

# **Development and Validation of MEMS Sun Sensor based on APS detector**

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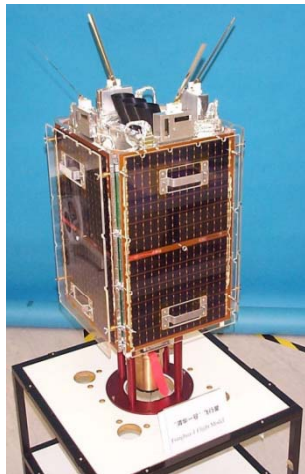
**Xin Fei, You Zheng**

Tsinghua University, Beijing, PRC

2010.8.14

# Microsatellite Research in Tsinghua University

**Successful Launched**   **Manufactured**   **Waiting for Launch**



**Tsinghua-1**

**(50kg, 2000)**

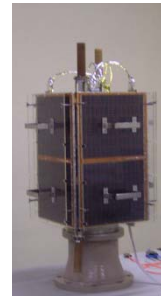


**NS-1**

**(25kg, 2004)**



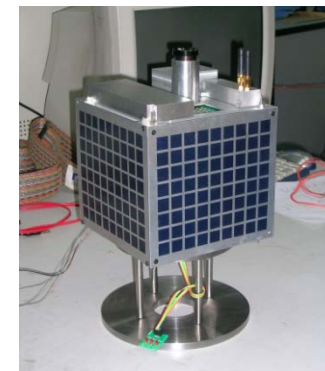
**KT-PS  
(35kg)**



**KT-PS1  
(35kg)**



**NS-2  
(25kg)**



**MEMSat  
(7kg)**

**NS-1 is smallest and high density satellite which has been launched in china untill now( or 3 Axis stable smallest satellite in the world)**

# MEMS device for Space

MEMS based Microgyro, Microaccelerometer, Microsensor, Micromagnetometer, Microthruster, Micro IR sensor, RF device and etc.

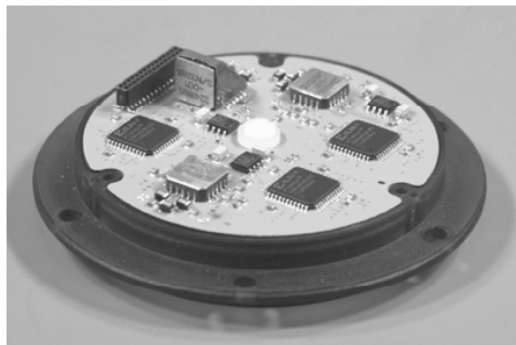
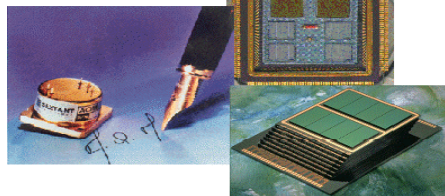
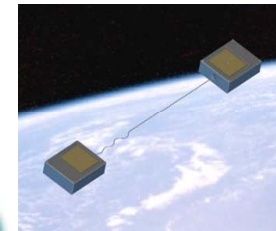
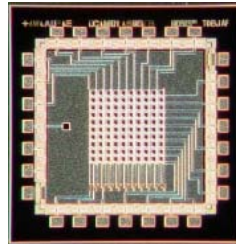
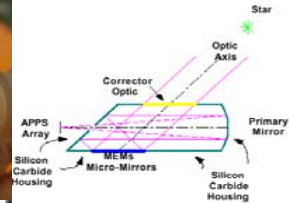
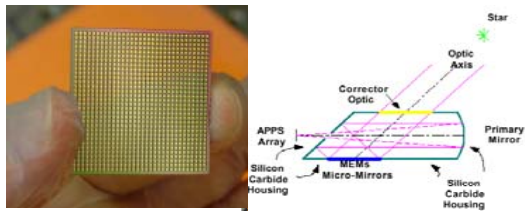
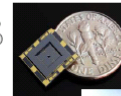


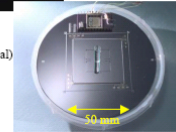
FIGURE 2.7 NMP ST6 ISC MEMS 3-axis gyro assembly. (Source: Charles Stark Draper Laboratory.)

Miniature Gyro

MEMS IMU  
• \$1.2K/0.36 lb (tactical goal)  
• <math>\pm 1m^3</math>  
• <math>0.1 \text{ deg/h}</math>  
• <math>< 1 \text{ W}</math>



Meso IMU  
IMU - \$2.5K/0.5lb (tactical goal)  
• <math>\pm 10m^3</math>  
• <math>0.01 \text{ deg/h}</math>  
• <math>< 1 \text{ W}</math>



Miniature Sun Sensor

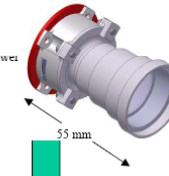


/JPL Publications/2001/NVM-2001/NVM-Tech-Symp ppt

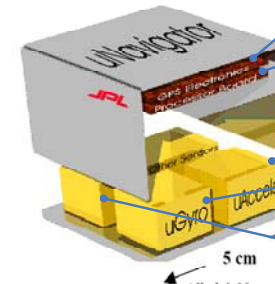
## 美国NASA/JPL实验室 Miniature Micro-Navigator

Miniature Star Tracker and Imager

- Radiation Tolerant
- Scientific Imaging
- Light weight and low power
- Autonomous Operation
- Hazard Avoidance



Miniature Micro-Navigator



- GPS Electronics
- Processor Board
- uAccels
- uGyro
- Other Sensors

# ESA/NASA-JPL/DARPA/ESCC



## Topic:

- MNTs flight history and future in-space demonstration
- MNTs programs, trends and roadmaps
- MNTs in space: specific problems and solutions
- New mission concepts and architectures enabled by MNTs
- Systems, subsystems, equipments and devices using MNTs
- MNTs based functioning demonstrators
- Availability & suitability of Micro-Systems Off The Shelf
- Technology transfer to & from other industrial sectors
- Micro-systems design and modelling
- Materials, technologies, foundries & services, packaging
- Micro- machining, rapid prototyping and fabrication
- Testing methods (inc. radiation), procedures & standards
- MNTs Reliability, evaluation and space qualification
- European Strategy for MNTs
- ESA projects and MNTs

# Contents

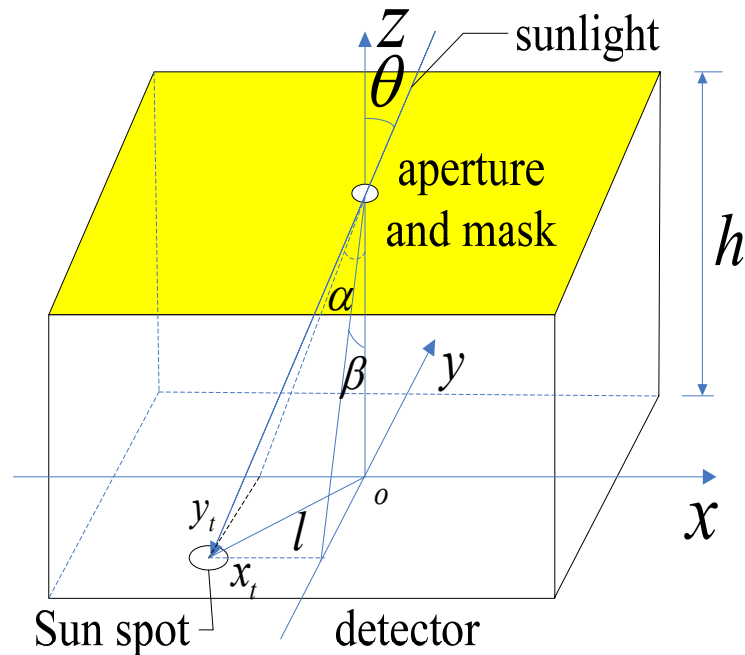
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- Introduction**
- Engineering Model Construction**
- FEIC algorithm**
- Performance Validation on lab and orbit**
- Summary**

# Introduction

## ➤ Principle

According to the location of sunlight imaging on the detector, determine the incidence angle of sunlight.



$$\alpha = \operatorname{tg}^{-1} \left( \frac{x_c}{h} \right) \quad \beta = \operatorname{tg}^{-1} \left( \frac{y_c}{h} \right)$$

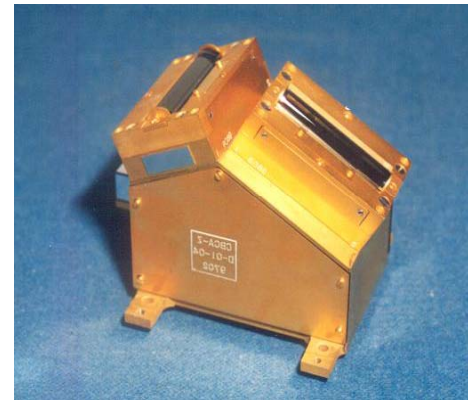
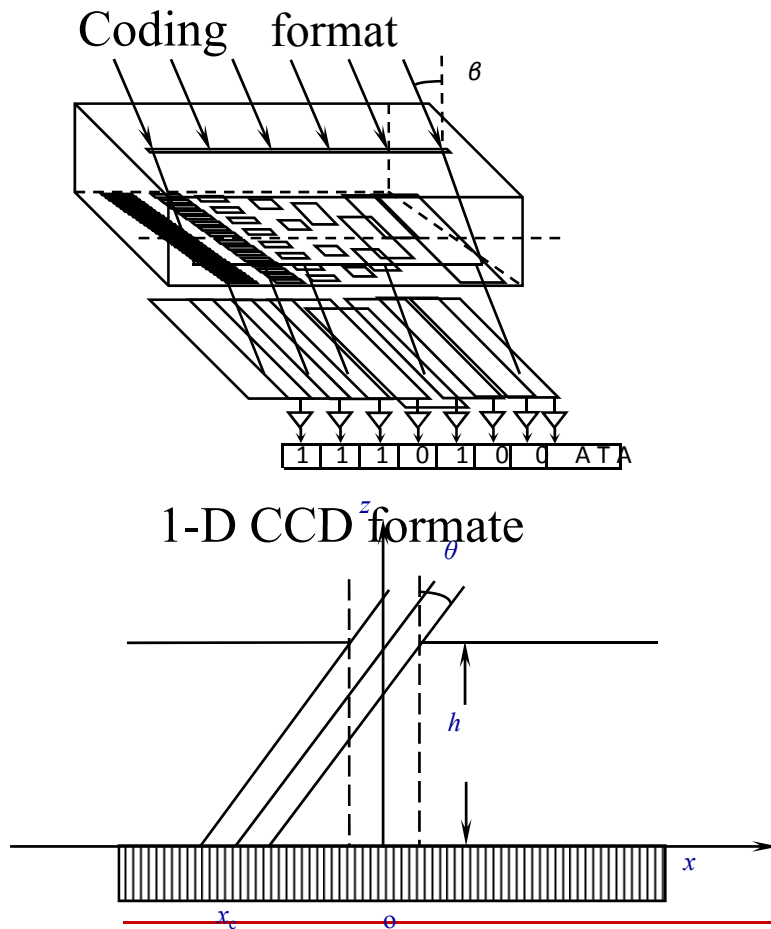
## ➤ Construction

1. Optical mask
2. Detector
3. Electricals

# Line array(1-D) SS principle

➤ Line array principle

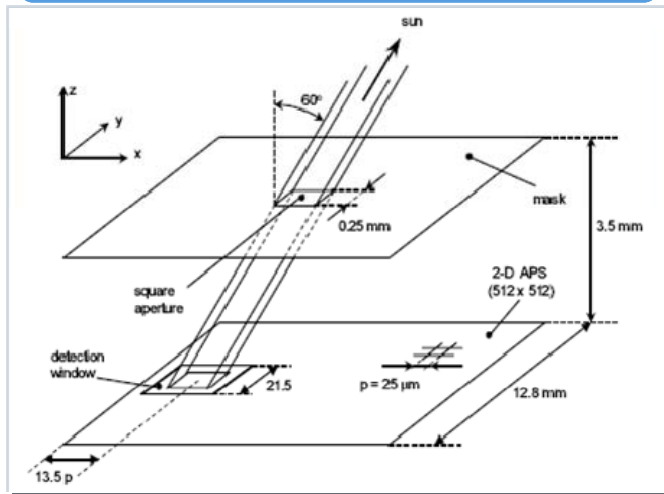
➤ prototype



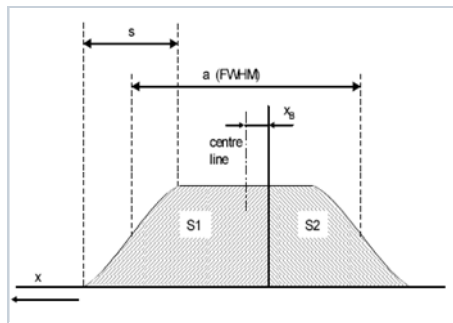
a) DSS2 ( China 502 )

# Area array(2-D) SS principle

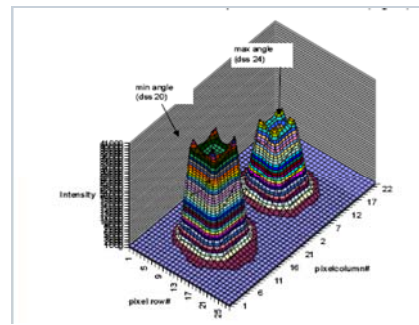
## ➤ 2-D Principle



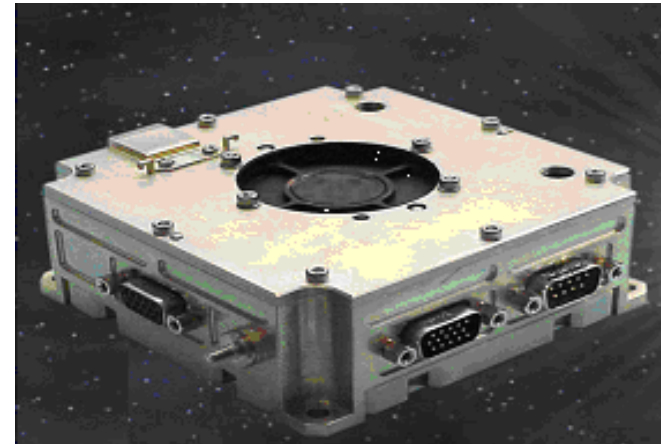
Imaging Principle



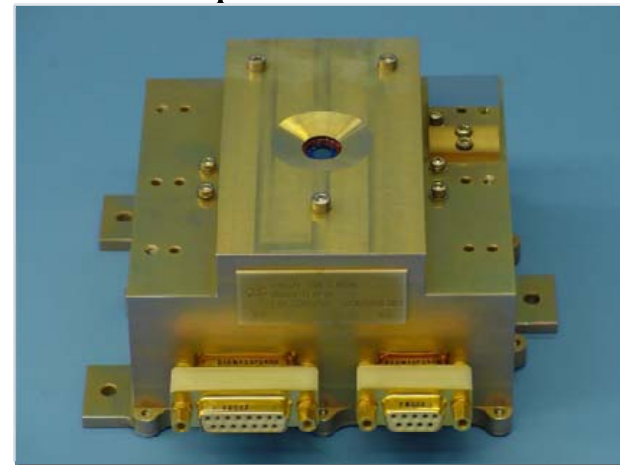
Energy distribution function



## ➤ prototype



Galileo Corp. (Smart Sun Sensor)



Soldern Corp. DSS



# Sun Sensor Trends

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- **2-D array instead of line array;**  
**improve the accuracy and reliability;**
- **APS COMS detector instead of CCD ;**  
**improve the space adaptability and speed;**
- **Multi-apertures instead of single;**  
**improve the accuracy, reliability and robustness.**

# Contents

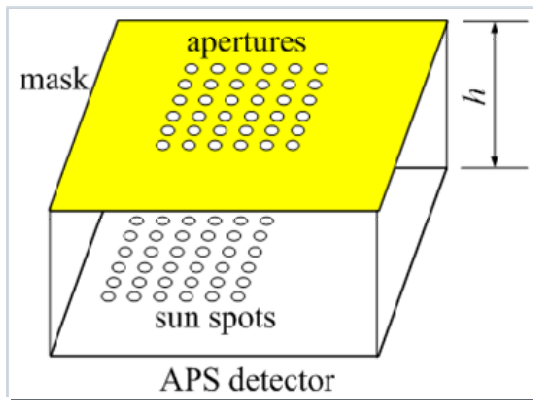
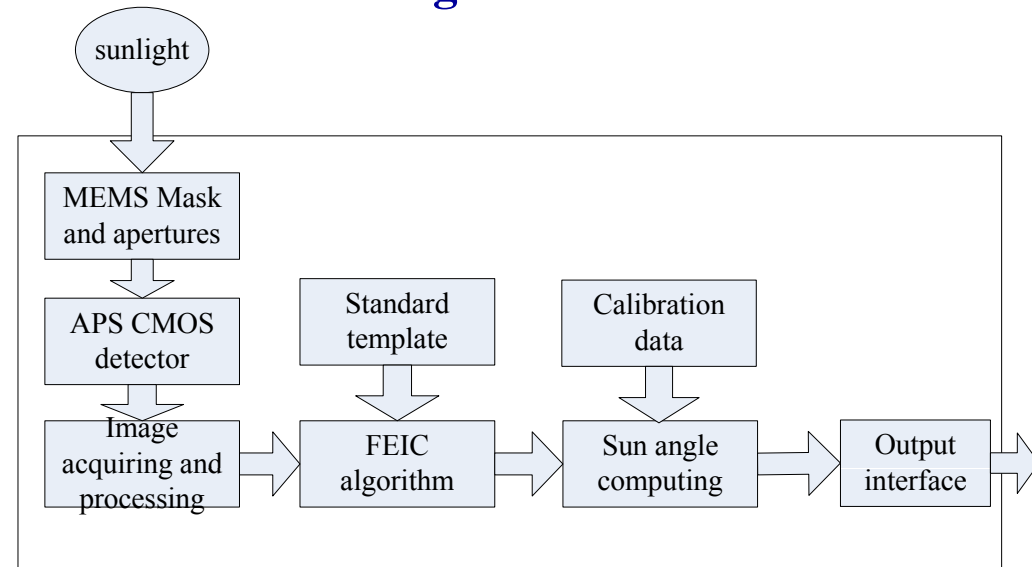
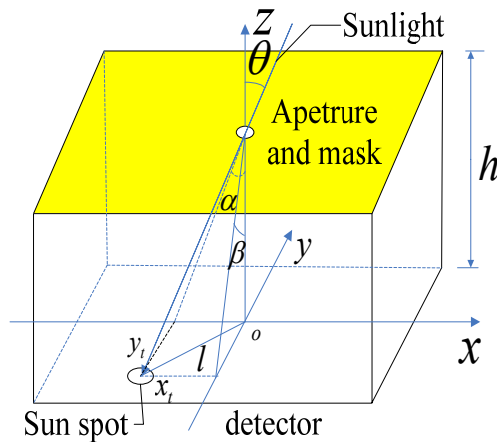
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- **Introduction**
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- **Summary**

# System Principle

## Function Diagram of Sun Sensor

### Construction Principle



### Characters

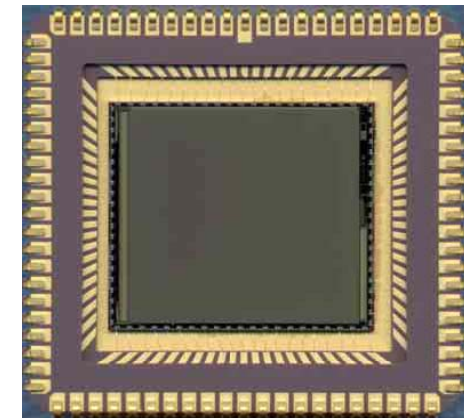
1. MEMS based apertures array mask
2. APS CMOS detectors
3. High accurate FEIC algorithm

# Detector Specification

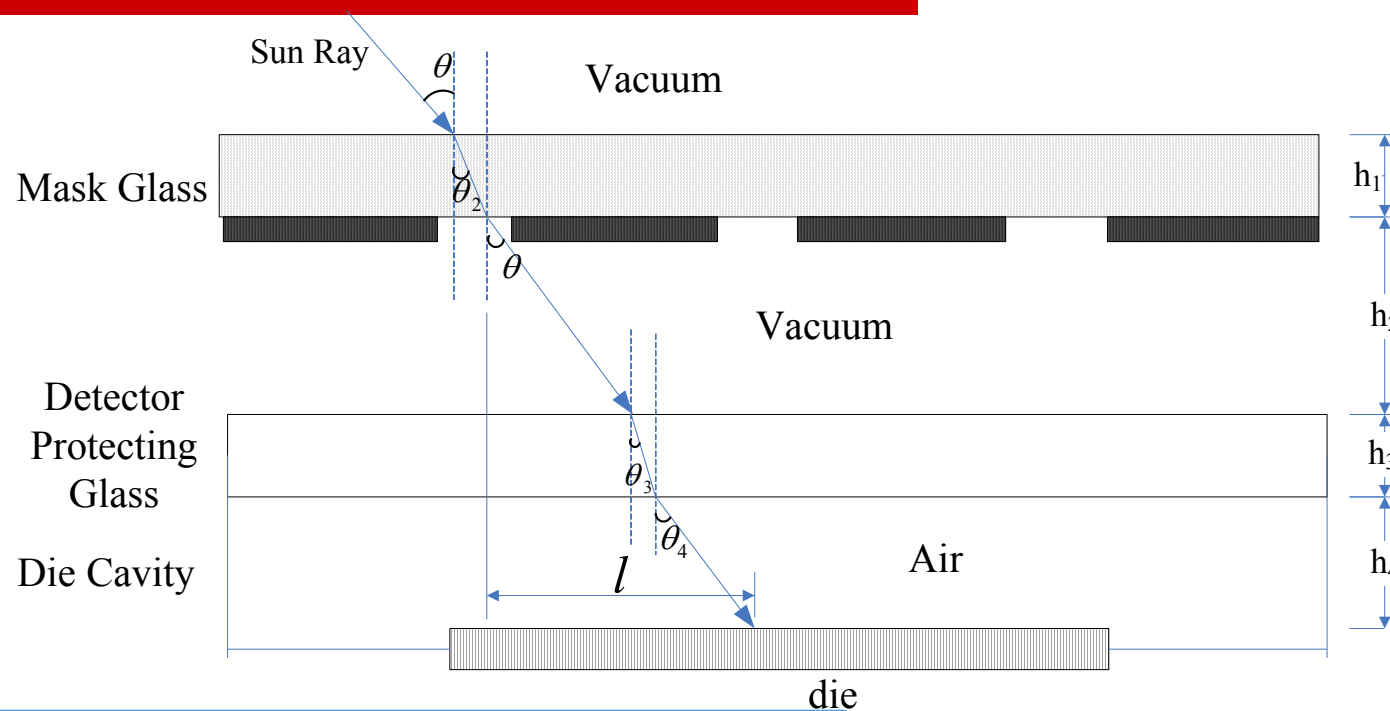
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## APS CMOS detector

Parameter	Specification
Detector technology	CMOS Active Pixels Sensor
Sensitive area format	1024 × 1024
Pixel size	15 $\mu$ m × 15 $\mu$ m
Spectrum range	400nm ~ 1000nm
Quantum Efficiency × FF	>20%
Dose Radiation Tolerance	>230krad(si)



# Sunlight tracking route in MEMS sun sensor



## ➤ Sunlight tracking route

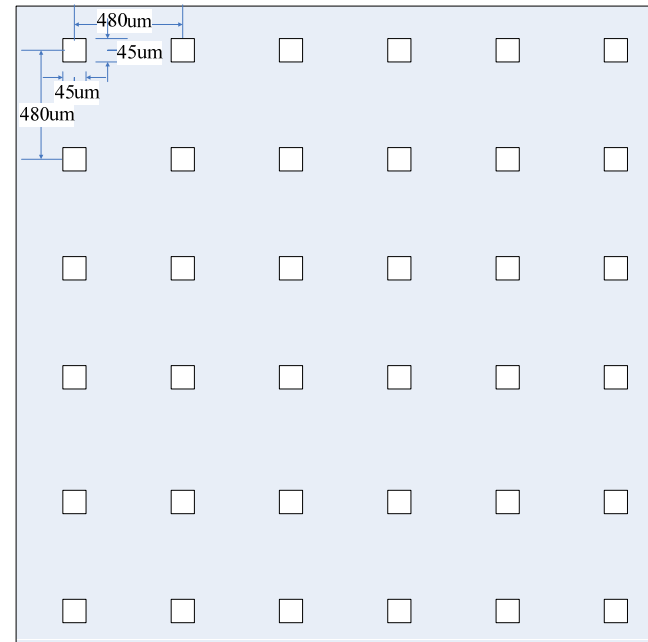
Before Sunlight arriving the detector, it pass through Mask glass, vacuum, Detector protecting glass, Die cavity Air, in turn.

# Multi-apertures Mask

## ➤ parameters

1. Radiation Glass thickness : 1.7mm;
2. Apertures format:  $6 \times 6$ ;
3. Apertures size  $45 \times 45 \mu\text{m}^2 = 3 \times 3$  pixels;
4. Apertures distance:  $480 \mu\text{m} = 32$  pixels;

## ➤ Sketch map



**Parameters is determined by accuracy, and rapid capturing and tracking algorithm.**

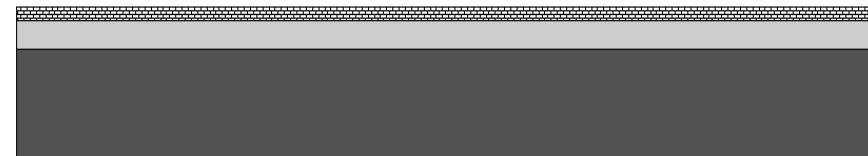
# MEMS Processing for mask



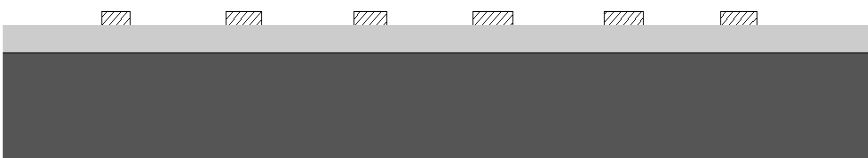
**Glass  
Substrate**



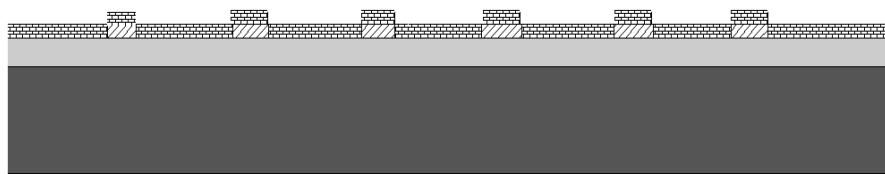
**Sputtering Cr  
film**



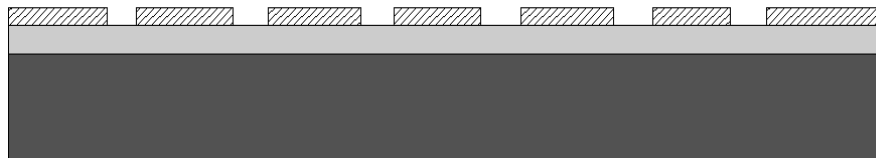
**Coating with  
photoresist**



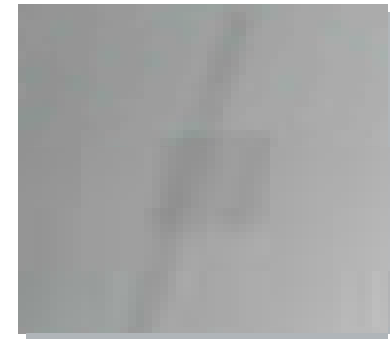
**The photoresist  
patterned  
lithographically**



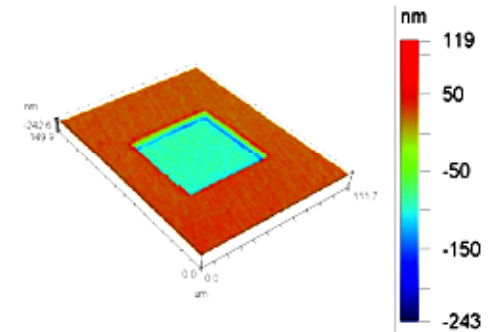
**Sputtering  
Au firstly,  
then Cr film**



**lift-off, forming  
apertures**



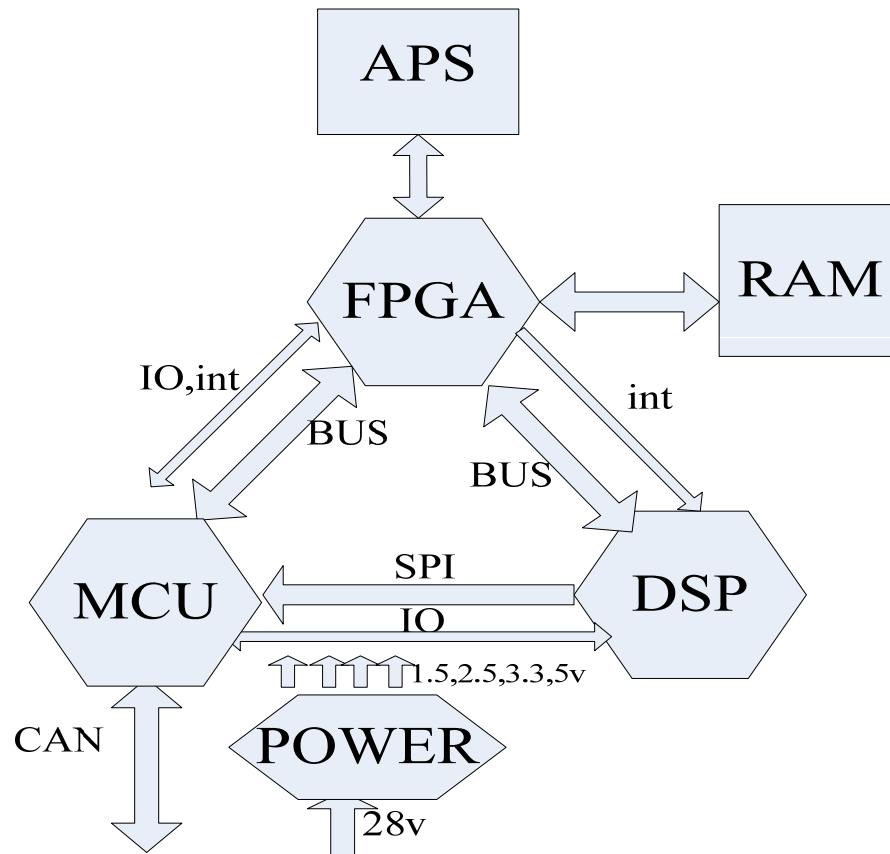
**Picture of Mask with  
6×6 apertures**



**two dimensional surface  
morphology of the single  
aperture**

# Electrics Design

## Electrics Sketch map

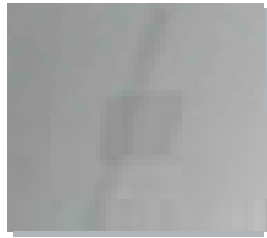


### ➤ Feature

1. Both DSP and MCU may access the image area via FPGA.
2. MCU backups the function of DSP.



# Sun Sensor construction



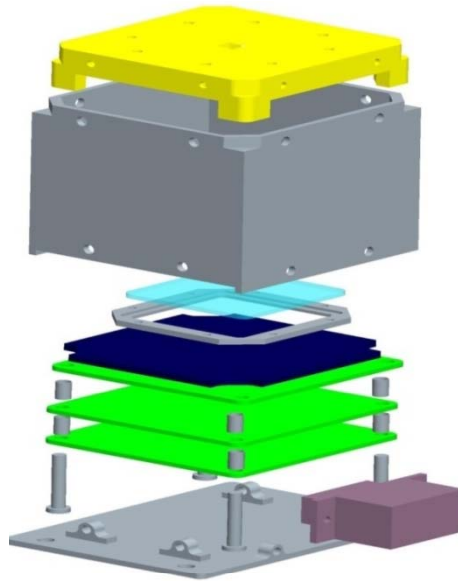
**MEMS** based mask



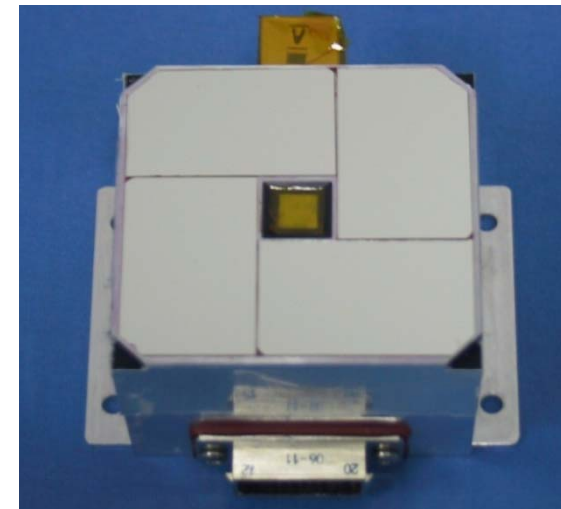
Large array Image sensitivity and readout



High speed Image processing and attitude computing



Layers Installation



sun sensor Production

# Contents

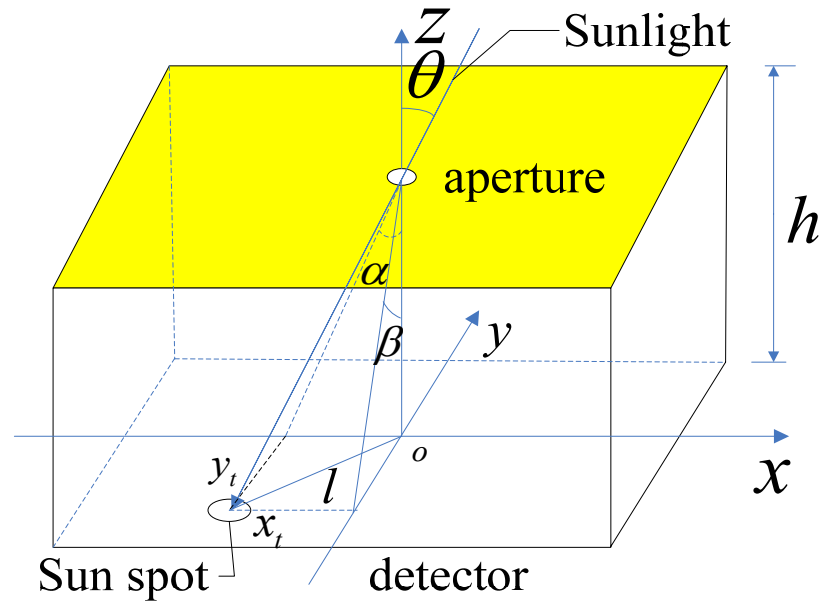
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# Imaging modeling without protected glass

➤ Sketch map

➤ equations



$$\theta = \arctan\left(\frac{l}{h}\right), \alpha = \arctan\left(\frac{x_c}{h}\right), \beta = \arctan\left(\frac{y_c}{h}\right) \quad (1)$$

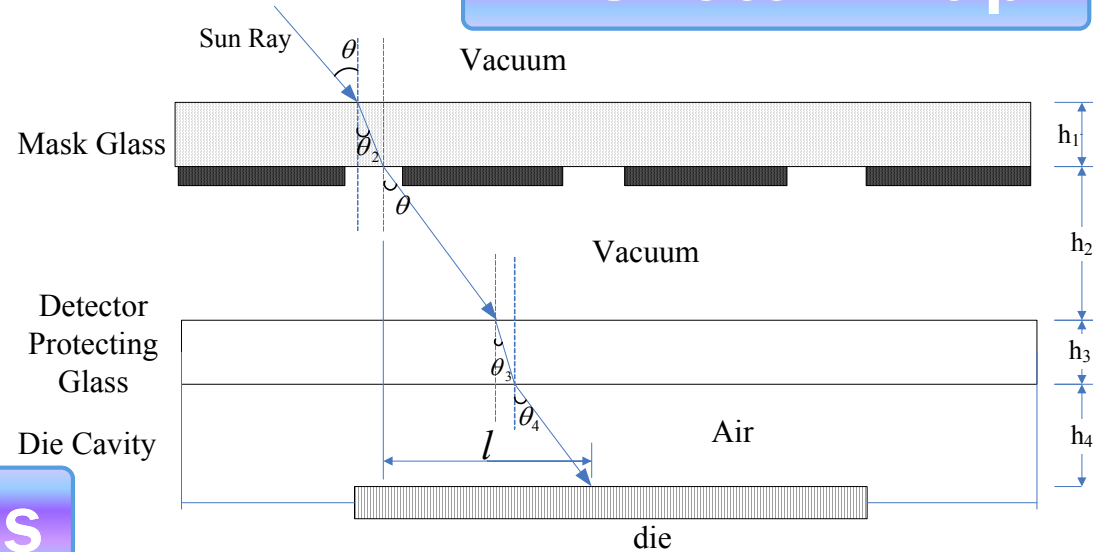
$$l = \sqrt{x_c^2 + y_c^2} \quad (2)$$

$$\tan \theta = \sqrt{(\tan \alpha)^2 + (\tan \beta)^2} \quad (3)$$

Eq. (1) applies to the case without protected glass ;  
Eq. (2),(3) applies to the both case with or without protected glass .

# Imaging modeling with protected glass

➤ Sketch map



➤ equations

$$l = h_2 \tan \theta_2 + h_3 \tan \theta_3 + h_4 \tan \theta_4 \quad \theta_2 = \theta \quad n_3 = \frac{\sin \theta}{\sin \theta_3} \quad n_4 = \frac{\sin \theta}{\sin \theta_4}$$

**Sun spot position  $l$  is unrelated with mask glass thickness, and closely related with detector protected glass thickness.**

# Sun angles computing algorithm with protected glass

## ➤ Function with glass

$$l = h_2 \tan \theta + h_3 \tan \left( \arcsin \left( \frac{\sin \theta}{n_3} \right) \right) + h_4 \tan \left( \arcsin \left( \frac{\sin \theta}{n_4} \right) \right)$$

**It is a Monotone function and no analytic solution.**

$$\tan \theta = f(l) \quad l = \sqrt{(x_c - x_0)^2 + (y_c - y_0)^2}$$

$$\tan \theta = \sqrt{(\tan \alpha)^2 + (\tan \beta)^2} = f \left( \sqrt{(x_c - x_0)^2 + (y_c - y_0)^2} \right)$$

➤ **Key point**

$$\alpha = \arctan \frac{-(x_c - x_0) \tan(\theta)}{\sqrt{(x_c - x_0)^2 + (y_c - y_0)^2}}$$

$$\beta = \arctan \frac{-(y_c - y_0) \tan(\theta)}{\sqrt{(x_c - x_0)^2 + (y_c - y_0)^2}}$$

- 1. The basic idea is through  $(x_c, y_c)$  to solve  $\tan \theta$**
- 2. System accuracy is determined by the accuracy of  $(x_c, y_c)$**

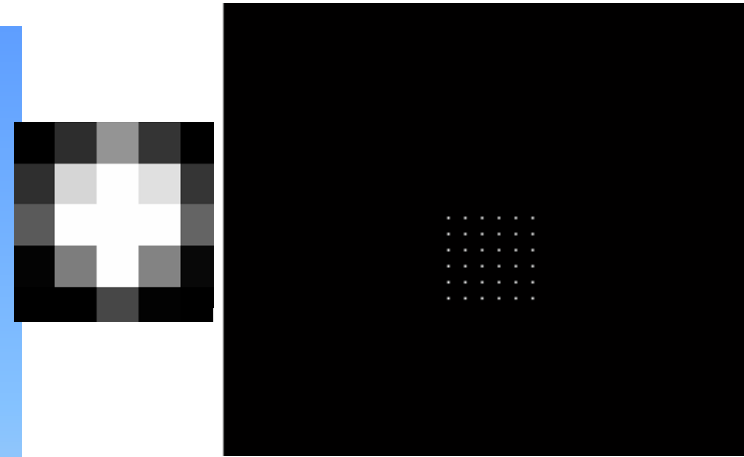
# MultiApertures image Characteristics

## ➤ Features

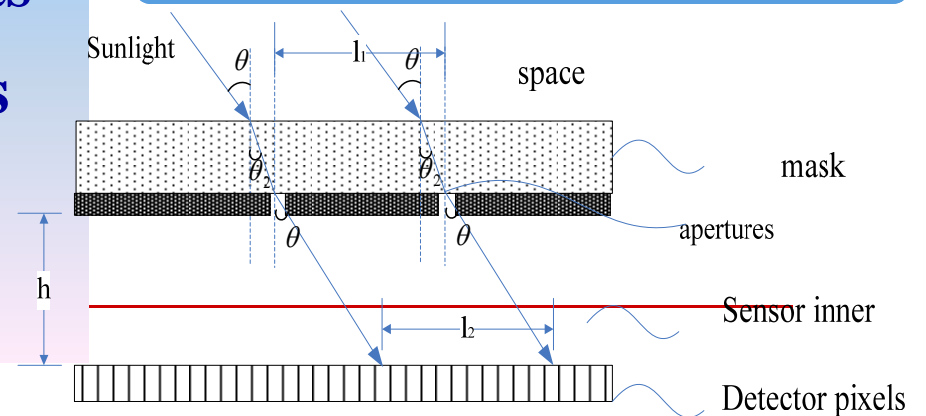
- 1: sun spots only occupy 2% of the whole sensor area;
- 2: sun spots is equal spacing distribution.
- 3: the relative position of sun spots in the sensor is invariable, which is design-related, however incident angle.

Tsinghua University

## ➤ Sun spots

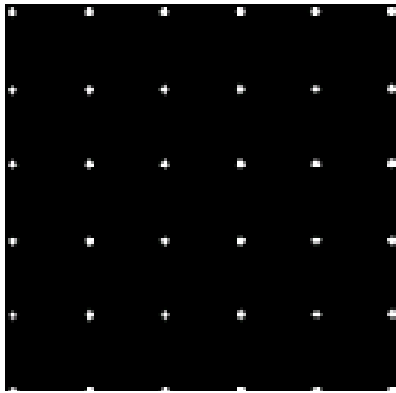


## ➤ Equal spacing distribution model



# Initial Capturing-templet construction

➤ Templet image



a) Templete image

Features Extraction



1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0
1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0
1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0
1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0
1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0
1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0

b) binary templet

➤ Features

Each sun spot signed by 1, circumambience area signend by 0, each 1 enclosed by three 0 constructs one identified area; templet total has 36 identified area.

# ⊕ Initial Capturing-algorithm processing

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**1. Image binary storing :** make Image binarization, and store the binary result pixels every other column and every other row.

**Image characteristics:** storing binary image meets the 2 bytes and 1024 bytes interval between the horizontal adjacent and vertical adjacent of sun spots.

1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0
1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0
1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0
1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0
1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0
1	0	1	0	1	0	1	0	1	0	1	0
0	0	0	0	0	0	0	0	0	0	0	0



# ⊕ Initial Capturing-algorithm processing

**2. templet:** Each sun spot enclosed by 3 dark points, for reliability, at least 31 templets need be recognized.

$A(i) > 0$  denote that existing a sun spot in the Successive 16 bytes.

$$P(i) = (A(i) \& A(i+1)) | (A(i) \& A(i+64 \times 8)) | (A(i) \& A(i+64 \times 8+1))$$

$P(i) = 0$  denote that the sun spot enclosed by dark area.

$$\begin{cases} A(l) > 0 \\ P(l) = 0 \end{cases} \quad l = \{i + k \times 2 + j \times 1024 + p \times 2\} \quad \begin{cases} k = -5, -4, \dots, 0 \\ j = 0, 1, \dots, 5; p = 0, 1, \dots, 5 \end{cases}$$

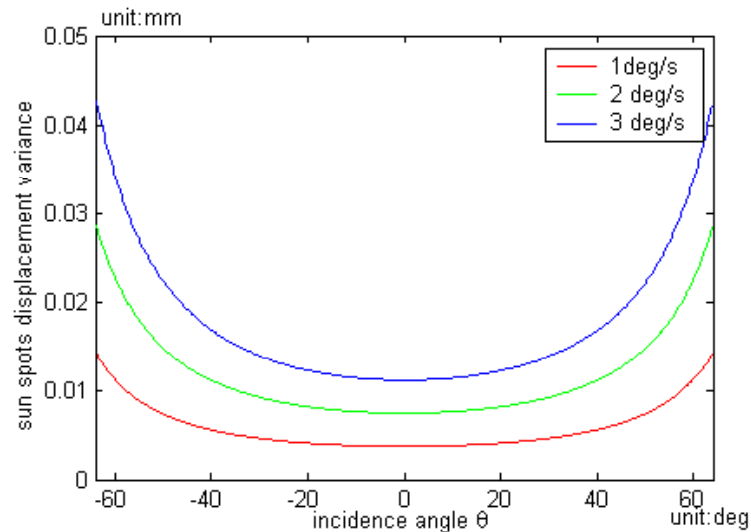
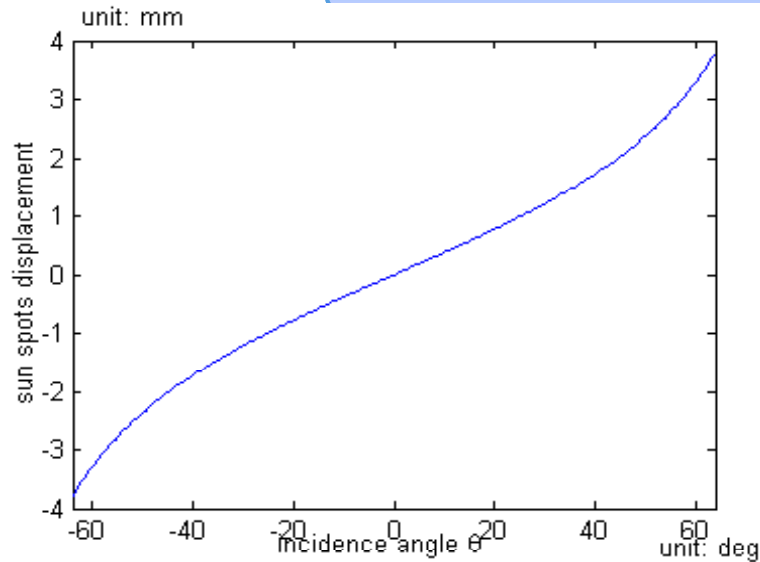
## ➤ Conclusion

To given domain of  $i, k$  if there is  $l$  meeting the requirement of  $\begin{cases} A(l) > 0 \\ P(l) = 0 \end{cases}$

, that indicates the capturing successful.

# Future Extraction(FE) Algorithm

## ➤ Sun spots position variation

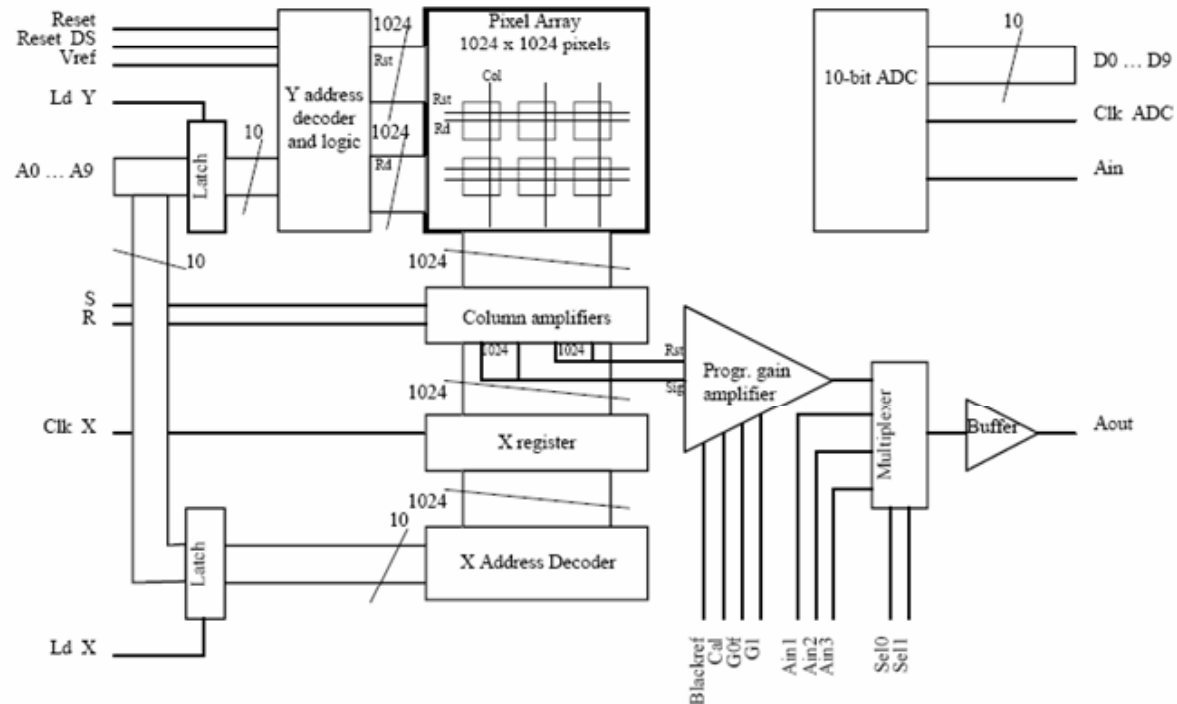


## ➤ characteristics

Sun spots positions have very small change, no more than 4 pixels, between the two successive frames, so the direct recursion is using as the future extraction algorithm.

# Future Extraction of APS

## ➤ APS readout

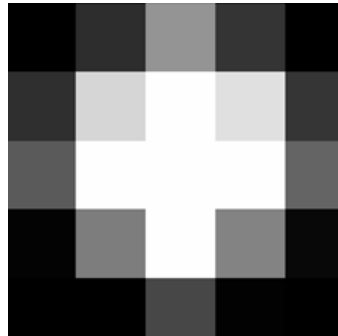


## ➤ Features

APS sensor has multi windows and random access and subsampling functions, with which the FE algorithm can directly image and readout the sun spots region.

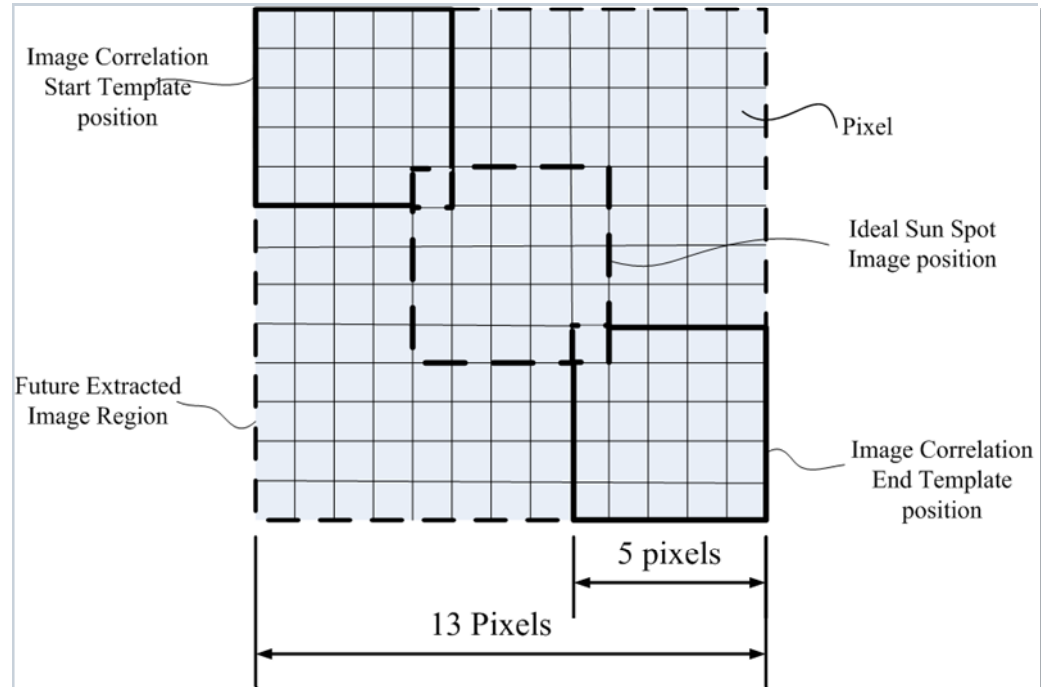
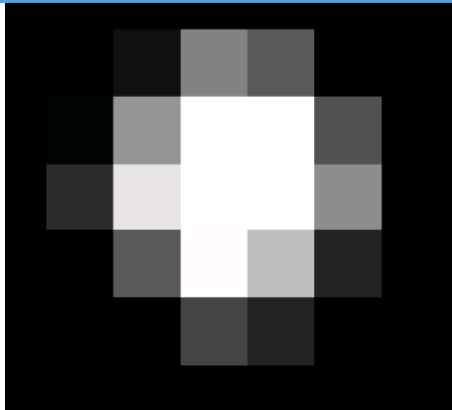
# image correlation(IC)

➤ Sun spot templet



➤ Image correlation

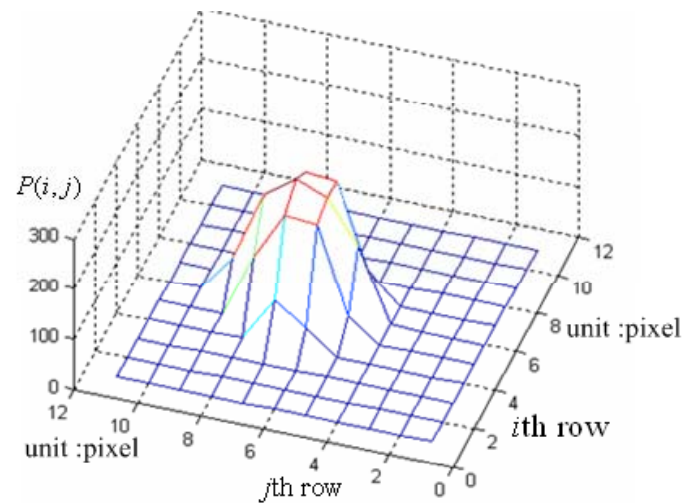
➤ FE Sun spot



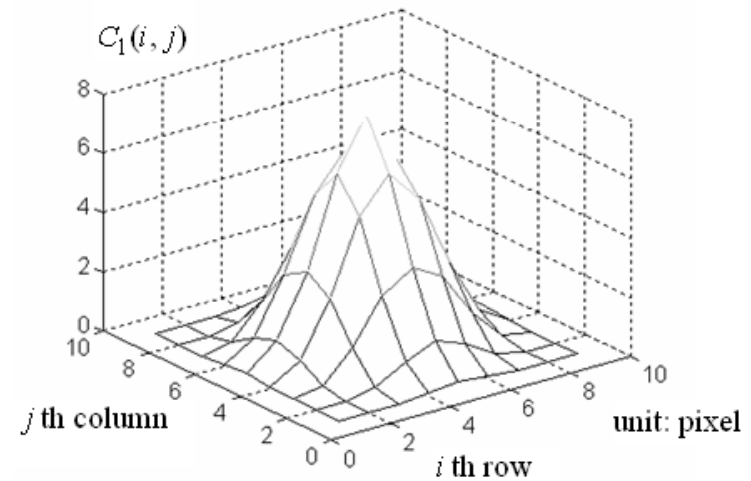
Through the image correlation of the templet and the FE sun spot, acquiring the small change of sun spot position.

# Energy distribution after correlation

➤ Tradition energy distribution



➤ energy distribution after IC



compared to the traditional energy distribution, the sun spot has higher SNR (signal to noise ratio) after IC (image correlation)。

# Centroid algorithm after IC

➤ Traditional Centroid

$$x_c = \frac{\sum_{i=x_{peak}-k}^{x_{peak}+k} \sum_{j=y_{peak}-k}^{y_{peak}+k} P(i, j) \times i}{\sum_{i=x_{peak}-k}^{x_{peak}+k} \sum_{j=y_{peak}-k}^{y_{peak}+k} P(i, j)}$$
$$y_c = \frac{\sum_{i=x_{peak}-k}^{x_{peak}+k} \sum_{j=y_{peak}-k}^{y_{peak}+k} P(i, j) \times j}{\sum_{i=x_{peak}-k}^{x_{peak}+k} \sum_{j=y_{peak}-k}^{y_{peak}+k} P(i, j)}$$

➤ Centroid after IC

$$\tilde{x}_1 = \frac{\sum_{m=0}^8 \sum_{n=0}^8 C_1(m, n) \times (m - 4)}{\sum_{m=0}^8 \sum_{n=0}^8 C_1(m, n)}$$
$$\tilde{y}_1 = \frac{\sum_{m=0}^8 \sum_{n=0}^8 C_1(m, n) \times (n - 4)}{\sum_{m=0}^8 \sum_{n=0}^8 C_1(m, n)}$$

**After IC, increasing SNR, degreasing regional pixels impact, the sun spot location accuracy will be greatly increased.**

# Multi sun spots weighted algorithm

## ➤ Weighted coefficient

$$x_c = \sum_i^{36} a_i x_{ci}$$

$$y_c = \sum_i^{36} b_i y_{ci}$$

$$a_i = b_i = \frac{\sum_{m=0}^8 \sum_{n=0}^8 C_i(m, n)}{\sum_{i=1}^{36} \sum_{m=0}^8 \sum_{n=0}^8 C_i(m, n)}$$

## ➤ Centroid after Weighted

$$x_c = \Delta x + \frac{\sum_{i=1}^{36} \sum_{m=0}^8 \sum_{n=0}^8 C_i(m, n)(m - 4)}{\sum_{i=1}^{36} \sum_{m=0}^8 \sum_{n=0}^8 C_i(m, n)}$$

$$y_c = \Delta y + \frac{\sum_{i=1}^{36} \sum_{m=0}^8 \sum_{n=0}^8 C_i(m, n)(n - 4)}{\sum_{i=1}^{36} \sum_{m=0}^8 \sum_{n=0}^8 C_i(m, n)}$$

## ➤ Features

**The energy-weighted method using 36-point not only improves the centroid accuracy better than 0.02 pixels, but also eliminates the impact to system when some the sun spots lost or disturbed.**

# Solving $\tan \theta$ through polynomial fitting

## ➤ principle

$$l_n = \sqrt{(x_{cn} - x_0)^2 + (y_{cn} - y_0)^2}$$

$$\tan \theta_n = a_5 l_n^5 + a_4 l_n^4 + a_3 l_n^3 + a_2 l_n^2 + a_1 l_n^1 + a_0$$

## ➤ method

To any  $\theta$  at given axis, measuring the corresponding sun spots position ,  $l$  acquire the polynomial coefficients.

## ➤ equipments

Sun simulator, two-axis gimbal, Theodolite etc.





# Contents

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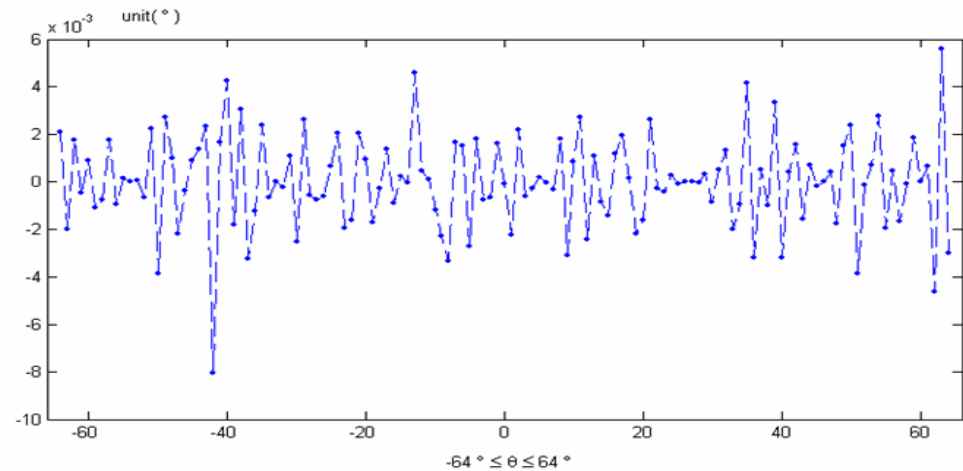
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# ⊕ The accuracy performance

Sun sensor in test set-up



Measurement error statistics



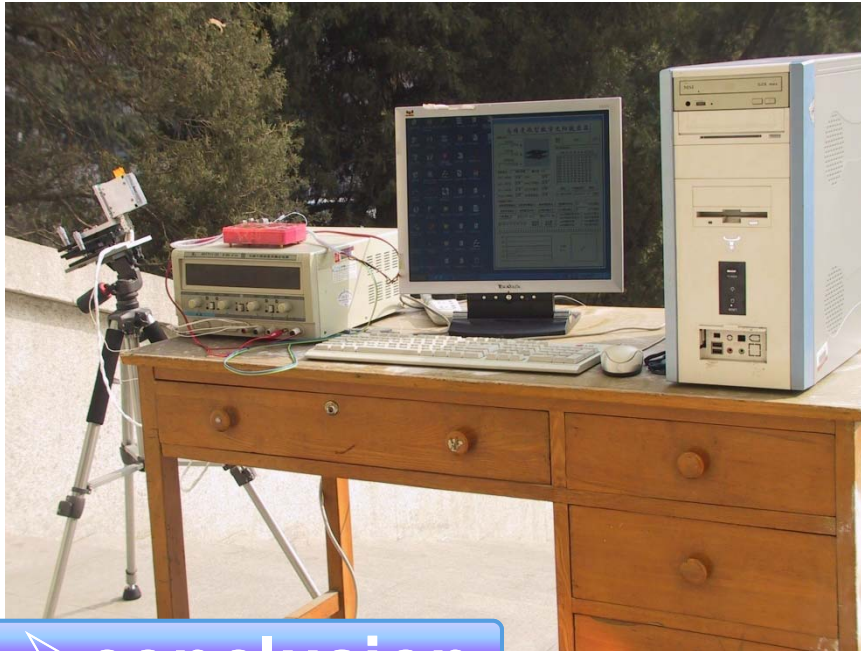
Sensor angle of incidence(° )

➤ conclusion

The residual random error is **0.008** degree of arc.

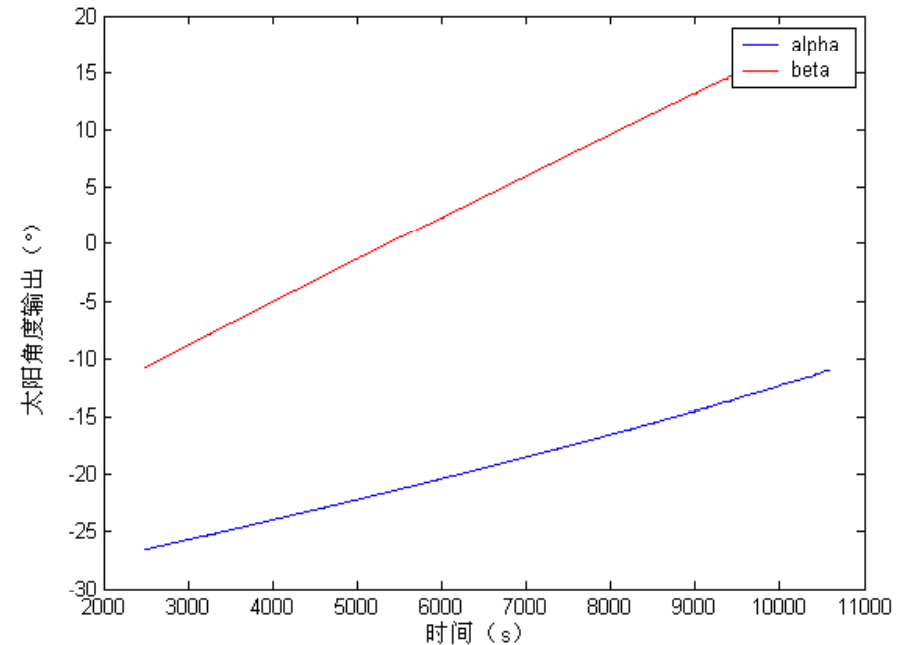
# ☉ Real sun experiment

Real sun experiment set-up



➤ conclusion

Sensor output result(Earth's rotation )



**Sun sensor output data is homologous to (ascending or descending) per second , and no jitter case, which determine that the system resolution of better than 0.004**

# ⊕ Immunity performance(1)

1: some apertures missing

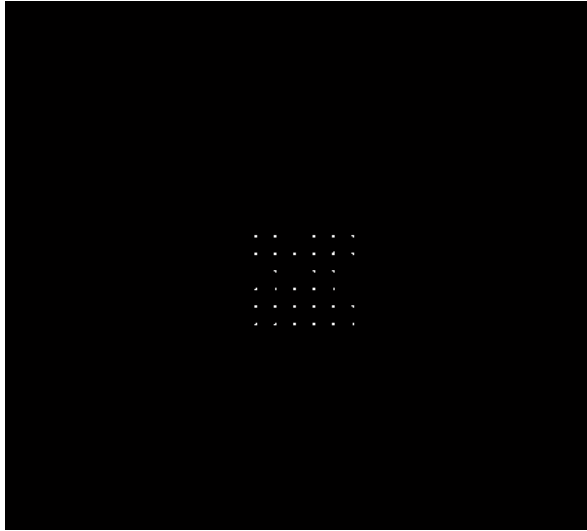
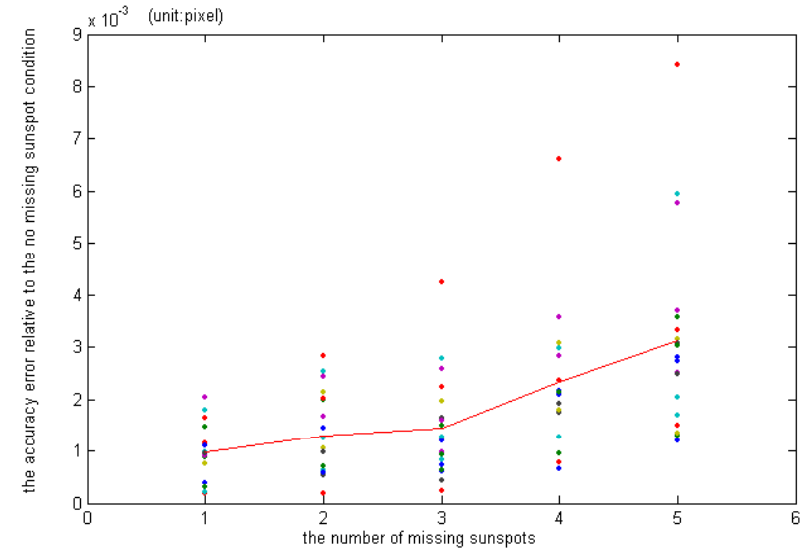


image with five sun spots missing

➤ conclusion

**with five sun spots missing , the sun spots position error is no more than 0.01 pixels which is little effect on system accuracy.**

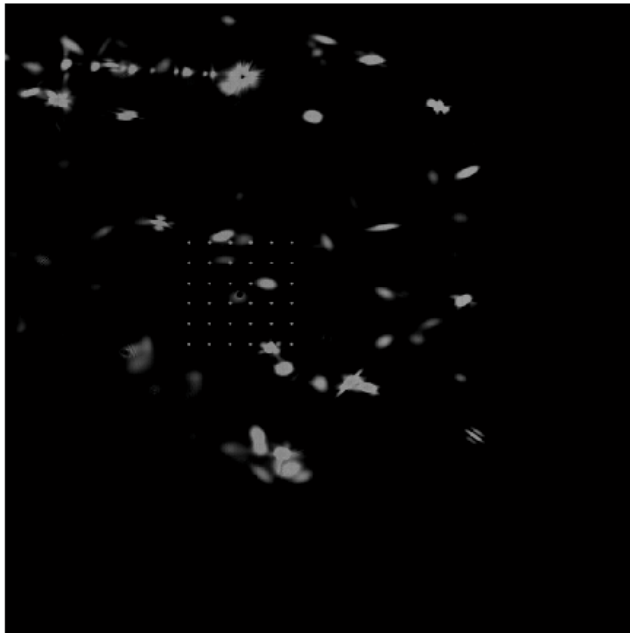


accuracy analysis with one–five sun spots missing at different sun angle of incidence.

# ⊕ Immunity performance(2)

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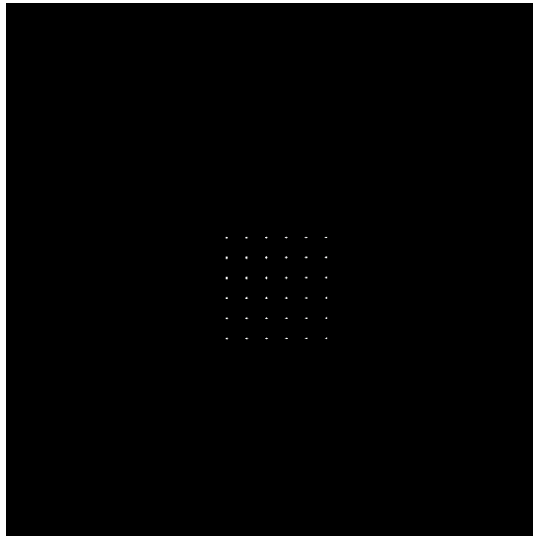
## 2: image deteriorated



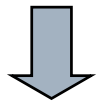
### ➤ conclusion

**Under above case, the sun sensor still works normally. It shows that the sun sensor has very strong immunity to image noise and harsh space environment.**

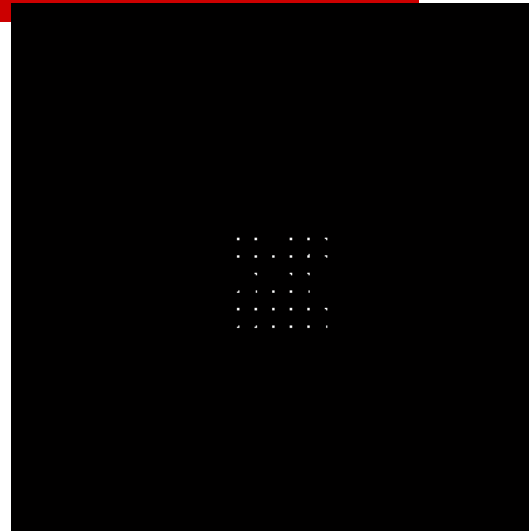
# ⊕ System performance summary



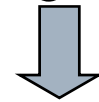
Normal case, 36 sun spots



① Normal case, system accuracy prior to 0.02 pixel



Some sun spots missing case



② Some sun spots missing case, keep same accuracy with ①

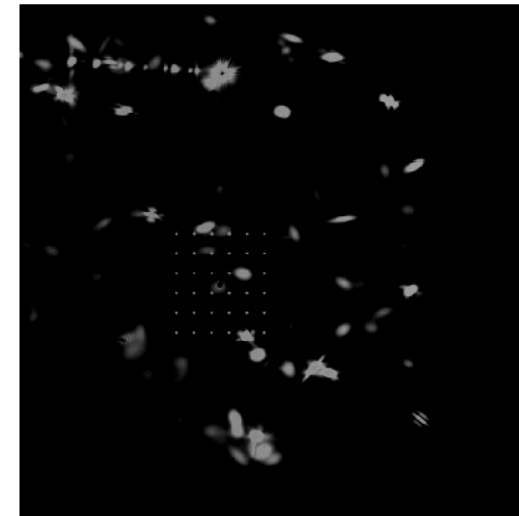
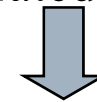


Image severely deteriorated case



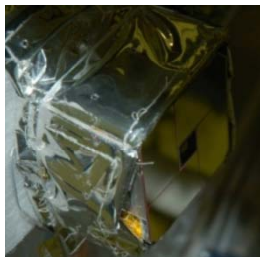
③ Image deteriorated case, keep system normally works.

# Testing in Space

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**SY-3 satellite**

**MEMS sun sensor**

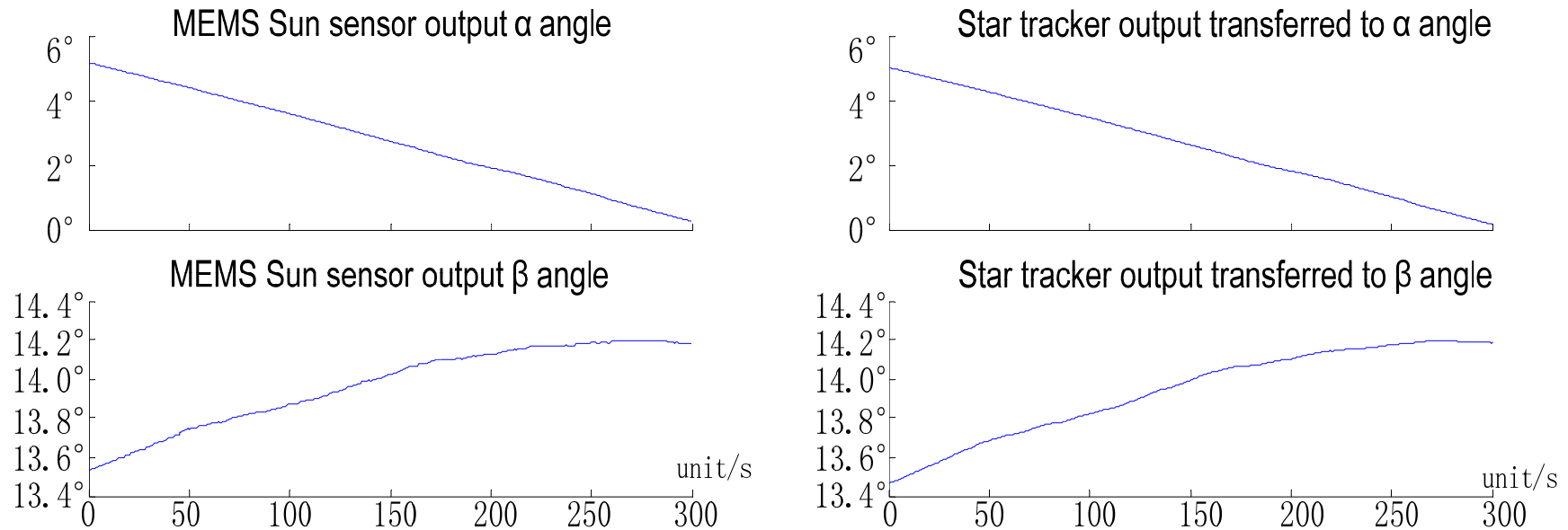


**The SY-3 satellite with our MEMS sun sensor was launch in JiuQuan launch center of China, Nov.5,2008**

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# Testing in Space

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7:00~7:05, Nov. 12 ,2008

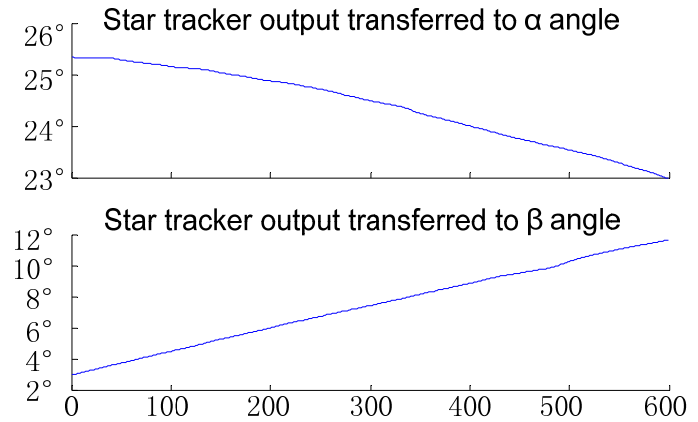
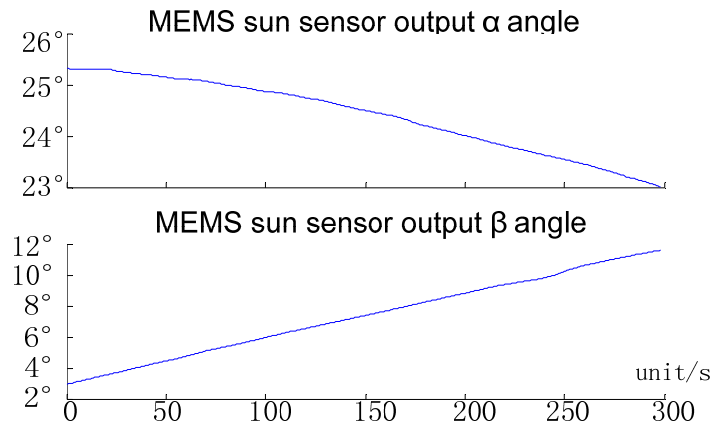
**The comparing of MEMS sun sensor output result and the star tracker output result, that has been transfer to the sun sensor coordinate.**

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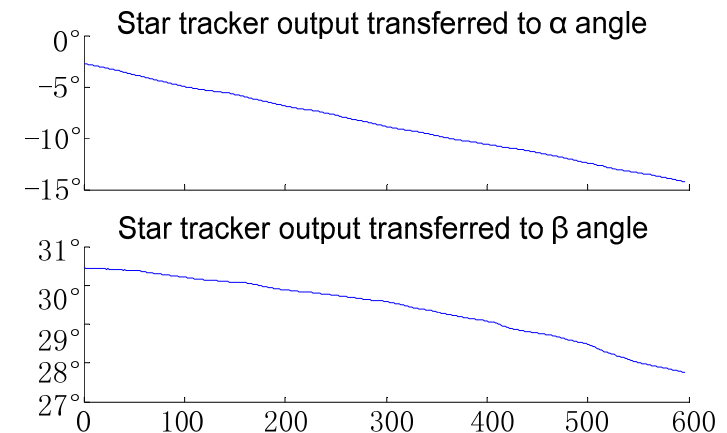
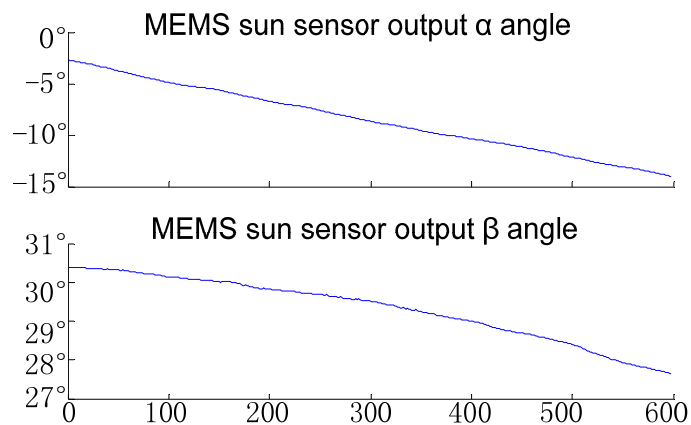


# Testing in Space

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8:36~8:41, Aug. 7, 2009

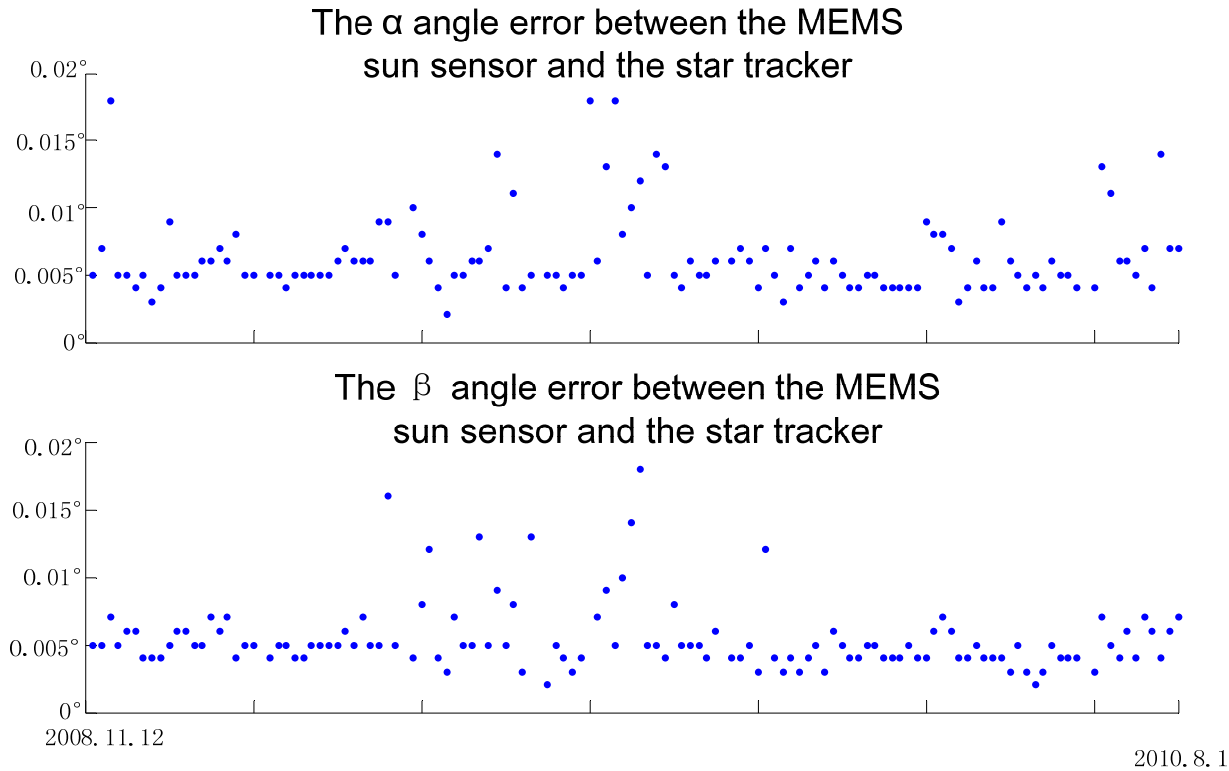


18:15~18:20, July. 8, 2010

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# Accuracy Analysis

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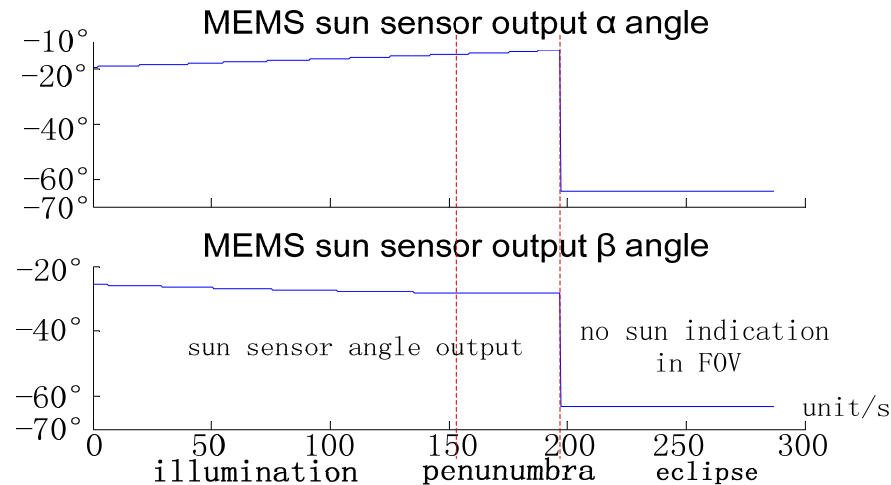


**On orbit test, the accuracy of MEMS sun sensor is prior to  $0.02^\circ$  .**

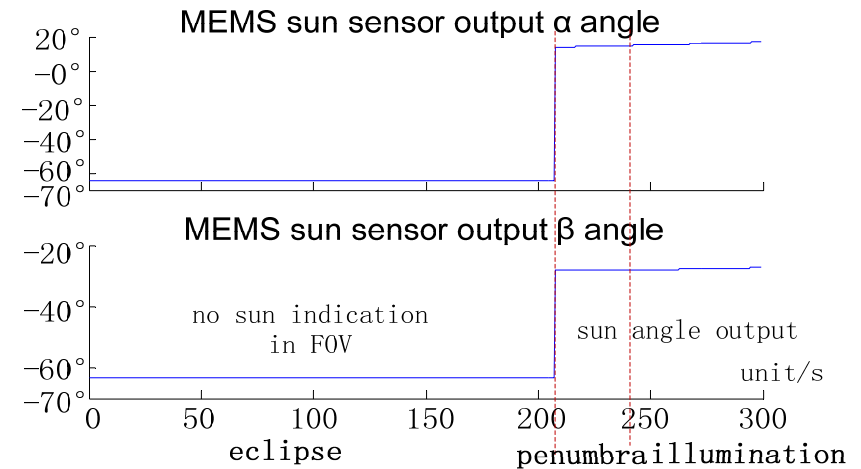
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# Penumbra and eclipse test

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July 9 ,2009, 6:35~6:40, the satellite start the process from the sun illumination via penumbra to eclipse.



July 12 ,2009, 7:04~7:09, the satellite start the process from the eclipse via penumbra to illumination.

**The experiment show the MEMS sun sensor can normally work on the process of the satellite into or out of the eclipse. ◦**

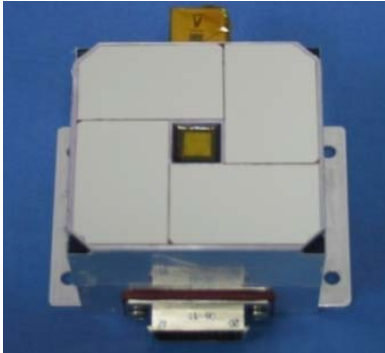
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# Summary

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Tsinghua university-THU  
(MEMS Sun Sensor)

parameter	THU MEMS Sun Sensor
accuracy	Prior to 0.02°
Field of View	128° × 128°
Resolution	Prior to 0.004°
mass	350g
Update data rate	10Hz

# Summary

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- **the MEMS based 36 apertures mark as the optical frond-end.**
- **Future extraction and image correlation-FEIC as the high accurate algorithm.**
- **Strong immune to some sun spots missing, image deteriorated and harsh space environment.**

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Thank  
you!  
Tsinghua  
University

**Reference:**

**Fei Xing , Zheng You, GaoFei Zhang, Jian Sun. Anovel active pixels sensor (APS) based sun sensor based on a feature extraction and image correlation (FEIC) technique. *Measurement Science & Technology* Vol: 19 Issue: 12 Pages: 125203 (9pp.) (SCI, UT ISI: 000260759000007)。**

Contact: Xing Fei: [xingfei@tsinghua.edu.cn](mailto:xingfei@tsinghua.edu.cn);

Department of Precision Instruments and Mechanology, BLD 9003, RM2008,  
Tsinghua University, BeiJing, China.

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**Tsinghua University**