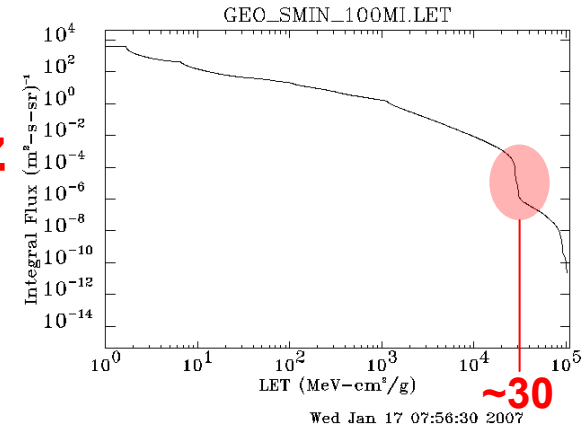
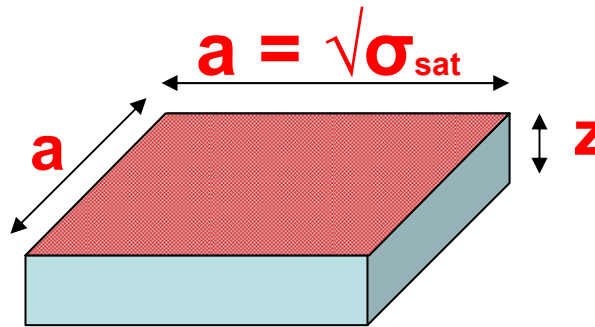
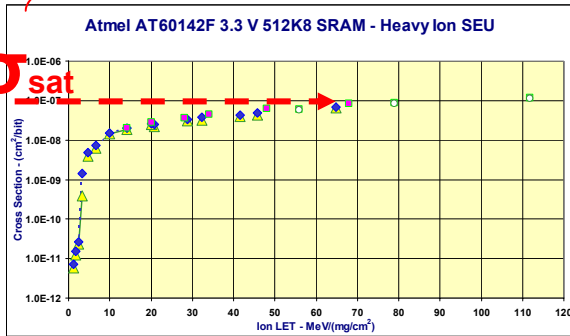
Predicted SEE rate in Space = $FLUX_{env}(LET_{th}) * \sigma_{sat}$

This is a 2D - Model

Space is 3-dimensionell



Basics of Heavy ion SEE rates Predictions 3D-Modelling (RPP&IRPP Method)



Accelerator Test Data + Sensitive Volume(SV) + Space environment
(RPP dimension)

RPP / IRPP method:



Integrating the environment in all directions into the SV, deriving if specific Ion path leave charge enough for an upset. Critical charge is derived from accelerator test data (IRPP)/ LET_{th} (RPP).

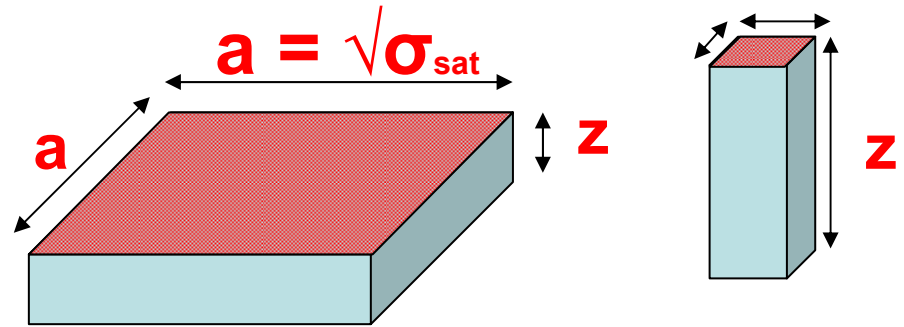
Predicted SEE rate in Space

Playing with RPP method using CREME96

(Critical charge method)

➤ All calculations made in CREME96

- Geosynchronous Orbit
- Atomic number 1-92
- Solar Quiet conditions
 - Solar Minimum
- 100mil Aluminium shielding



➤ Device Input parameters

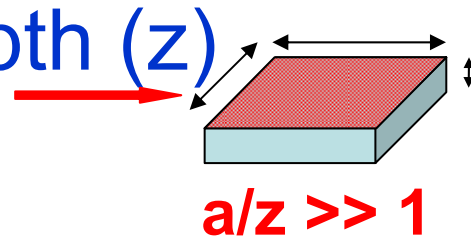
- σ_{sat} Saturation cross section
- LET_{th}
- Z Sensitive depth

$a/z \gg 1$

$a/z < 1$

=> Critical charge for upset $Q_c(Si) = 0.0103 \times LET_{th} \times Z$

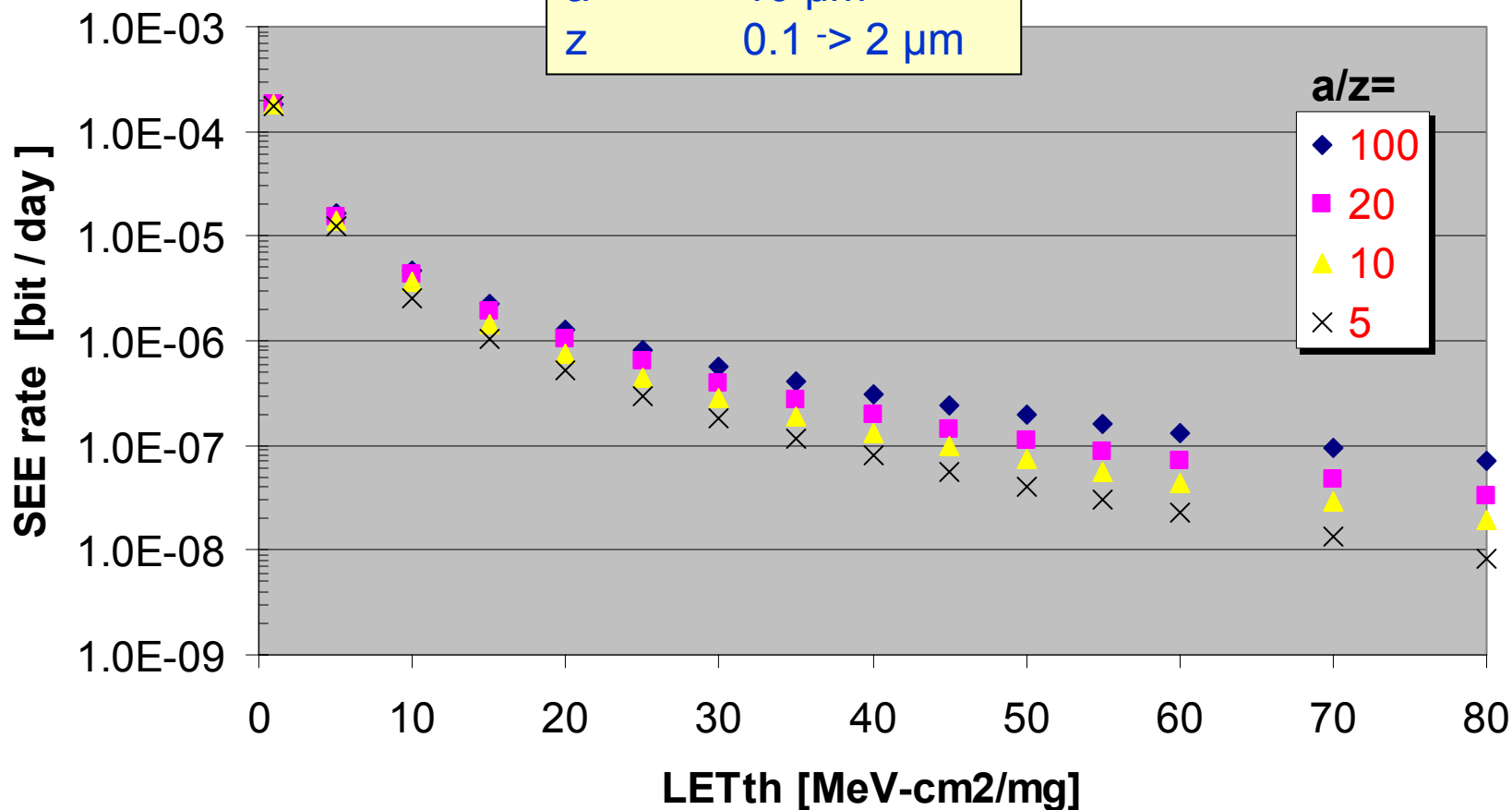
Playing with the sensitive depth (z) at a big flat SV



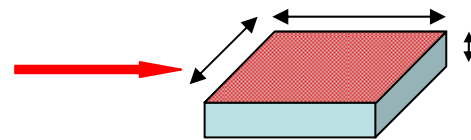
Horizontal direction to SV
worst case path

=> small dependency on z

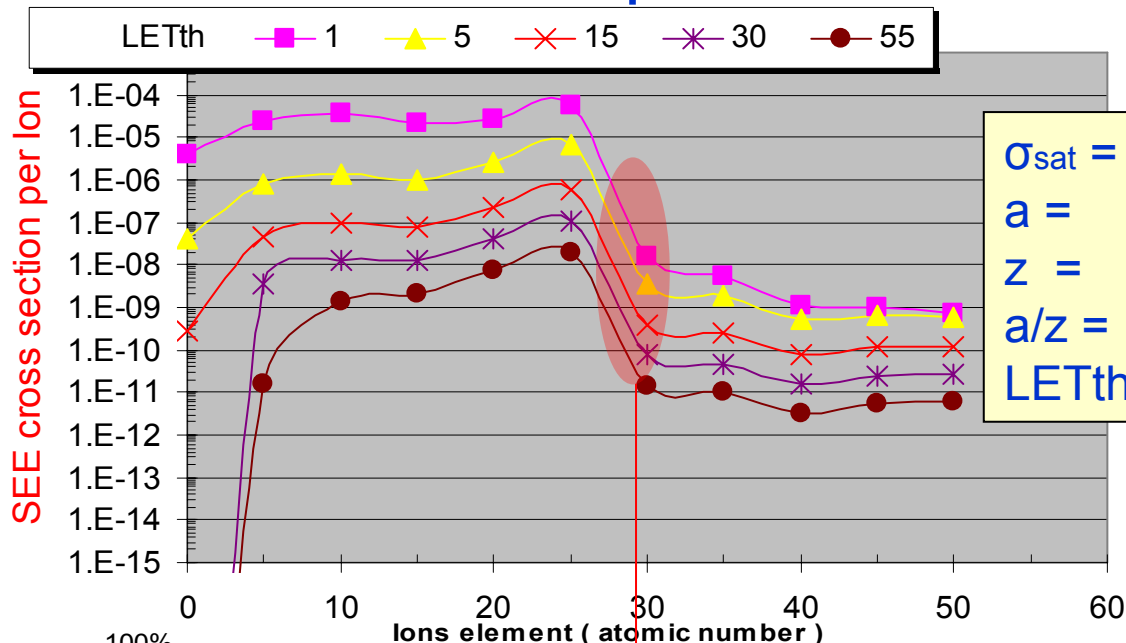
$\sigma_{\text{sat}} = 100 \mu\text{m}^2$
 $a = 10 \mu\text{m}$
 $z = 0.1 \rightarrow 2 \mu\text{m}$



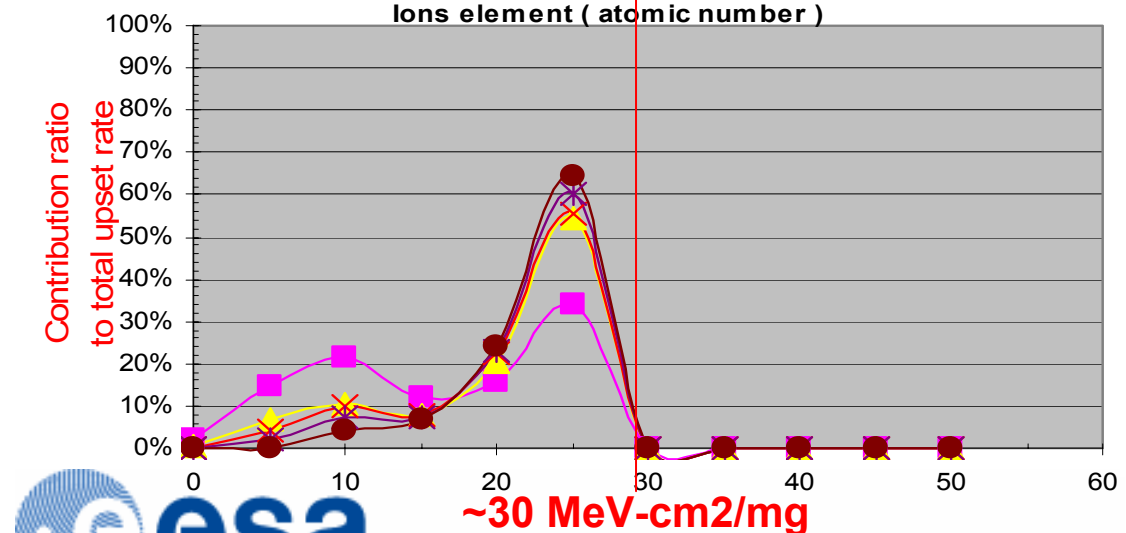
Contribution to SEE rate of Each Ion Element of Space Environment



$a/z \gg 1$



$\sigma_{sat} = 100 \mu\text{m}^2$
 $a = 10 \mu\text{m}$
 $z = 2 \mu\text{m}$
 $a/z = 5$
 $\text{LET}_{th} 1 \rightarrow 55 \text{ MeV-cm}^2/\text{mg}$

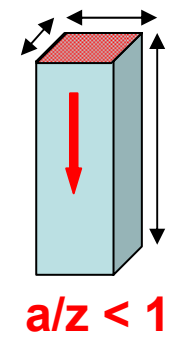
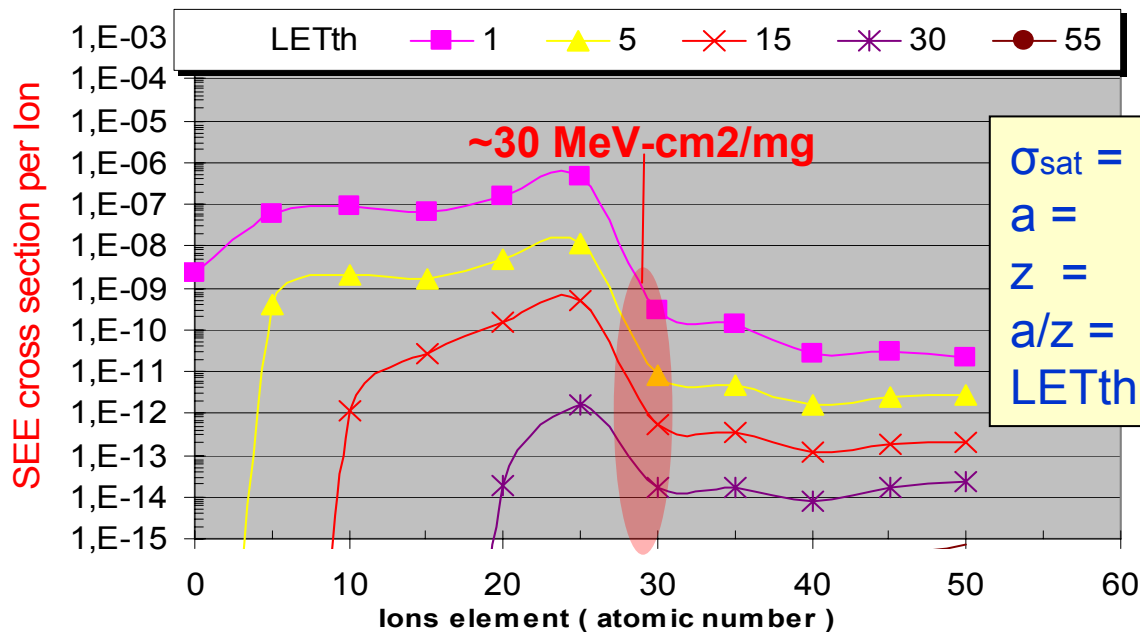


$\sim 30 \text{ MeV-cm}^2/\text{mg}$

- Even for LET_{th} of 55, the main contribution of upsets comes from lighter ions (with max LET of 30)
 - Horizontal ion paths into the SV give the main contribution to SEE rate
 - This is a worse path compared to acceleration data.

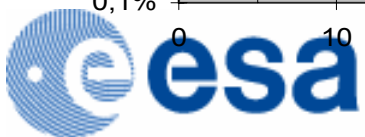
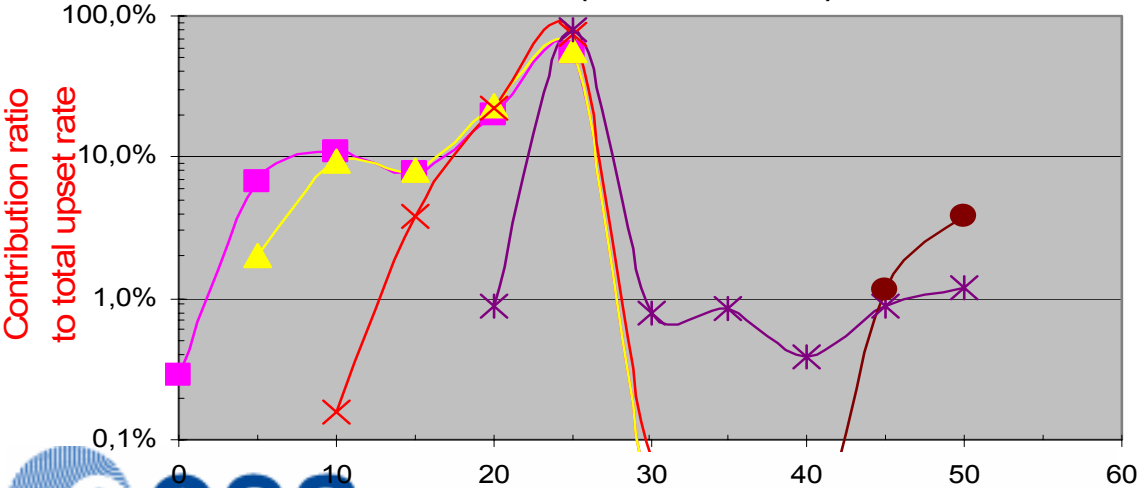


Contribution to SEE rate of Each Ion Element of Space Environment



➤ For LET_{th} of 55, no contribution of upsets comes from lighter ions (with max LET of 30)

- Horizontal ion paths into the SV give no contribution to SEE rate
 - The worse ion path is vertical into SV. Same as in accelerator test data



Playing with the sensitive depth (z)

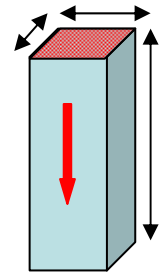
at a small SV

Vertical direction to SV worst case path

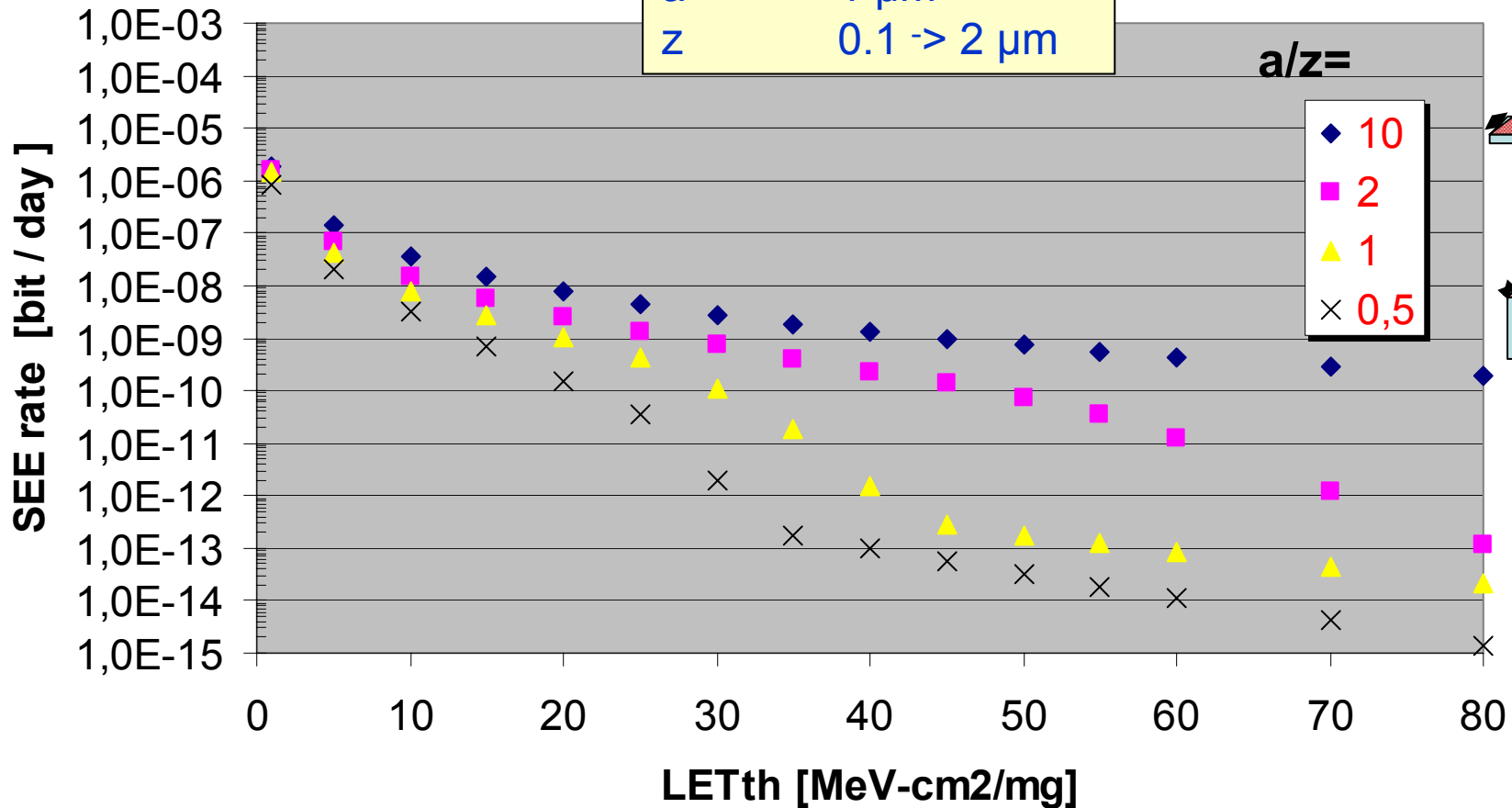
⇒ Huge dependency on z

⇒ Especially for higher LETth

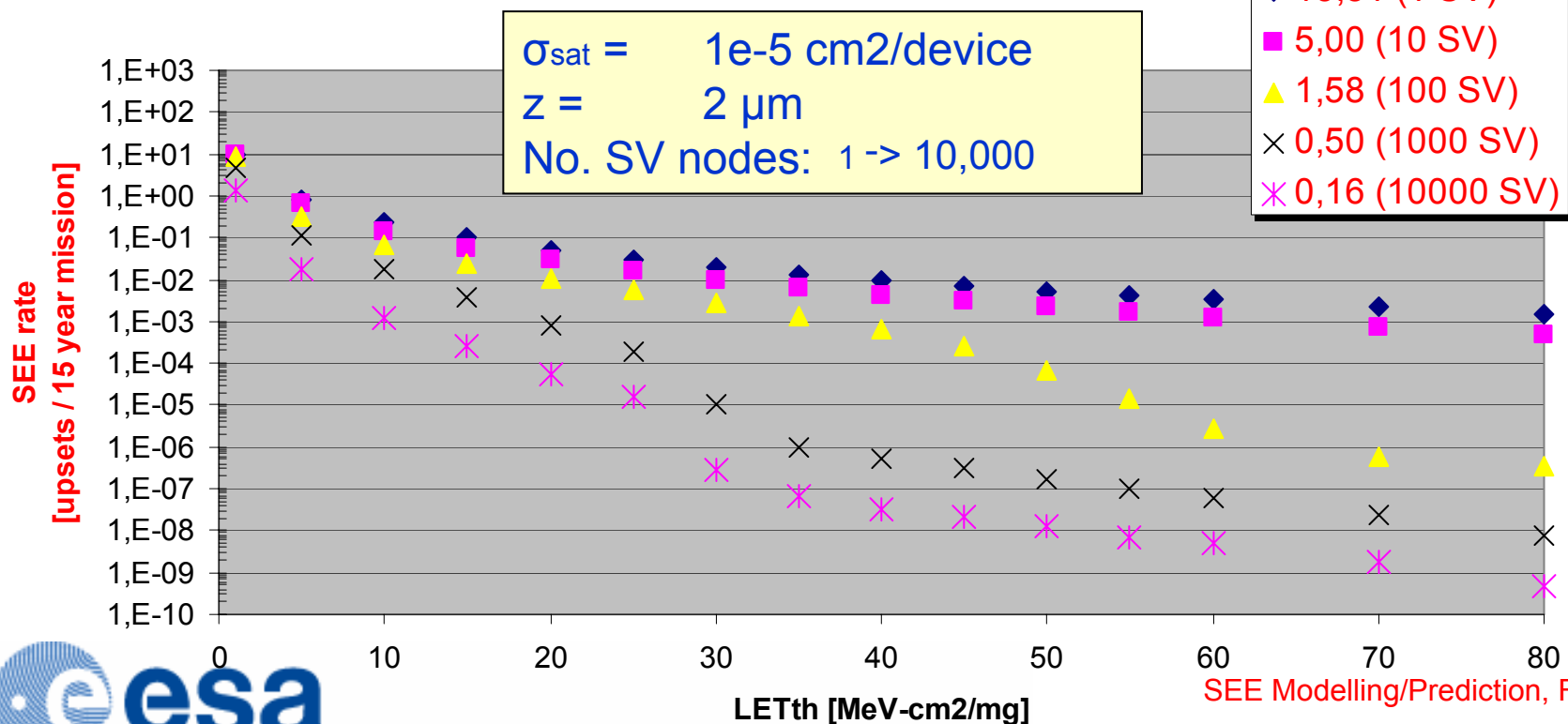
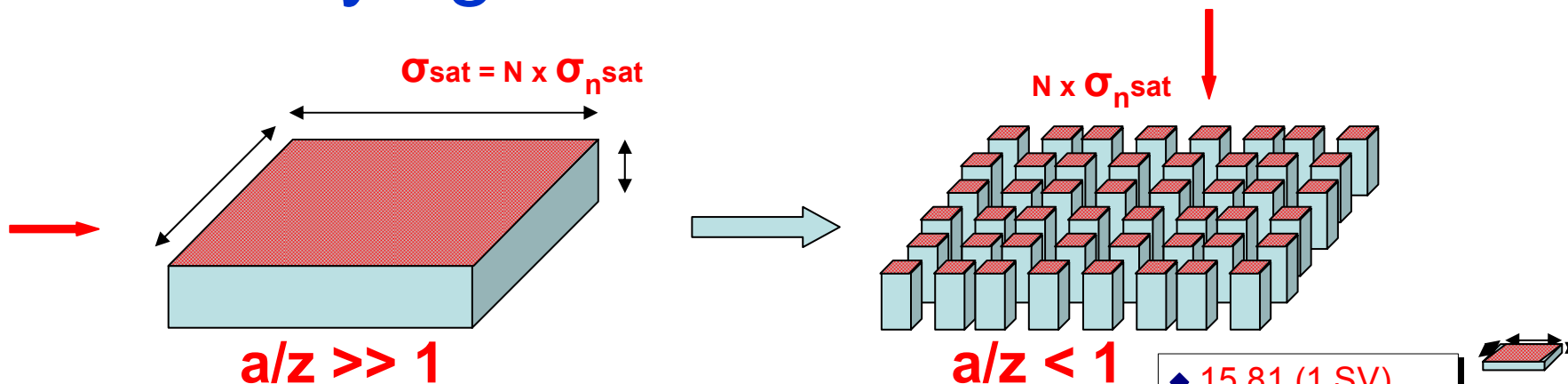
$\sigma_{\text{sat}} = 1 \mu\text{m}^2$
 $a = 1 \mu\text{m}$
 $z = 0.1 \rightarrow 2 \mu\text{m}$



$a/z < 1$

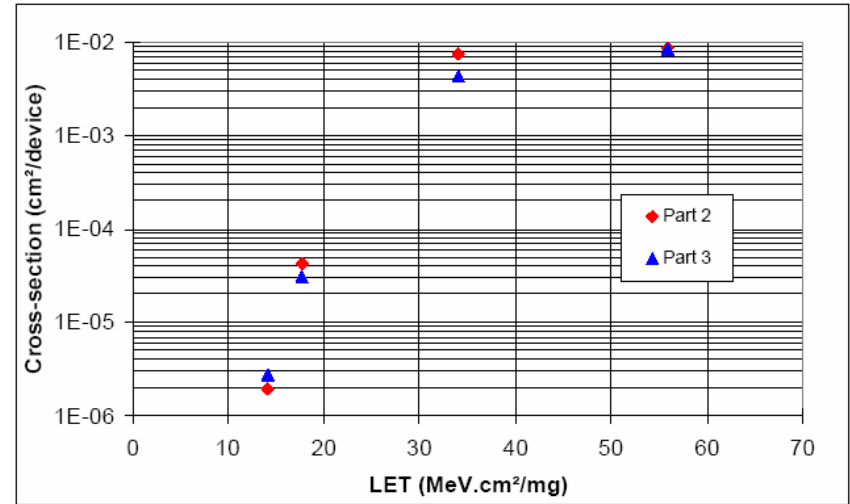
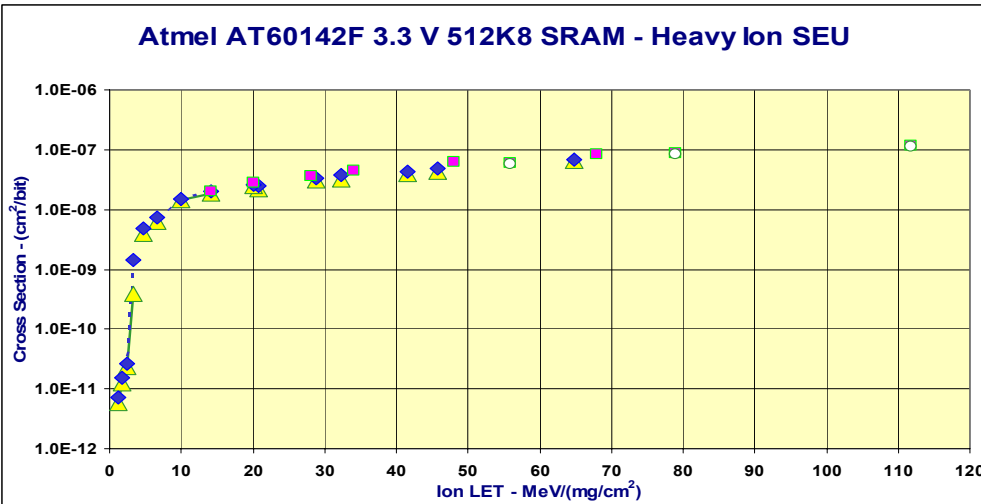


Playing with the number of SV



Example data using Integral-RPP method (IRPP)

➤ Weibull parameters into same space environment (CREME96)



Atmel AT60142F 512K8 SRAM

Used weibull fit:

- Onset = 1.5 MeV-cm²/mg
- Power = 1.29
- Width = 436.4
- σ_{sat} = 6e-7 cm²/bit

Atmel AT40K FPGA -Cfg memory cells

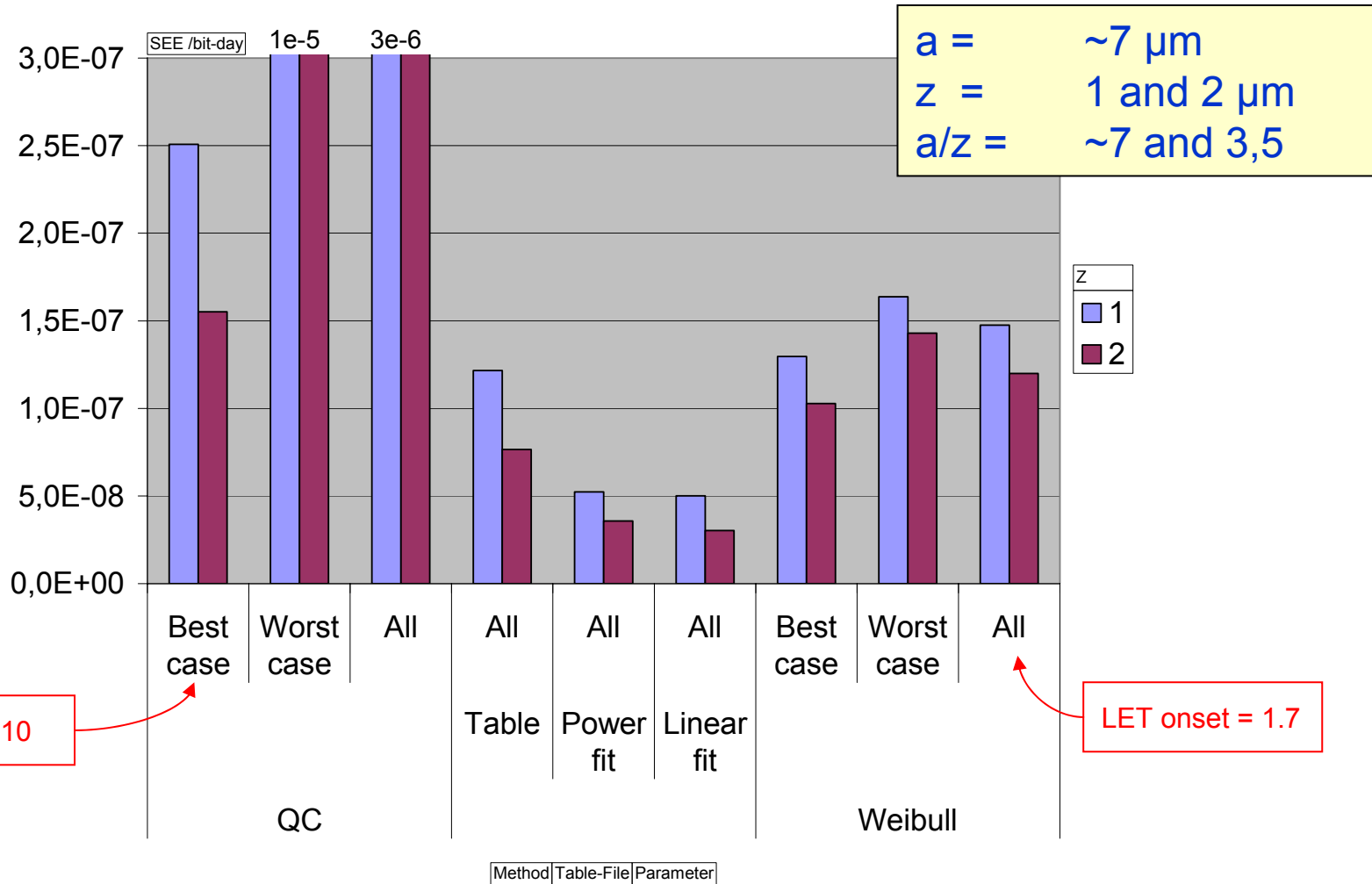
Used weibull fit:

- Onset = 9.7 MeV-cm²/mg
- Power = 4.6
- Width = 25
- σ_{sat} = 2.5e-8 cm²/bit

Predicted SEE rates with different fitting of test data using CREME96

AT60142F

Bits / Device (Alla) Funnel 0 Zname | devXS(cm2) (Alla) LETth (Alla)



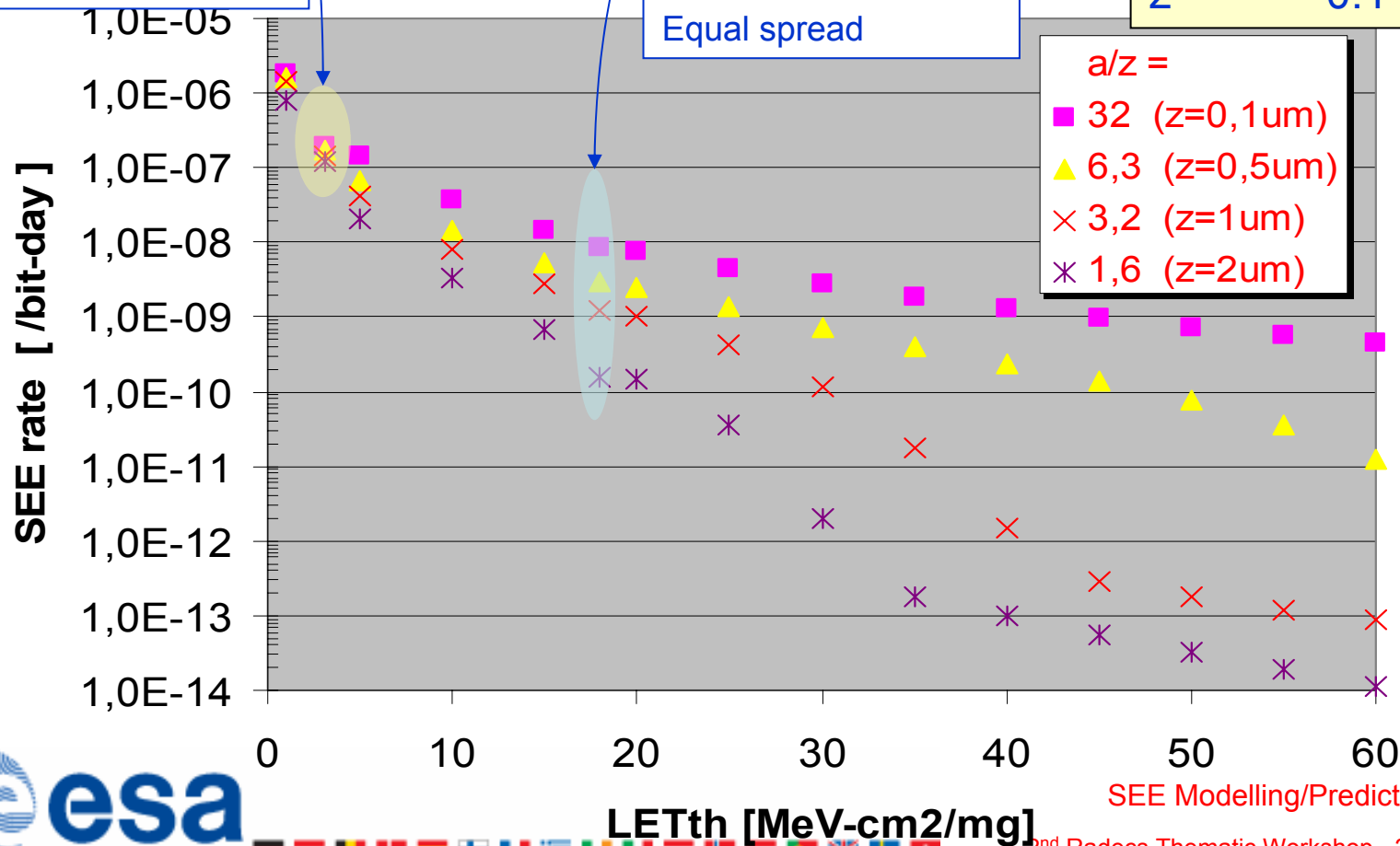
Playing with the sensitive depth (z)

IRPP Method

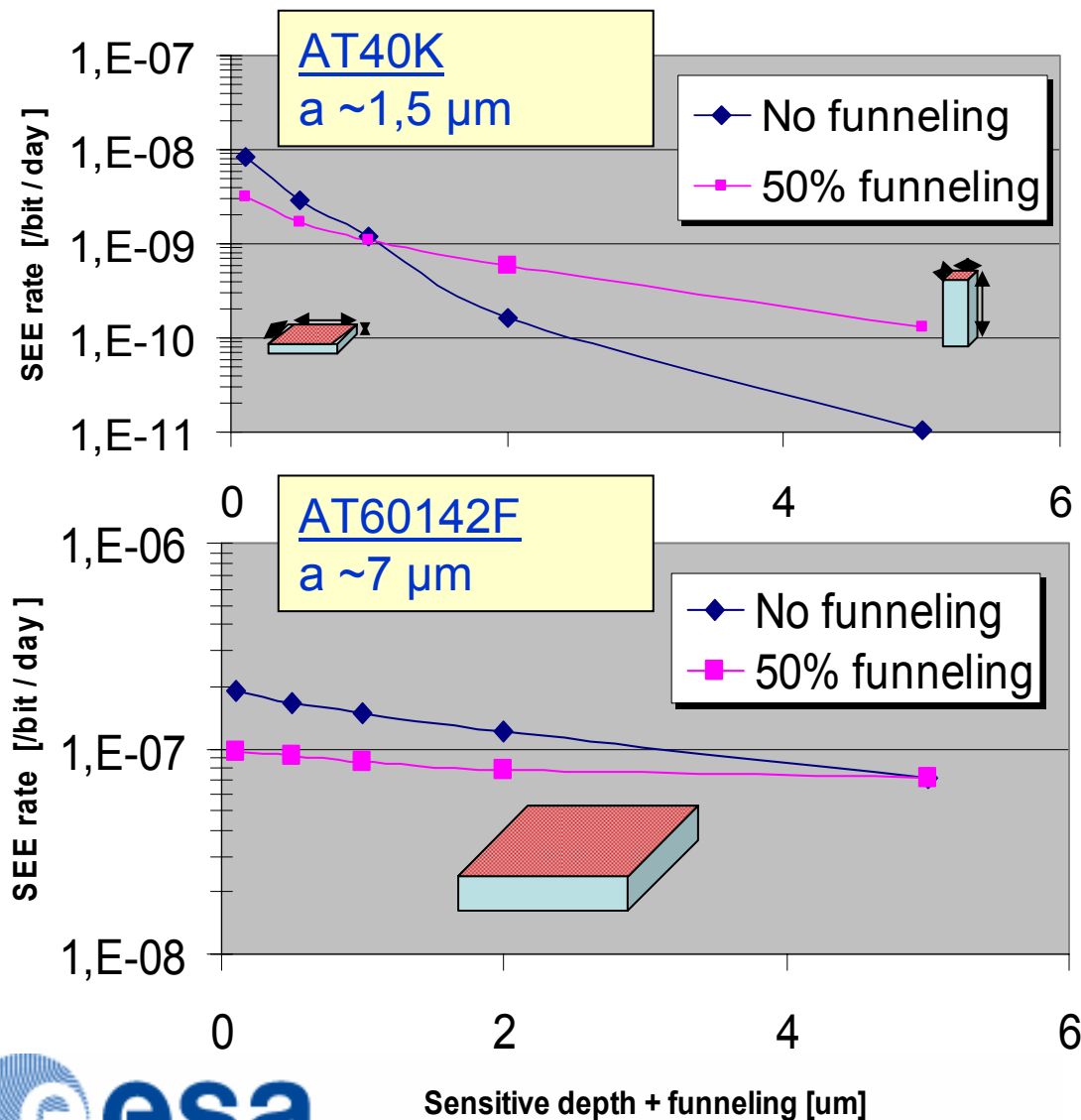
AT60142F
 $\sigma_{\text{sat}} = 60 \mu\text{m}^2$
 $a/z: 70 \rightarrow 3,5$
 Smaller spread

AT40K
 $\sigma_{\text{sat}} = 2,5 \mu\text{m}^2$
 $a/z: 15 \rightarrow 0.8$
 Equal spread

$\sigma_{\text{sat}} = 10 \mu\text{m}^2$
 $a = 3,33 \mu\text{m}$
 $z = 0.1 \rightarrow 2 \mu\text{m}$



Playing with the sensitive depth (z) and funnel IRPP Method



➤ Effect of Funnel switch “polarity” when going from flat SV to deep SV

Conclusions & Topics for the Round Table

- With the IRPP/RPP method, basically no ions above Iron contributes to the predicted SEE rate.
 - Is this also the case in space?
- For higher LET thresholds (above 20-30 MeV-cm²/mg) and shrinking technologies it is impossible to do reliable SEE rate predictions!?!
 - At least when we test as we do today
- We have requirements for LET threshold and Fluence (1E+7 ions/cm²)
 - But with shrinking technologies the final SEE rate is more dependent on the geometry of the sensitive volume
 - For older technologies it might have been relevant to use LET threshold as screening parameter, but how relevant is it for shrinking technologies?