



Microbolometers for mid, thermal, and far IR sensing

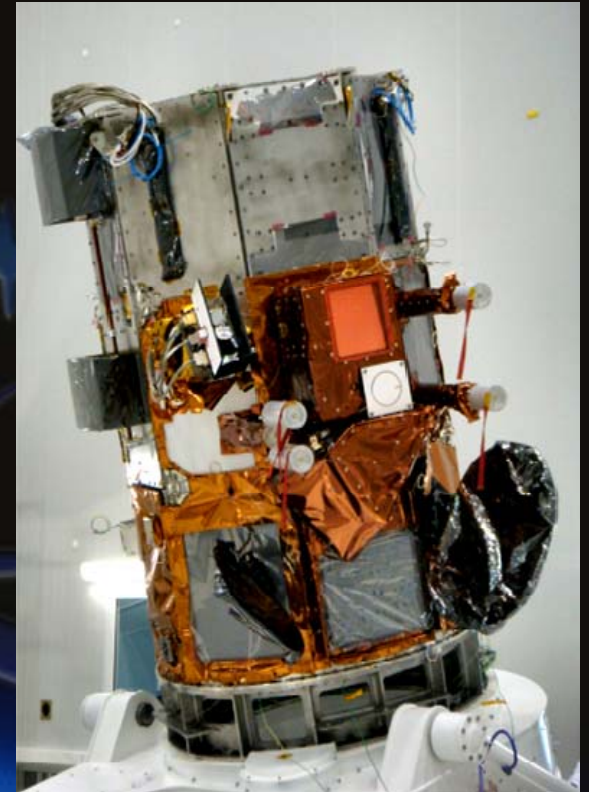
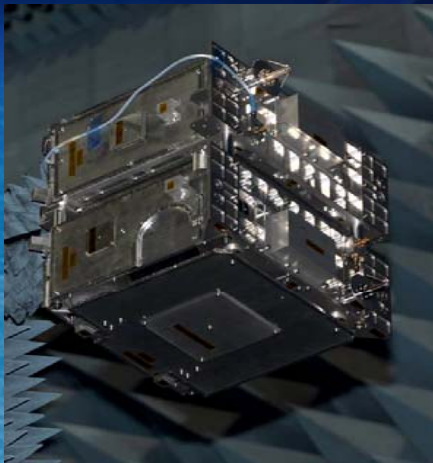
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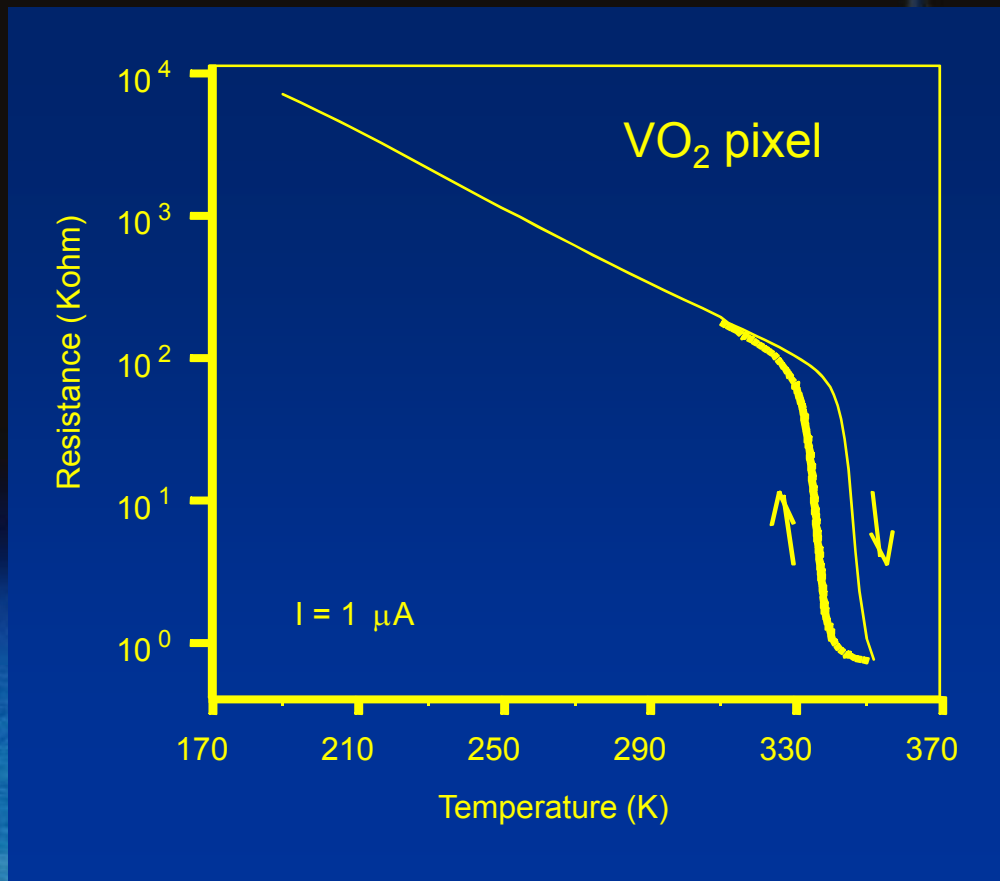
Introduction

- The advantage of microbolometer, beside the room temperature operation, is the absence of a cutoff wavelength
- Lately there have been requirements for coregistration in the MWIR, TIR, and FIR
- Technology development at the CSA has focused on *linear* arrays of microbolometers
- Sensors were developed for two missions: SAC-D Aquarius (MWIR / TIR) and JC2Sat (FIR)
- Details on these sensors are presented





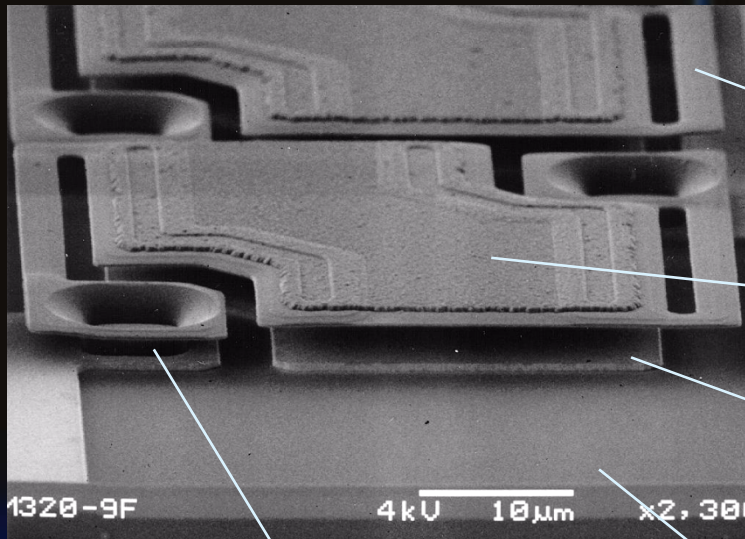
Bolometric sensing



- Thermal mechanism
- Spectral response defined by absorptance
- Absorptance can be adjusted using coating or cavity effect



Use of MNT



Si₃N₄ microbridge

Thin film bolometer and coating

Reflector

Monolithically embedded ROIC

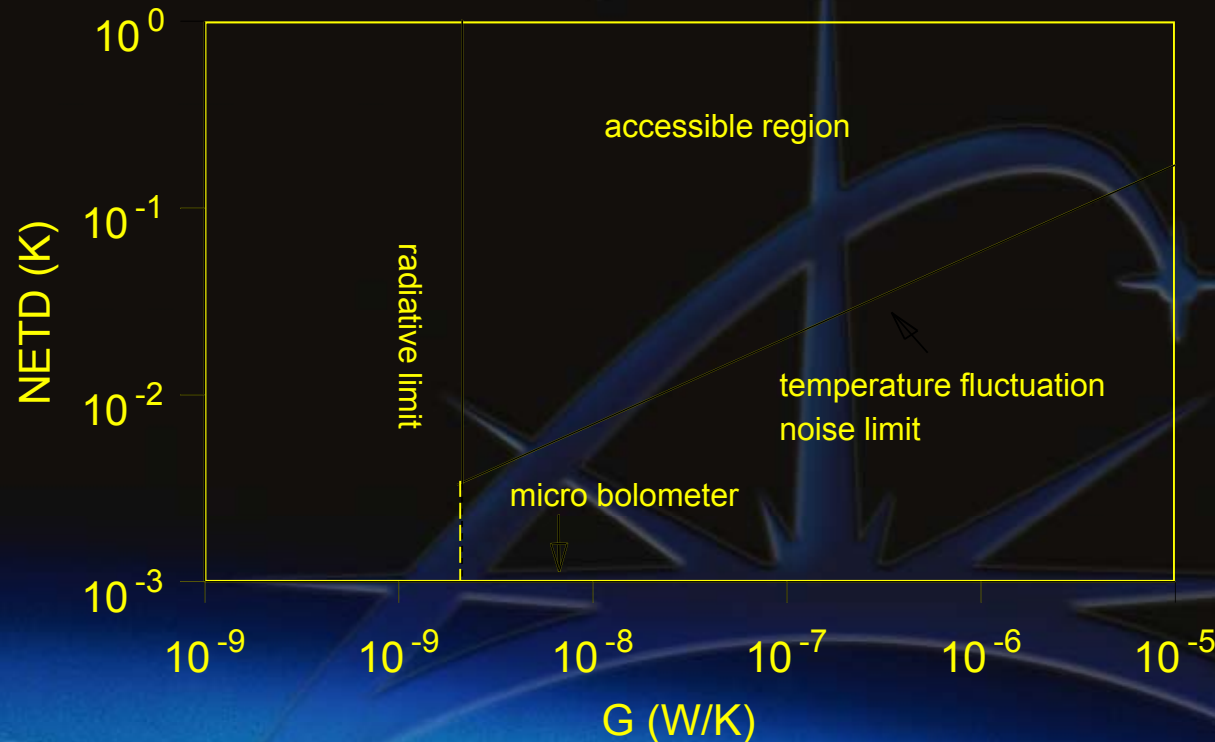
vias

$$r = \frac{I_b (dR / dT) \eta}{G (1 + \omega^2 \tau^2)^{1/2}}$$

▷ Vacuum gap provides cavity effect and thermal isolation



Performance limitations



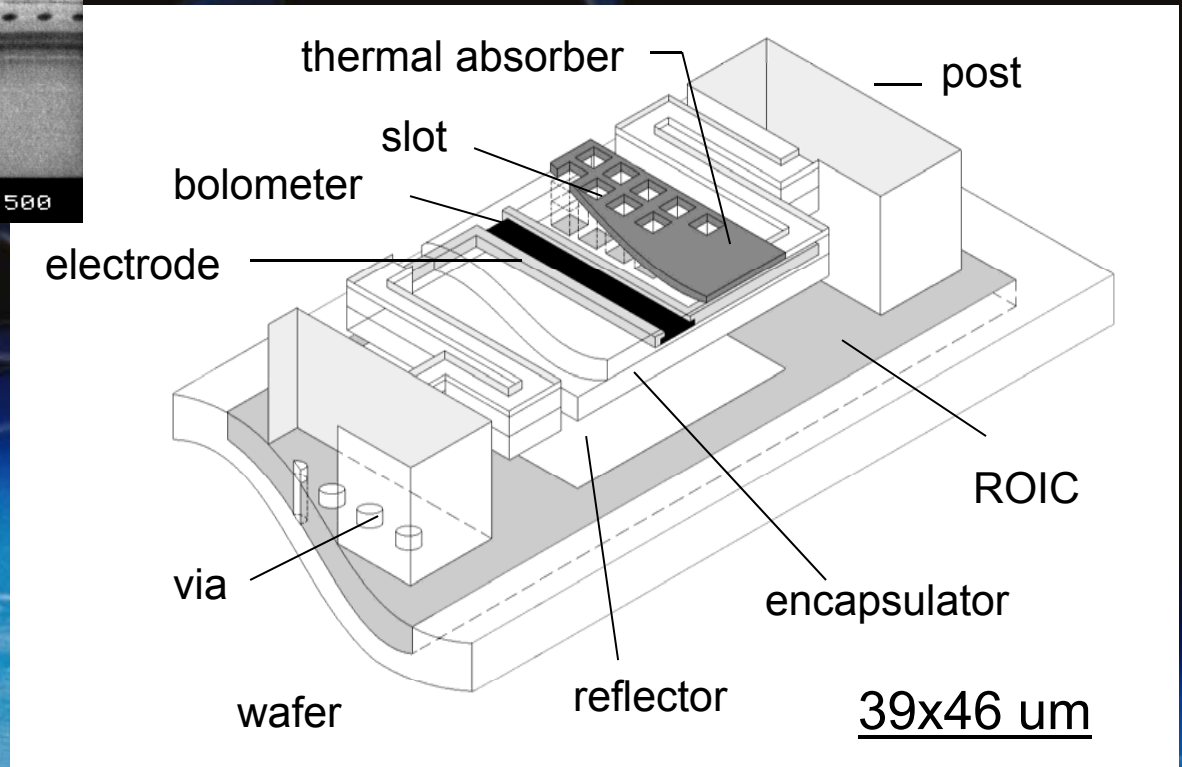
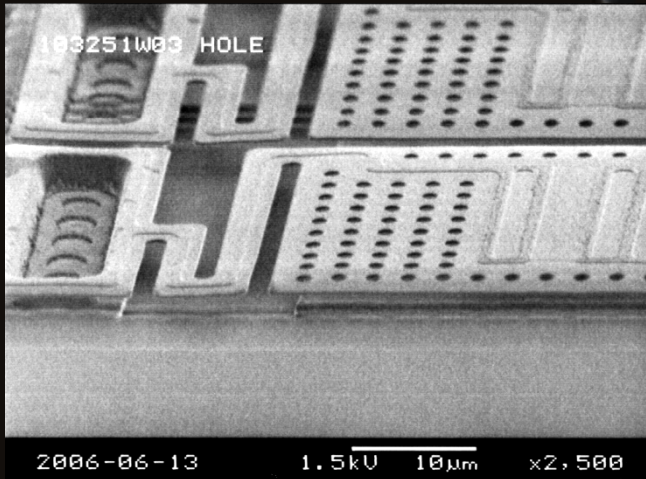
30 ms integration, F/1 optics, 90% transmission

(Kruse, SPIE 2552, 556)

- ▷ Need for low G , lower limit of effective $G \sim 10^{-8}$ W/K
- ▷ C must be kept to small values to ensure adequate speed (C/G)
- ▷ Designs with reduced C and G are structurally challenging

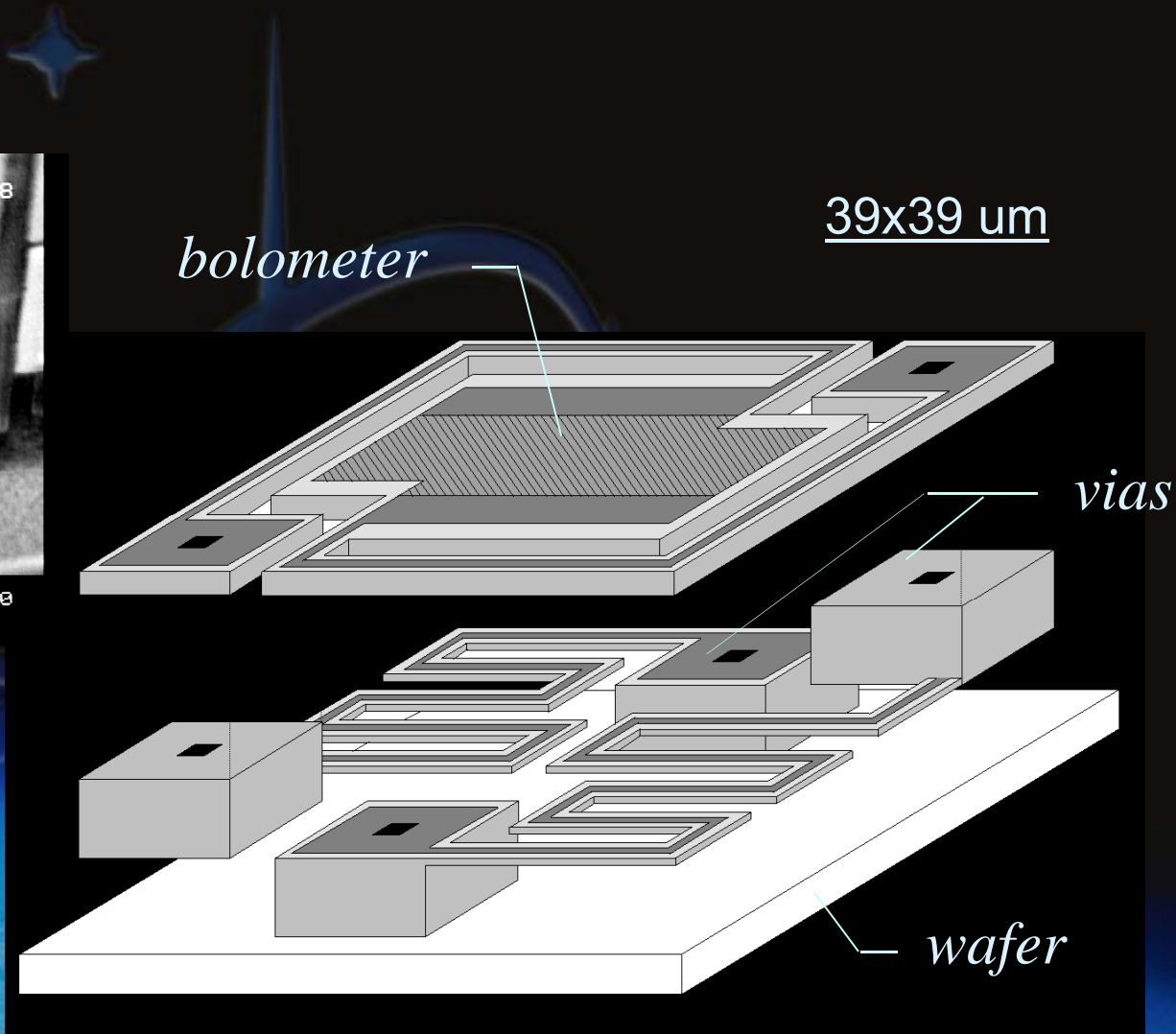
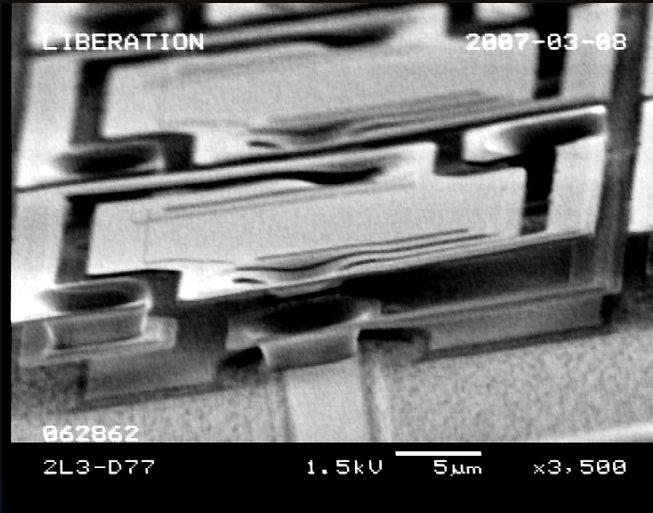


Single stage design





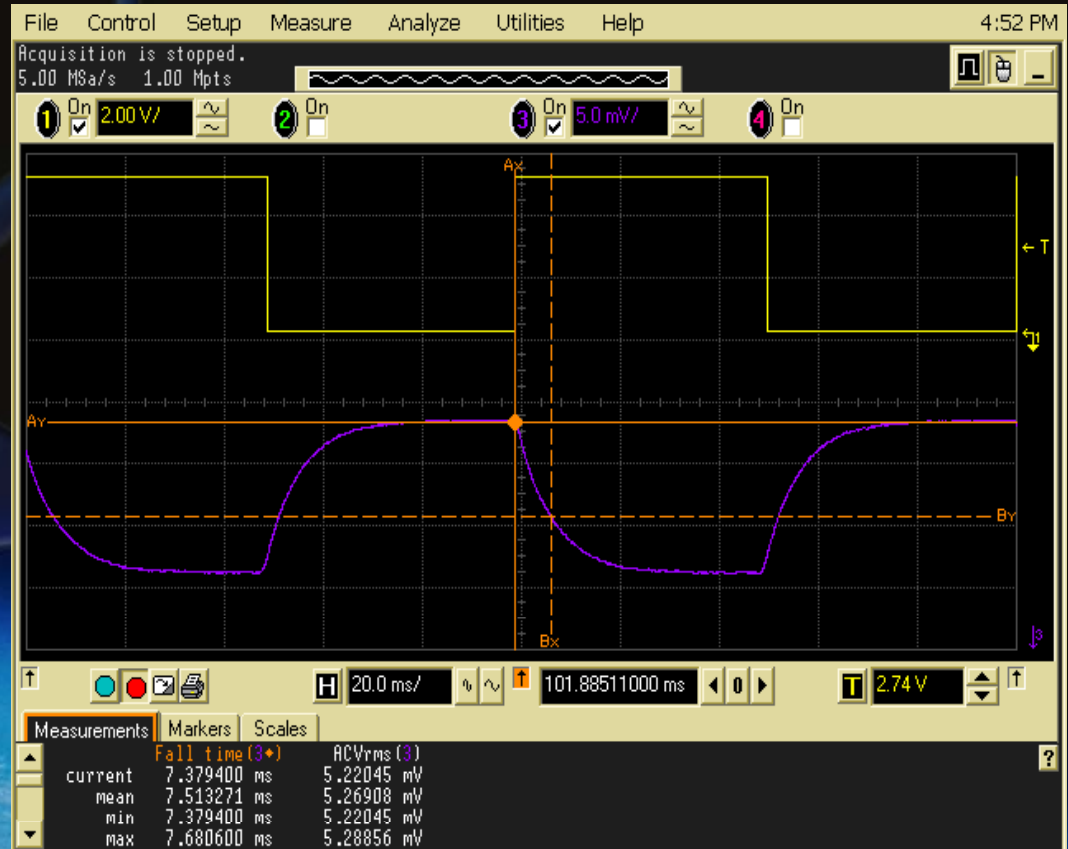
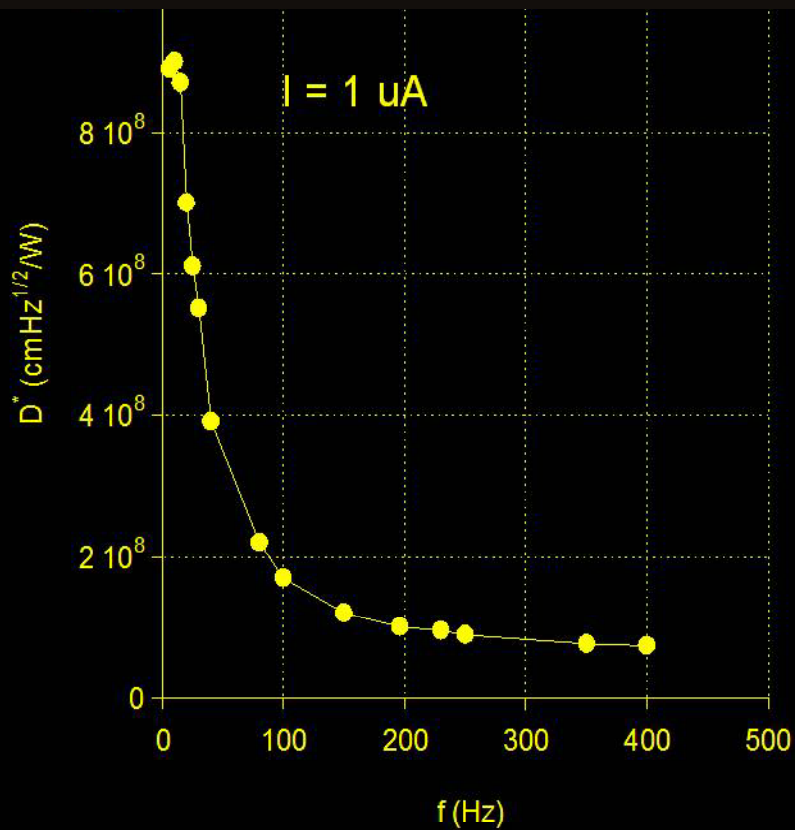
Double stage design



- Maintain D^* with an increased fill factor
- Avoid along track blurring



Typical characteristics





Pixel signal chain

- *0.5 μm CMOS ROIC*
- *Voltage to frequency conversion*
- *Integrate while read operation with dual registers*
- *Single supply 5V operation*

2.5 V +/- 0.7 V

Pixel
bridge

I1T

Bypass

Swap

4x stage

Reset

Switched
capacitor
integrator

3.5 V

Compa-
rator

Counter

I2T

1.5 V

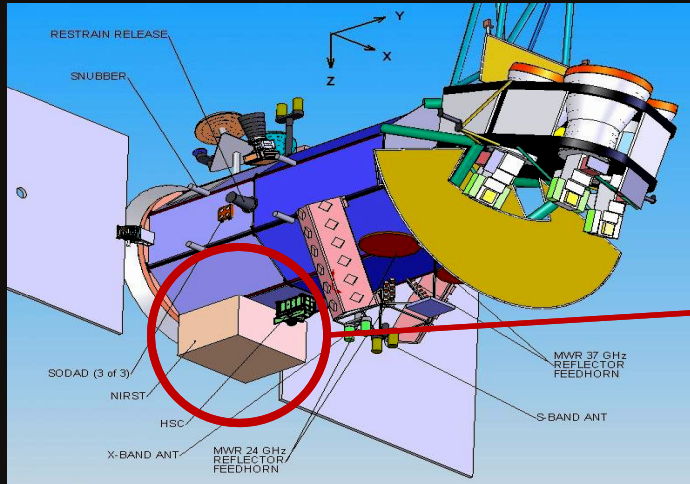


ROIC

Format	512 x 3
Line spacing	2.88 mm
Microbolometer pitch	39 μm
Supplied voltage	5 V +/- 10% all operating conditions
Power consumption	0.7 W / line
Integration / readout	parallel (dual register)
Integration time	10 - 140 ms
Master clock speed	0.8 - 2 MHz
Signal chain response	1.79 μV / lsb (3 mV input, 0.8 MHz, 140 ms integration time)



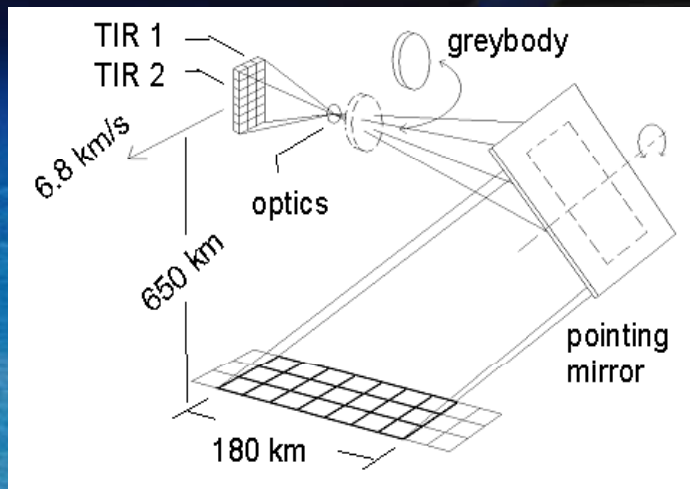
MWIR-LWIR sensing



Linear arrays of 512x3 microbolometers developed for SAC-D / NIRST

NIRST

- *is a pushbroom scan radiometer with boresight pointing and in-field spectral separation*
- *evolves on sun synchronous orbit @ 657 km alt, ascending node @ 6 pm*
- *retrieves temperatures of wildfire, volcanic activities, and sea surface*





MWIR sensor

Format / pitch	512 x 3 / 39 μm
Number of active lines	1
Spectral band 1	3.4 - 4.2 μm
Out-of-band transmission	< 4%
Operating temperature	283 - 291 K
Non-operating temperature	233 - 313 K
Time constant	12 ms
Responsivity	2×10^4 V/W
Detectivity	10^9 cm.Hz ^{1/2} /W
NETD (f/1 optics, 39 μm pixel, 4 μA bias, 50 ms integration, 0.6 transmission)	2.3 K @ 400 K scene
Radiometric temperature range	300 - 700 K
Saturation temperature	> 600 K



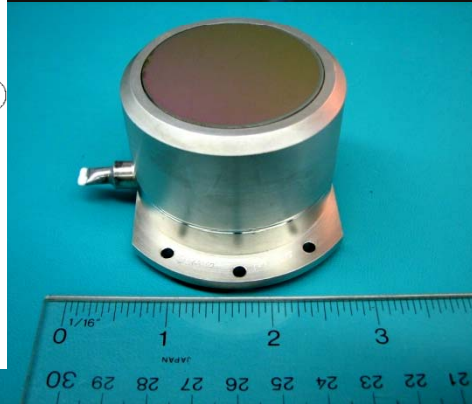
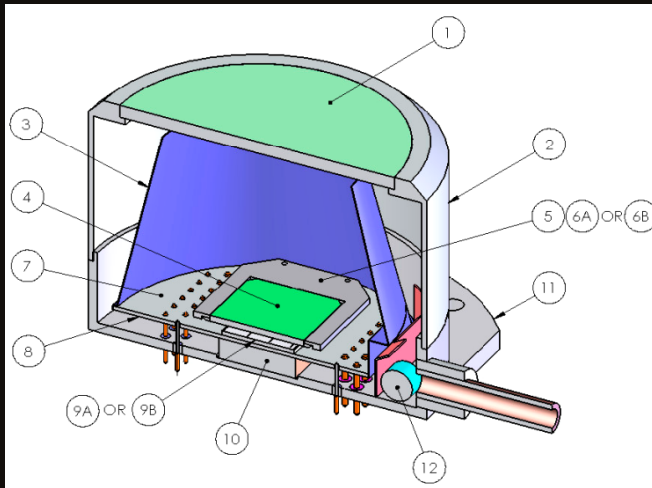
LWIR sensor



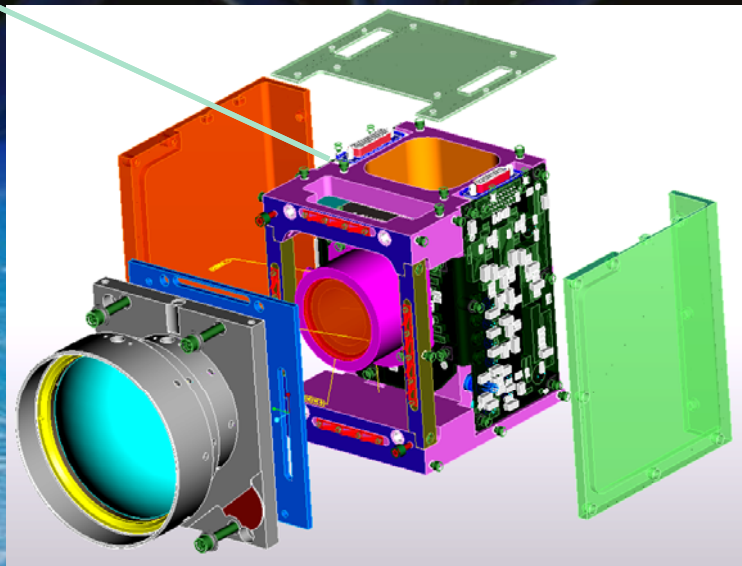
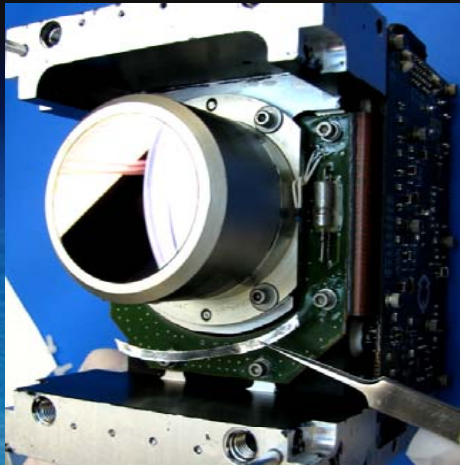
Format / pitch	512 x 3 / 39 μm
Number of active lines	2
Spectral band 2	10.4 - 11.3 μm
Spectral band 3	11.4 - 12.3 μm
Operating temperature	283 - 291 K
Non-operating temperature	233 - 313 K
Out-of-band transmission	< 1%
Time constant	12 ms
Responsivity	$4.2 - 5.8 \times 10^4$ V/W
Detectivity	$2.1 - 2.9 \times 10^9$ cm.Hz ^{1/2} /W
NETD (f/1 optics, 39 μm pixel, optimum current bias, 50 ms integration, 0.6 transmission, readout noise included)	< 0.4 K @ 300 K scene, band 2 < 0.5 K @ 300 K scene, band 3
Radiometric temperature range	200 - 400 K
Saturation temperature	> 600 K



Radiometric sensor

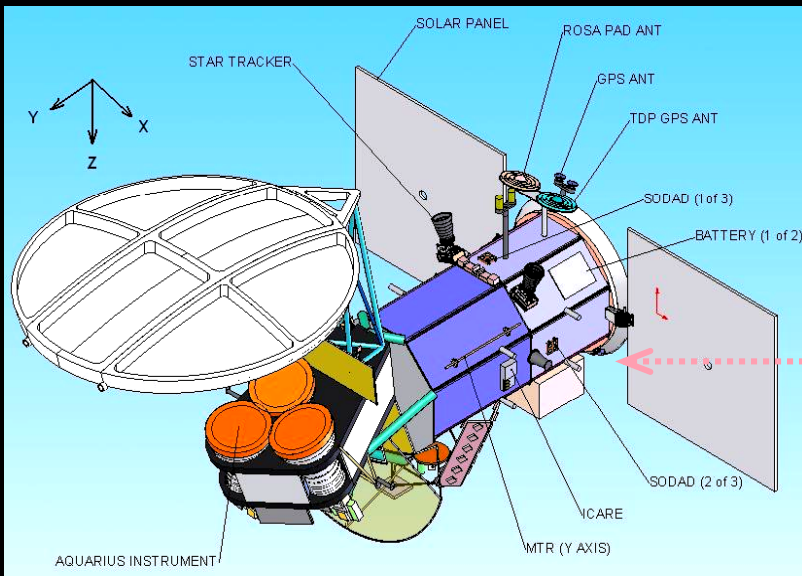
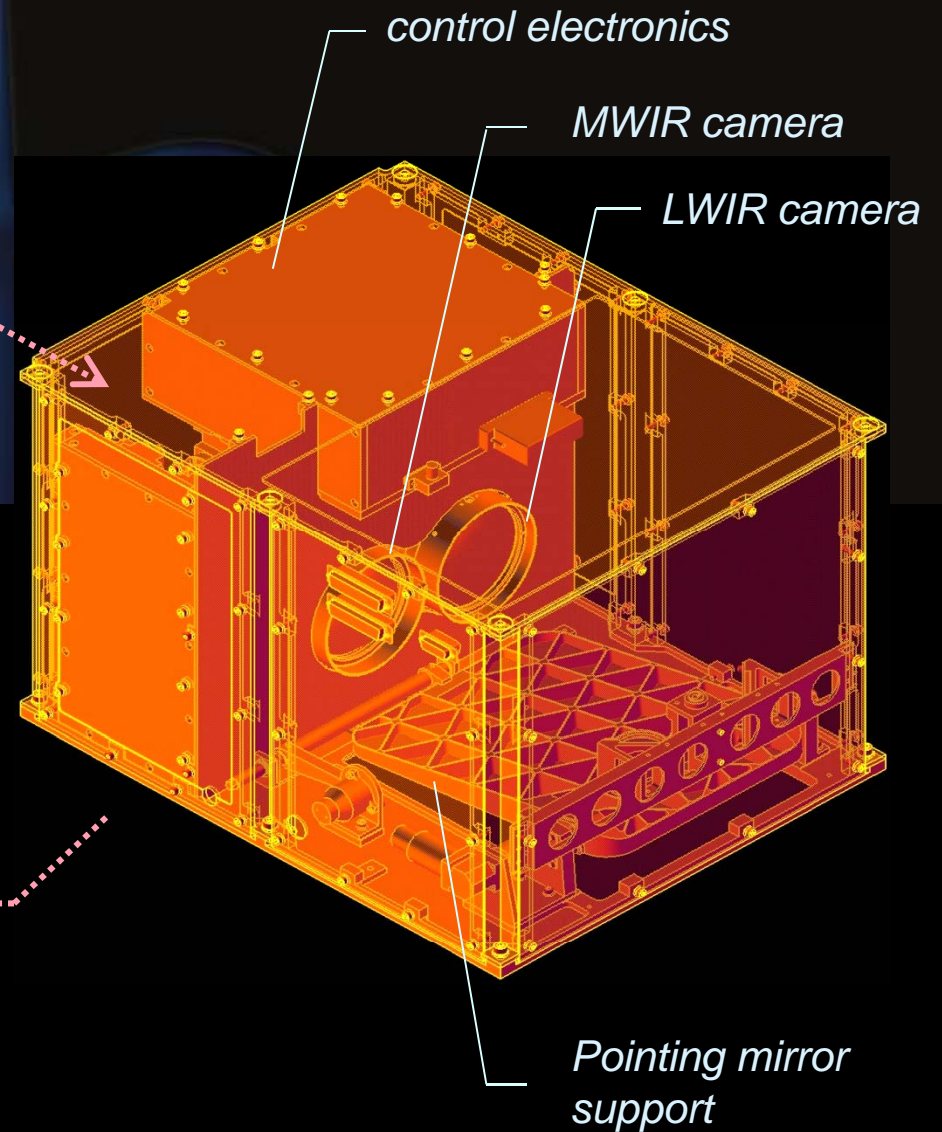
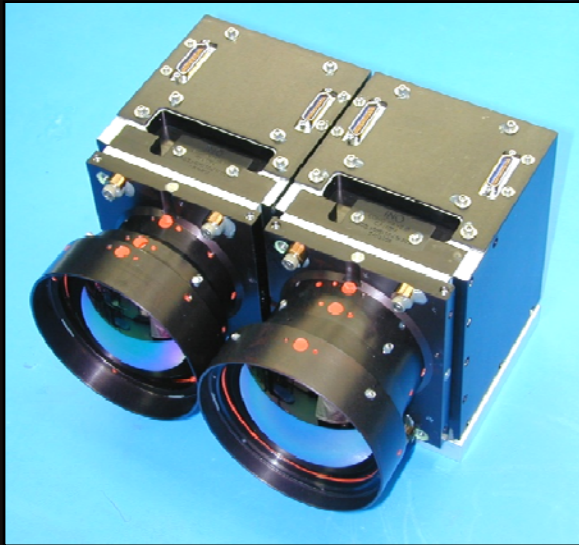


(1) IR window; (2) package cover; (3) cold shield; (4) spectral separation filter; (5) filter assembly; (8) routing circuit wafer; (9) 512x3 microbolometers with ROIC; (10) thermoelectric cooler; (11) assembly header with vacuum tube; and (12) vacuum stabilizing getter.





Instrument details





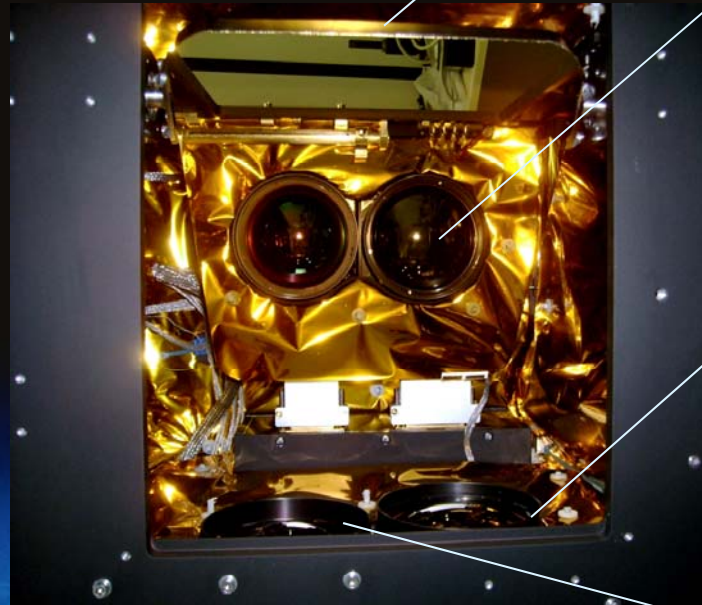
Instrument details

Pointing mirror

Reflection of cameras pointing to Earth surface

LWIR camera

MWIR camera

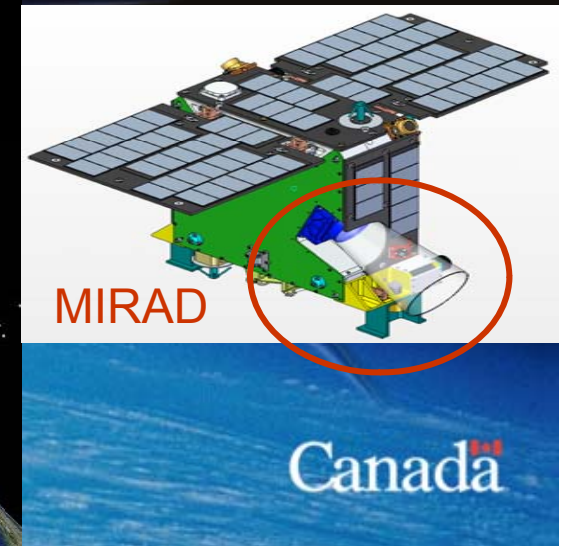
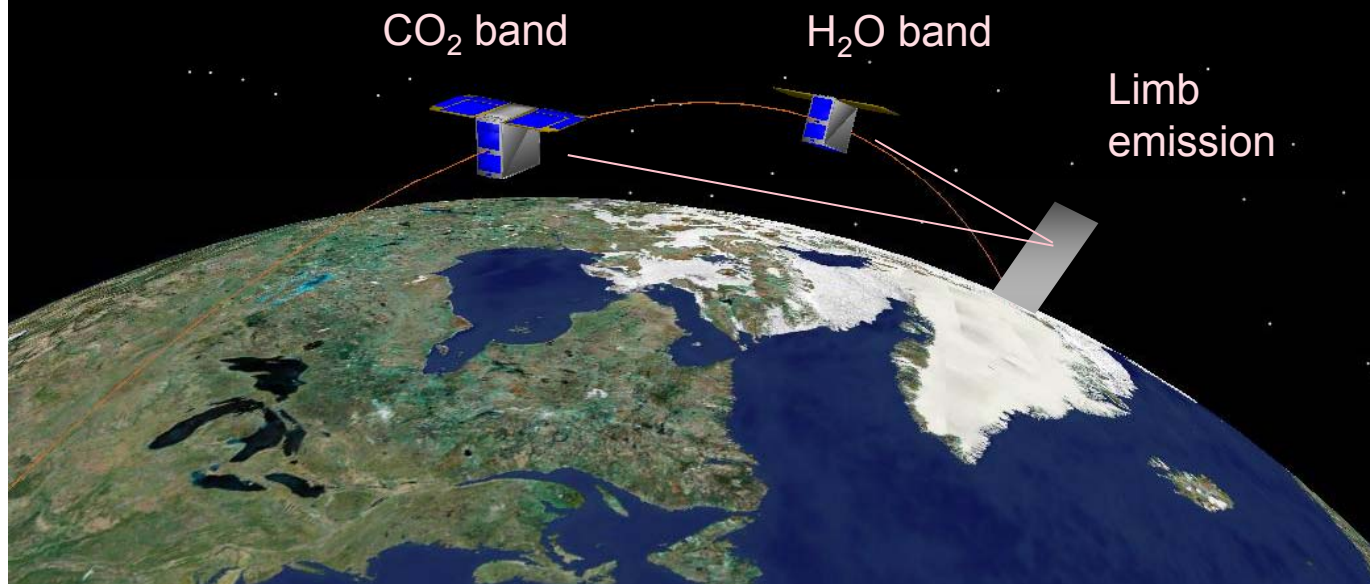




Far IR sensing

JC2Sat-FF

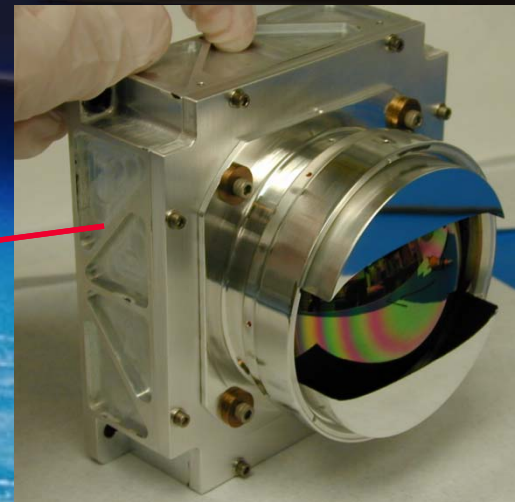
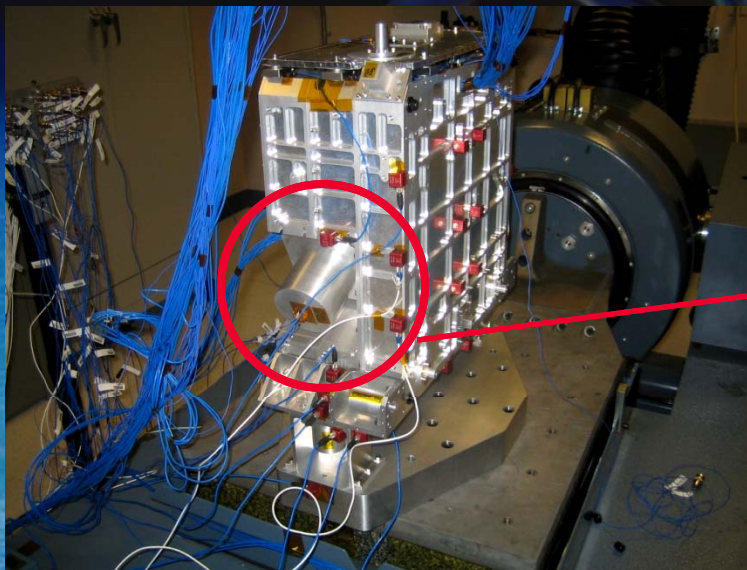
- Consist of two miniature far IR radiometers distributed on two FF nanosats
- Coregister limb radiances in the CO₂ 15 um band (sat 1) and H₂O 25 um band (sat 2)
- Generate a database of limb radiance profiles with global, seasonal, day/night time variations
- Provide characteristics of navigational horizon and atmospheric OR in the far IR





Radiometers

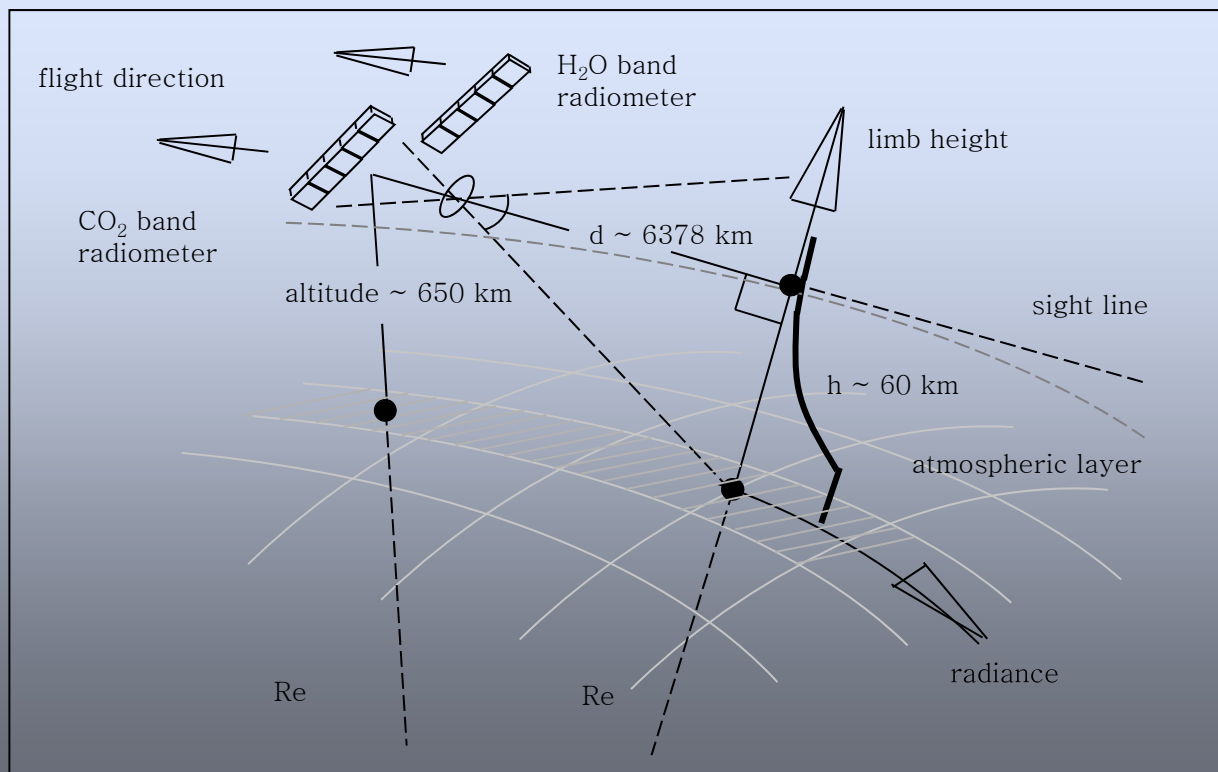
- ▶ Miniaturized to fit in nanosats
- ▶ Differ in operating spectral band, one band per nanosat
 - *FIR1: 14-16 μm*
 - *FIR2: 24.3-26.1 μm*
- ▶ Coregister the radiance profiles of the Earth's limb
- ▶ Enable direct FIR measurement using microbolometers





Operation

- Perform measurements during the formation keeping phase (30 days)
- Acquire and transmit data to CPU when powered
- Measure in sequences of 14 s with 1 orbital period between sequences (latitude coverage in 28.5 days)
- Downlink data to two ground stations (~ 15 min / station / sat / day)





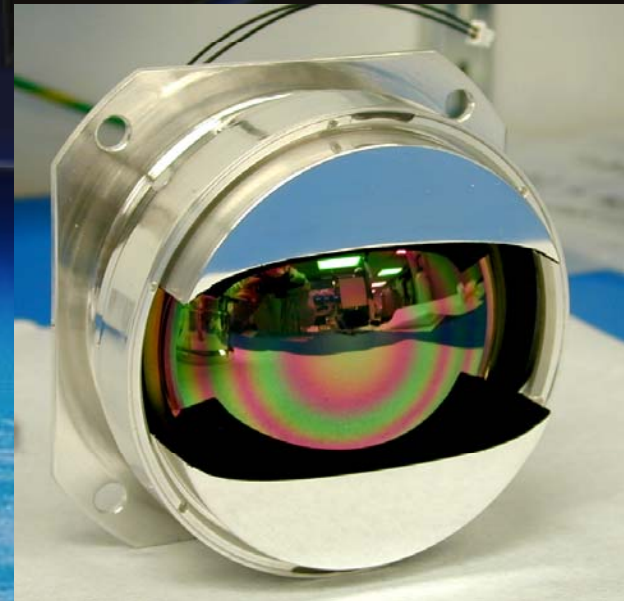
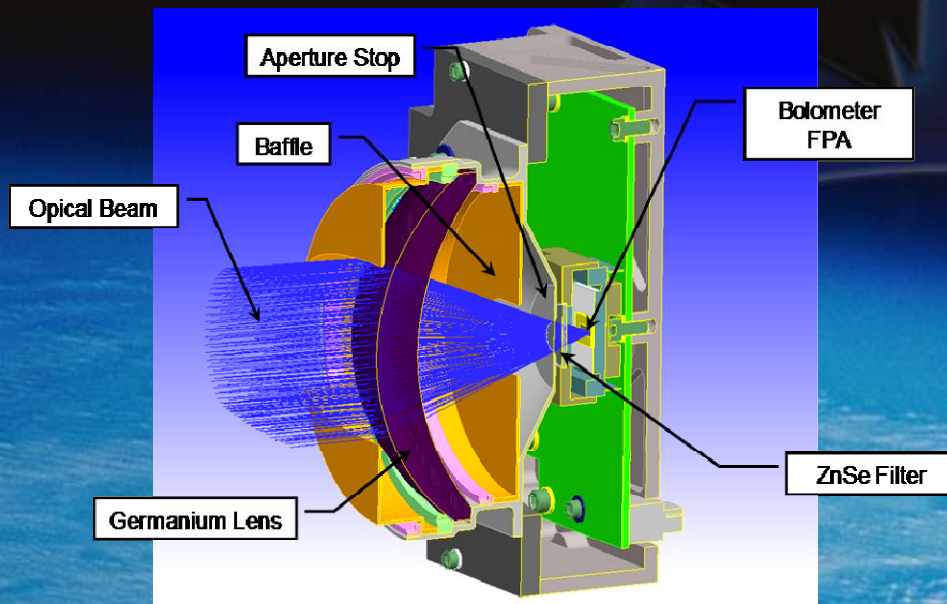
Budgets

	<i>Spacecraft</i>	<i>Radiometer (allocation)</i>	<i>Radiometer (design)</i>
<i>Power (W)</i>	14	1	0.39 (average) 0.58 (max)
<i>Mass (kg)</i>	20	0.50	0.48
<i>Volume (mm³)</i>	350x350x150	105x105x100	105x105x100



Optics

- Largest fraction of the volume and mass budgets
- Optics must be small but provides enough throughput for the sensor
- Optical materials must be compatible with the design of small optics
 - *Ge allows for diffraction limited single lens design in band 1*
 - *CdTe has an acceptable transmittance in band 2 and can be diamond turned into aspherical surfaces*





Optics



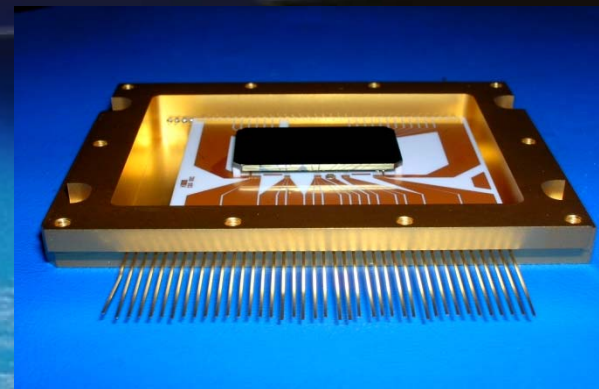
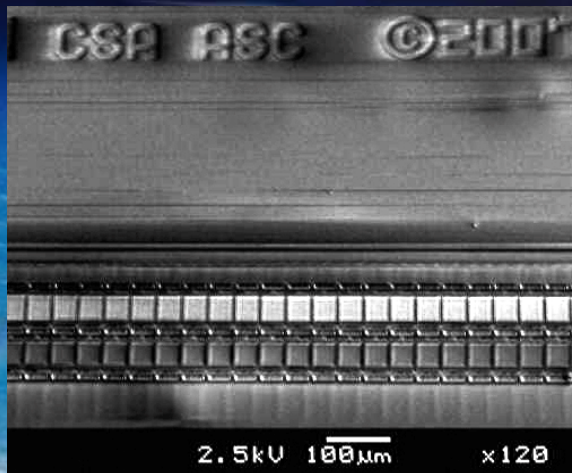
	<i>Band 1</i>	<i>Band 2</i>
<i>Waveband</i>	Corrected from 14 um to 16 um	Corrected from 24.2 um to 26.2 um
<i>IFOV (per pixel)</i>	0.98 mrad	0.98 mrad
<i>F/# (Working)</i>	F/1.47	F/1.3
<i>Detector</i>	1X256 pixels, 1D=13.312 mm (50 um/pl, 2 um gap)	1X256 pixels, 1D=13.312 mm (50 um/pl, 2 um gap)
<i>EFL</i>	51.02 mm	51.02 mm
<i>Entrance Pupil Diameter</i>	34.013 mm	39.246 mm
<i>FFOV in 1D</i>	14.87 degrees	14.87 degrees
<i>Elements</i>	1 aspherical lens (in Ge)	2 aspherical lenses (both in CdTe)
<i>Optics mass</i>	171.0 g	82 g
<i>Dimension</i>	71 mm diameter 52.1 mm length	67 mm diameter 84.4 mm length
<i>Total transmittance</i>	39%	34%
<i>Athermal range</i>	10 ⁰ C to 28 ⁰ C	10 ⁰ C to 28 ⁰ C



Sensors

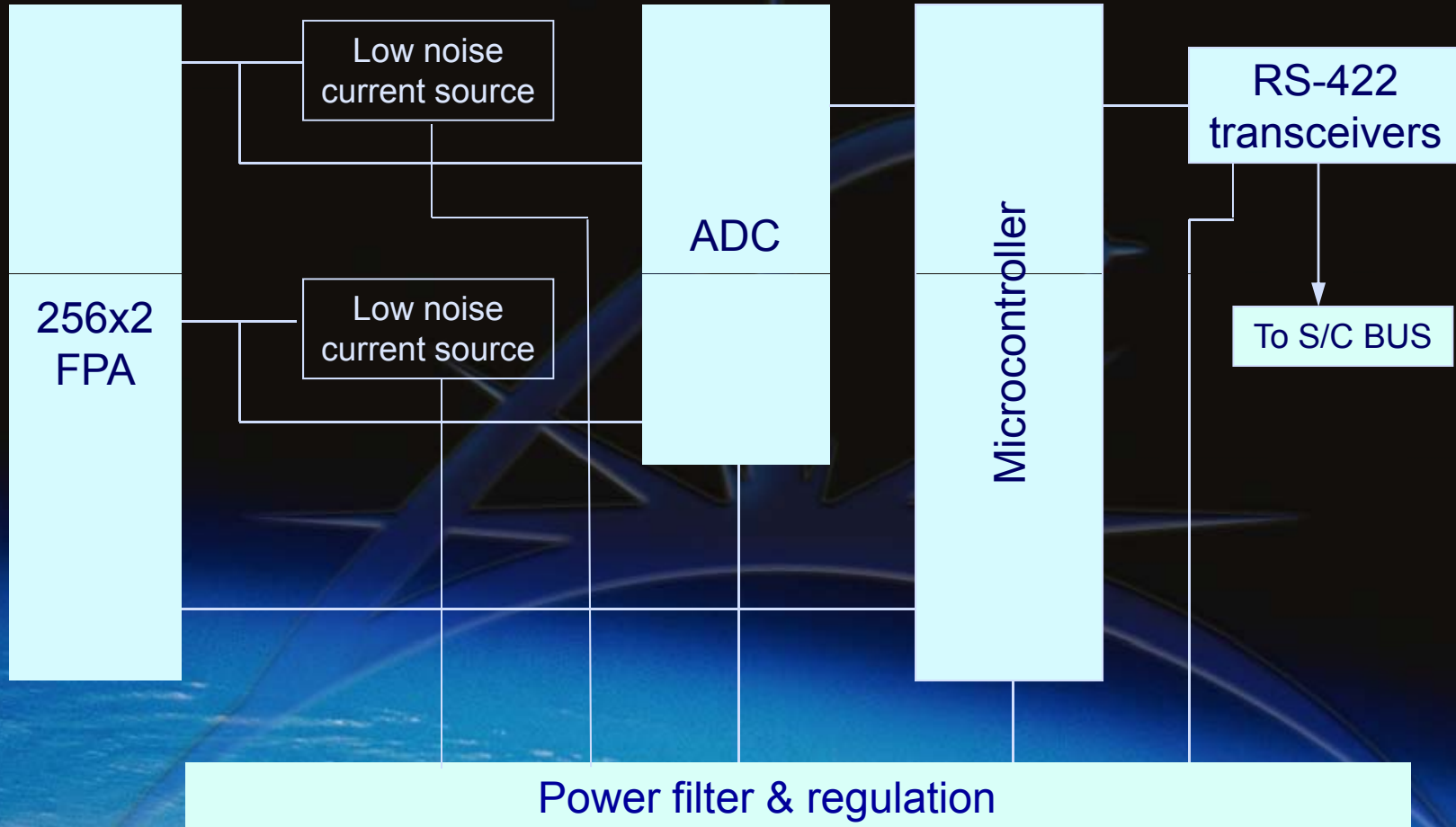
- $IFOV \sim 1$ mrad; resolution: $d \tan IFOV \sim 2.8$ km
- No. of pixels needed to capture the limb: $2h / 2.8$ km ~ 43 pixels
- Equivalence of ± 5 deg error ~ 175 pixels
- A minimum of 218 pixels is required

*Format selected: linear array of 256x2 pixels
One row of active pixels, one row of reference pixels*



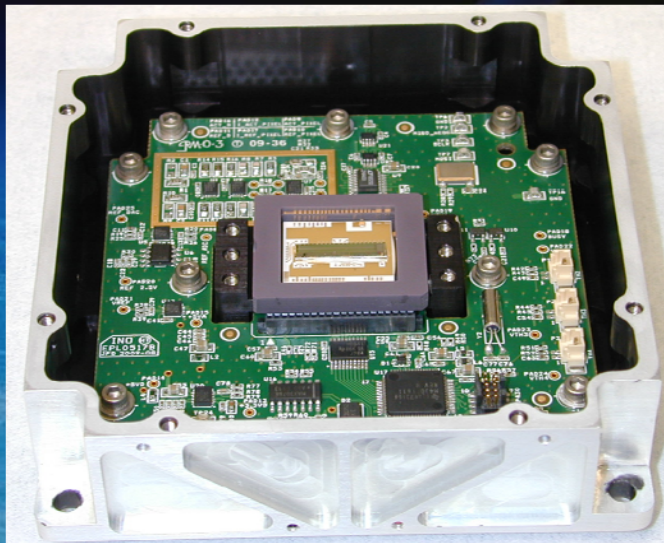
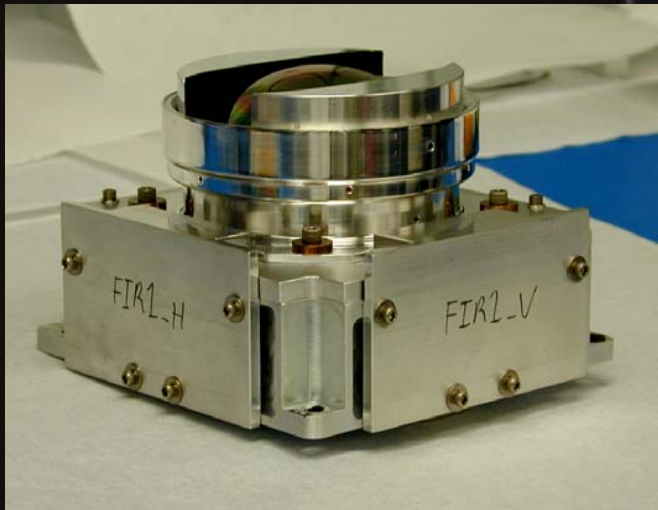


Electronics





Electronics

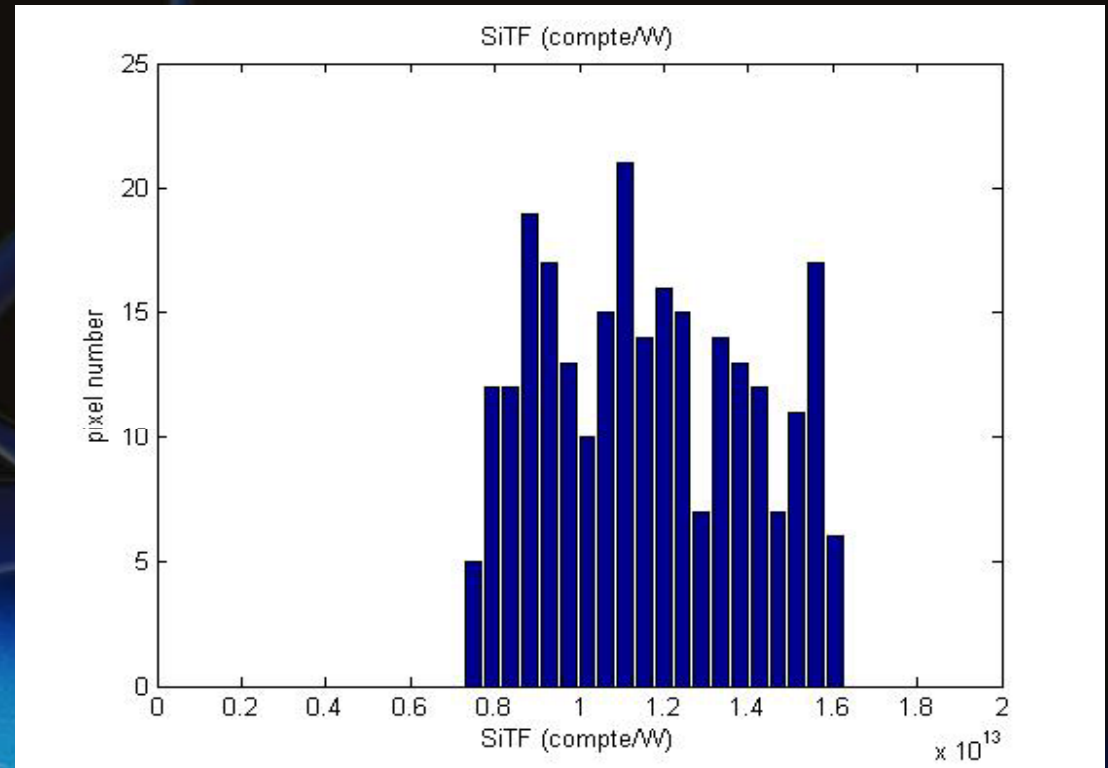
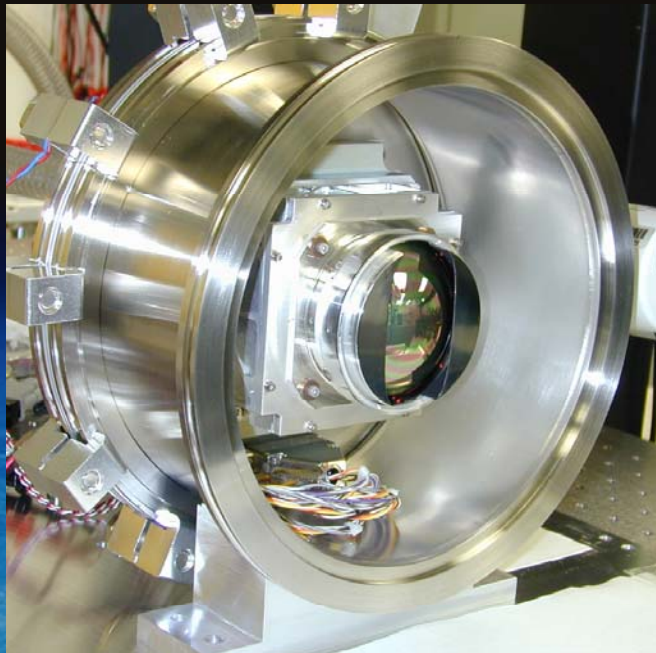


- Remove offset due to die temperature drifts
- Monitor average sensor resistance and adjust supplied current
- Monitor temperatures of lens assembly, sensor, and microcontroller for data interpretation
- Filter and regulate supplied power
- Control process of data acquisition and transmission



Tests

T cycling: -30 to 70 C
Vibration: 14.1 grms





Summary

- Technology development at the CSA has focused on linear arrays of resistive microbolometers
- Details on the developed sensors were presented
- The main challenge - monolithic ROIC - was successfully addressed
- Suitability of microbolometers for use in different infrared bands was confirmed
- Flight sensors were delivered to SAC-D Aquarius mission (MWIR / TIR) and JC2Sat mission (FIR)